

*The Evolution of the Mass-Size relation to $z = 3.5$
in the GOODS-North Field*



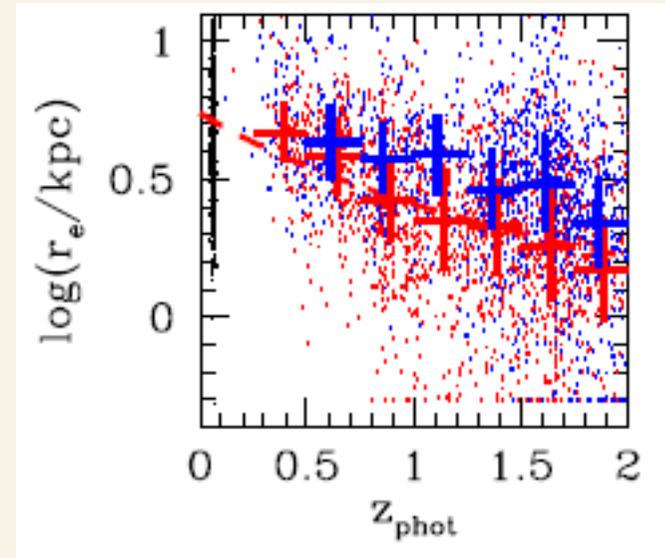
Moein Mosleh

THE EVOLUTION OF THE MASS-SIZE RELATION TO $Z=3.5$ FOR UV-BRIGHT GALAXIES AND SUB-MM GALAXIES IN THE GOODS-NORTH FIELD

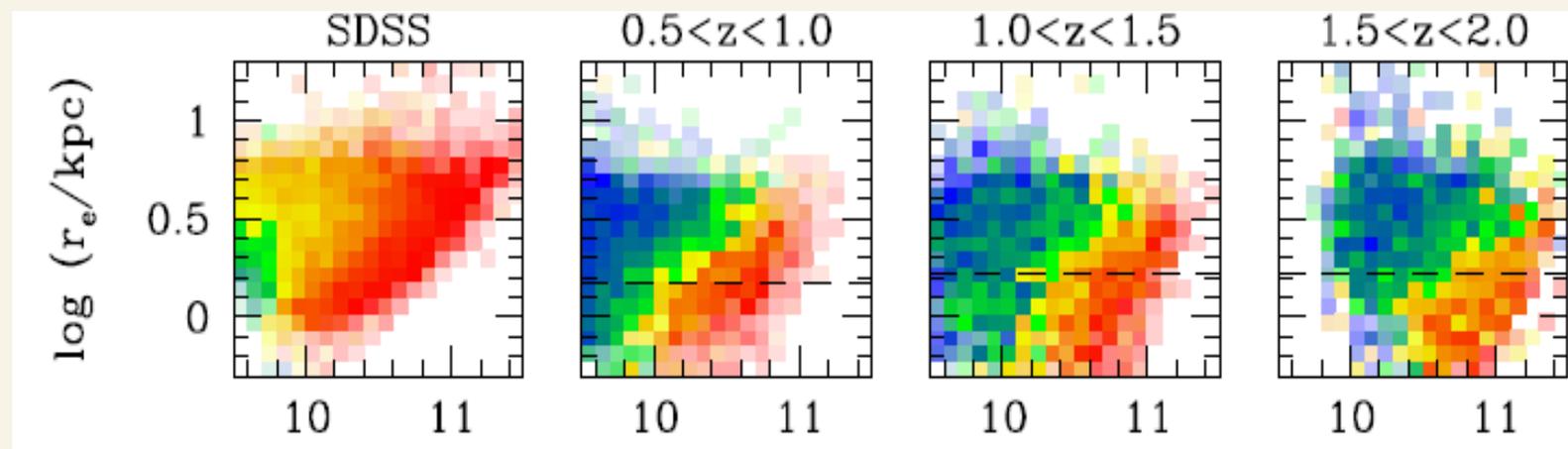
MOEIN MOSLEH¹, RIK J. WILLIAMS², MARIJN FRANX¹, MARISKA KRIEK³

Introduction

- ❖ *Galaxies at high redshift have significantly smaller sizes than the local galaxies of similar masses*
- ❖ *Quiescent Galaxies are significantly smaller than star-forming galaxies of similar masses*
- ❖ *These studies are almost bases on Photometric redshift especially at high z*
- ❖ *Size measurements of Sub-mm or LBG galaxies at high z are also rare*



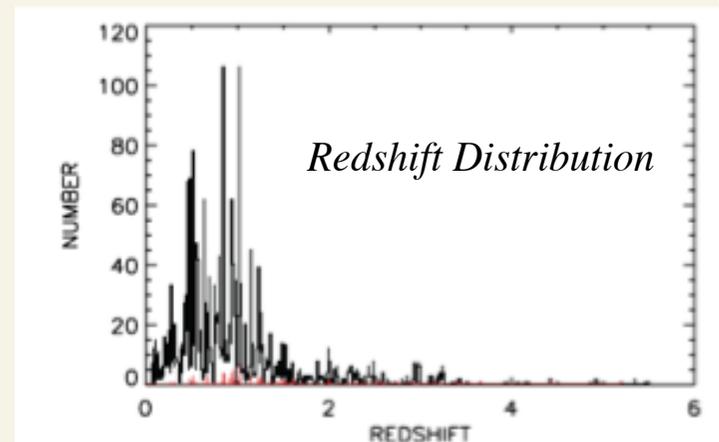
Williams et al. 2010



DESCRIPTION OF DATA

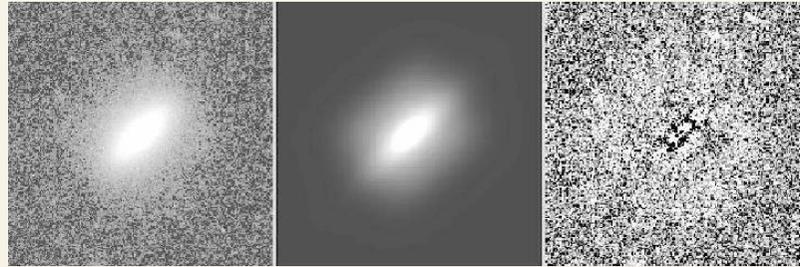
Using most complete spectroscopic catalog of galaxies in the GOODS-North Field. (Barger et al. 2008)

- ★ *Using U band (Capak et al. 2004)*
- ★ *ACS Magnitudes (F435w (B), F606W (V), F775W (I), F850LP (Z))*
- ★ *Ks-band magnitudes from WIRCam (CFHT, 3.6 m) (6909)*
- ★ *NUV – FUV from GALEX (1016)*



SIZES

- ✓ *Using GALFIT (Peng et al. 2002) to fit Sersic models*
- ✓ *Ks-band image from Subaru 8M Telescope*
- ✓ *FWHM $\sim 0.5''$*
- ✓ *Pixel scale = 0.12 arcsec/pixel*

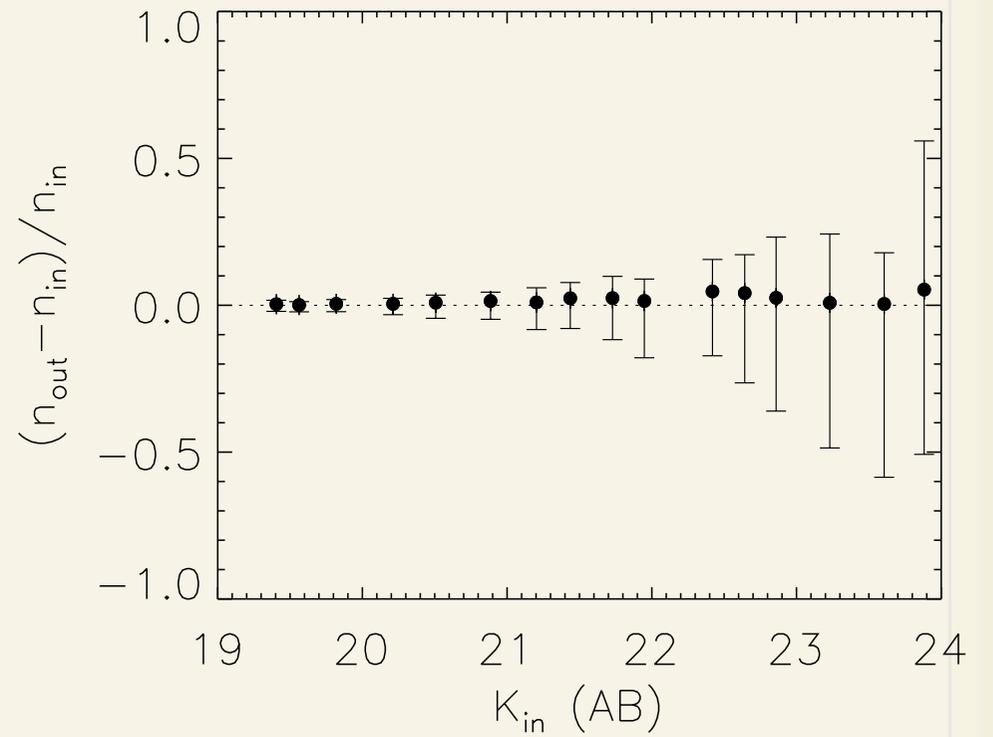
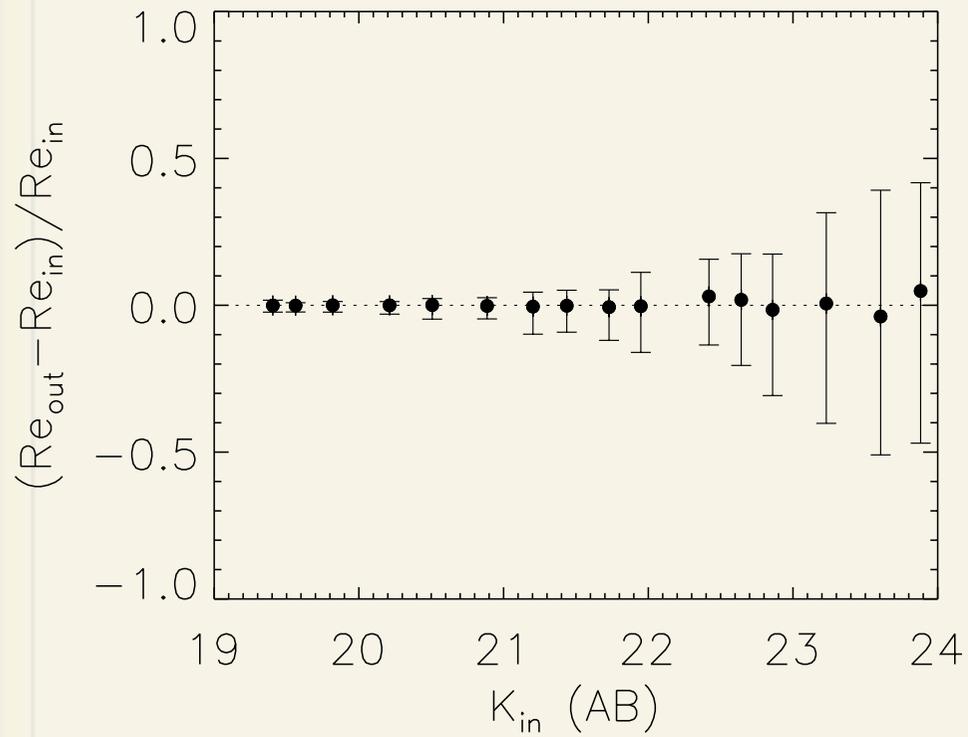


STELLAR MASS ESTIMATES

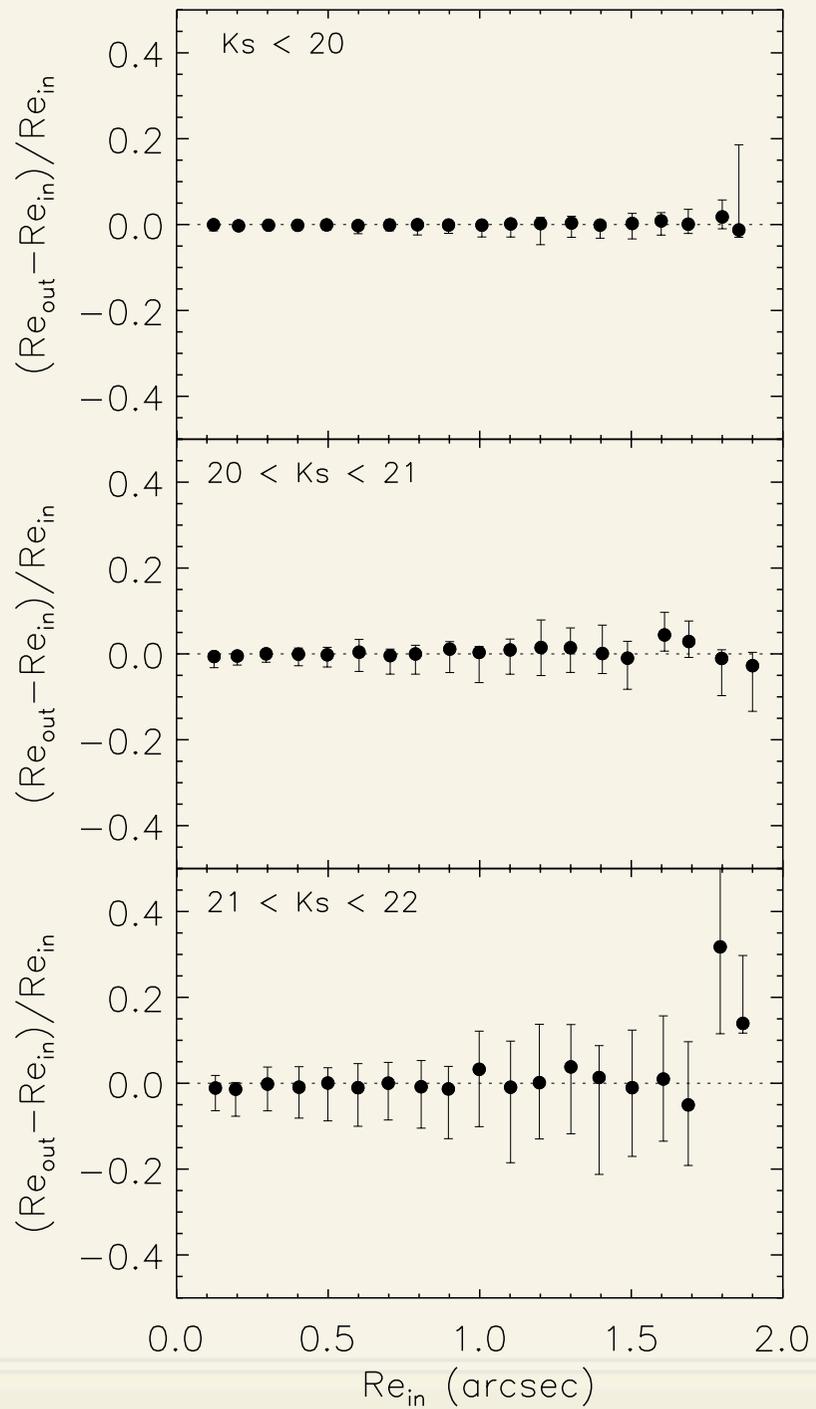
- ✓ *Using FAST (Fitting and Assessment of Synthetic Templates code) (Kriek et al. 2009)*
- ✓ *Using Salpeter IMF & Solar Metallicity*
- ✓ *Masses are corrected by a factor of -0.2 dex (for consistency with a Kroupa (2001) IMF)*
- ✓ *Bands: U, B, V, I, Z, Ks*

Simulations

Running simulation for testing the accuracy of parameter measurements



Simulations



SUB-SAMPLES

◆ *UV- Bright Galaxies :-*

- *Lyman Break Galaxies (LBGs) ($2.7 < z < 3.4$)*

Selected from Reddy et al. 2006

- *BM/BX ($1.4 < z < 2.7$)*

Selected from Reddy et al. 2006 Catalog

- *GALEX/LBGs ($0.6 < z < 1.4$)*

Selected by using (FUV-NUV) versus (NUV-F435W) color- color diagram

◆ *Sub-mm Galaxies (SMG)*

Selected from samples of Chapman et al. 2005 & Pope et al. 2006

◆ *Star-forming BzK (sBzK) ($1.4 < z < 3.2$)*

TABLE 1
SUB-SAMPLES

Sample	Redshift	$\log(M_*/M_\odot)$	No. of Sources
GALEX/LBG	$0.6 < z < 1.4$	8.8 - 11.0	105
BM/BX	$1.4 < z < 2.7$	9.8 - 10.8	41
LBG	$2.7 < z < 3.5$	10.2 - 10.8	5
SMG	$0.5 < z < 3.$	10.0 - 11.7	14
sBzK	$1.4 < z < 3.2$	9.8 - 11.7	70

Size Evolution

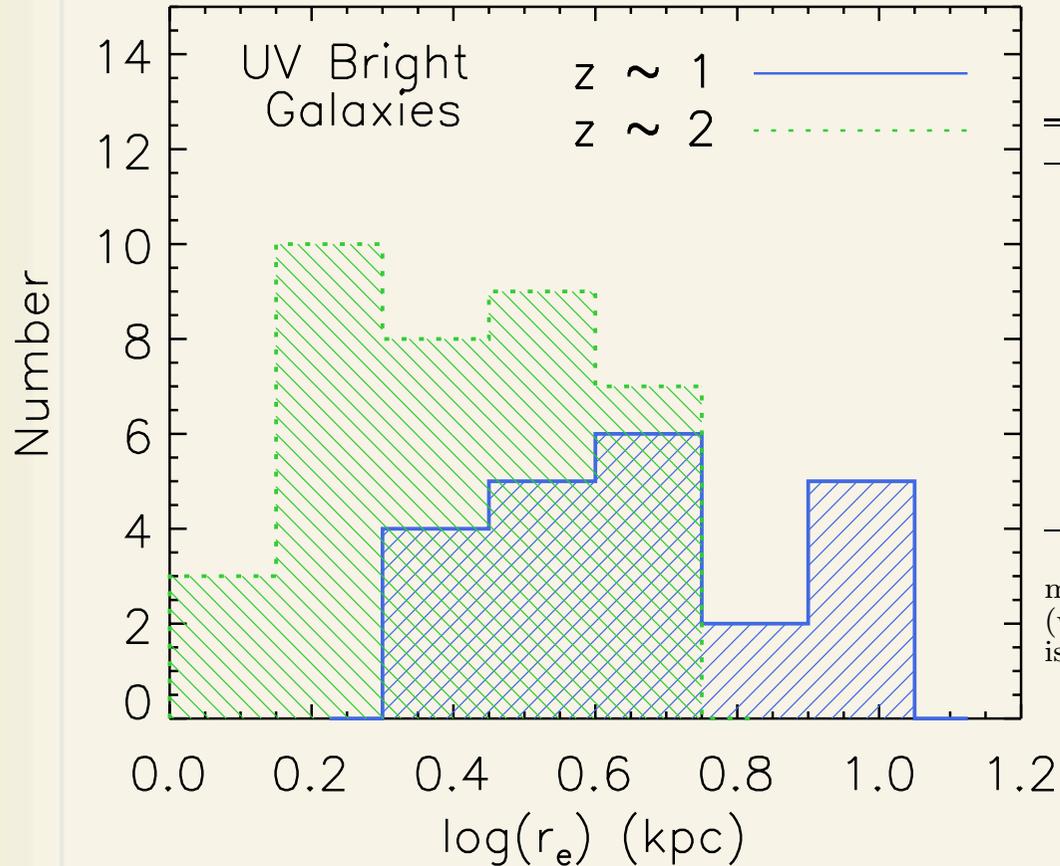


TABLE 3
MEDIAN SIZES OF DIFFERENT SAMPLES

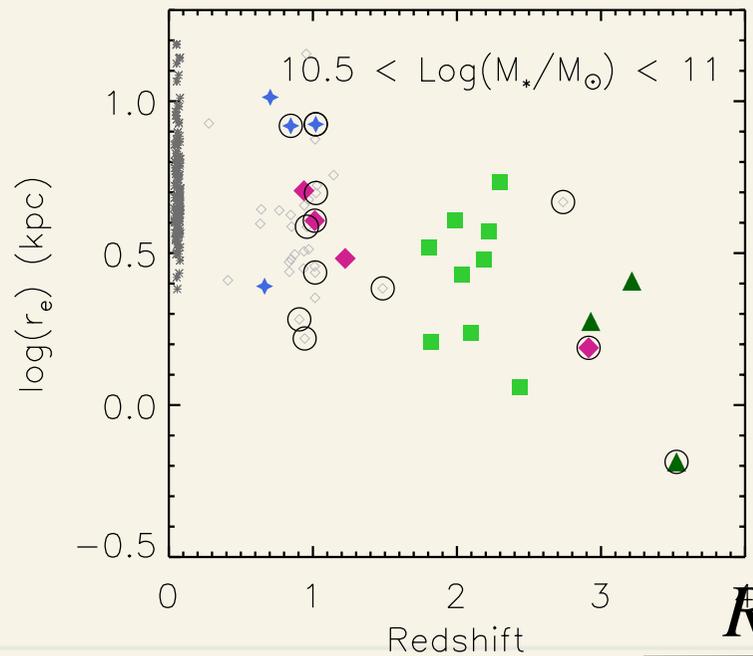
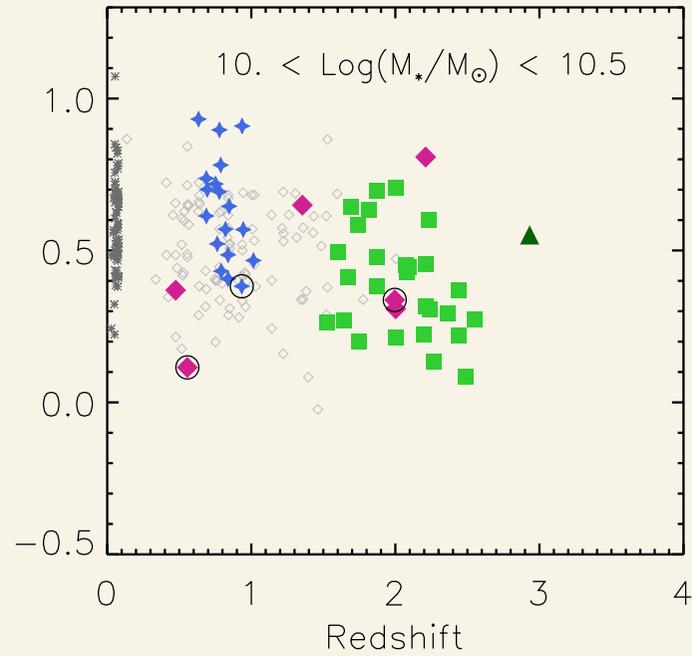
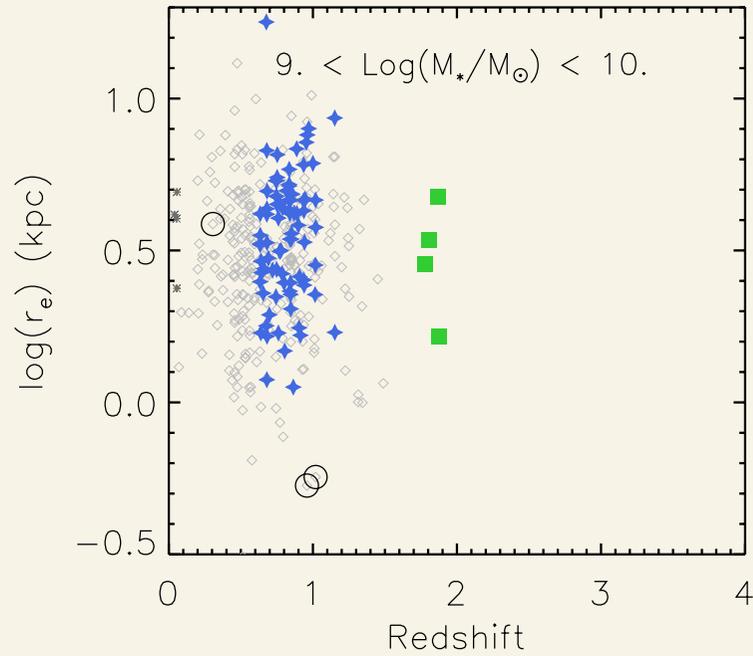
Sample	Redshift	Size	$median(\log(M_*/M_\odot))$
GALEX/LBG	$0.6 < z < 1.4$	4.42 ± 0.52	10.2 ± 0.3
BM/BX	$1.4 < z < 2.7$	2.68 ± 0.19	10.4 ± 0.2
LBG	$2.7 < z < 3.5$	2.22 ± 0.61	10.6 ± 0.2
CDFS	$0.5 < z < 1.5$	2.33 ± 0.1	10.4 ± 0.3
CDFS	$1.5 < z < 2.5$	2.40 ± 0.13	10.4 ± 0.3
CDFS	$2.5 < z < 3.5$	1.73 ± 0.18	10.4 ± 0.3
Quiescent(CDFS)	$0.5 < z < 1.5$	1.32 ± 0.07	10.5 ± 0.3
Quiescent(CDFS)	$1.5 < z < 2$	1.12 ± 0.32	10.7 ± 0.2
Quiescent(VD08)	$2 < z < 2.5$	0.9 ± 0.21	11.23
SMG	$0.5 < z < 3.$	2.90 ± 0.45	10.8 ± 0.5

Notes. The median sizes of samples are for galaxies with stellar masses between $10^{10} - 10^{11} M_\odot$ except for the quiescent(VD08) (van Dokkum et al. 2008b) and SMG galaxies. The CDFS sample is all galaxies at the same stellar mass range in CDF-S field.

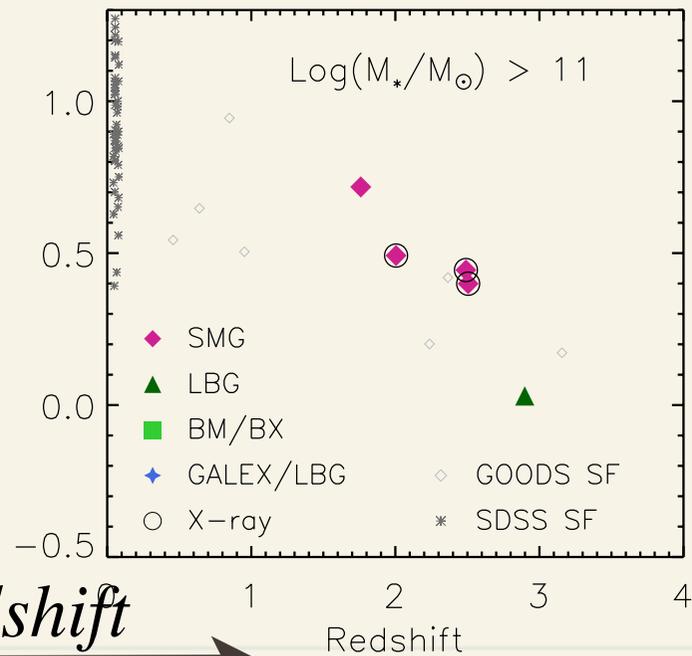
Sizes of UV-Bright galaxies evolve by a median factor of 0.60 ± 0.08 between $z \sim 2$ to $z \sim 1$

Size Evolution

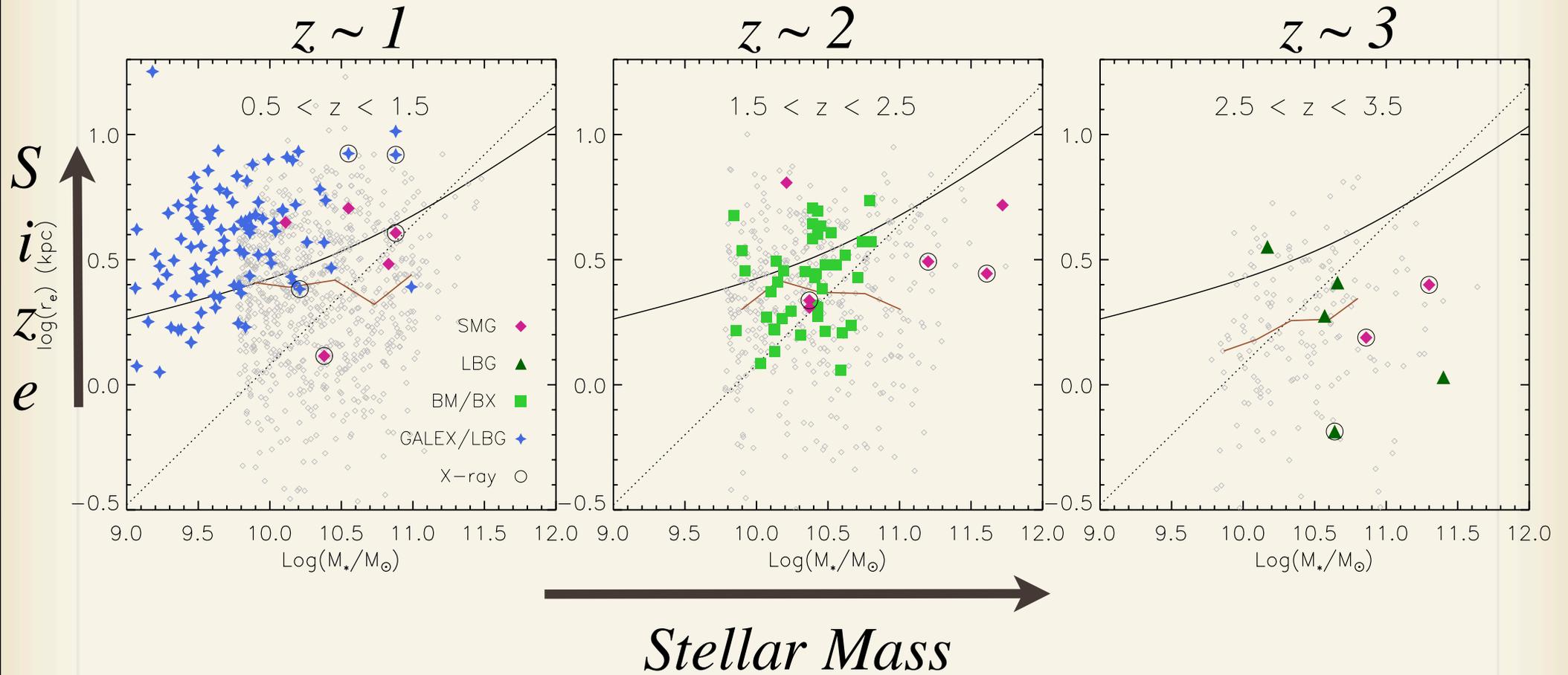
S
i
z
e



Redshift



Stellar Mass-Size Relation



Stellar Mass-Size Relation

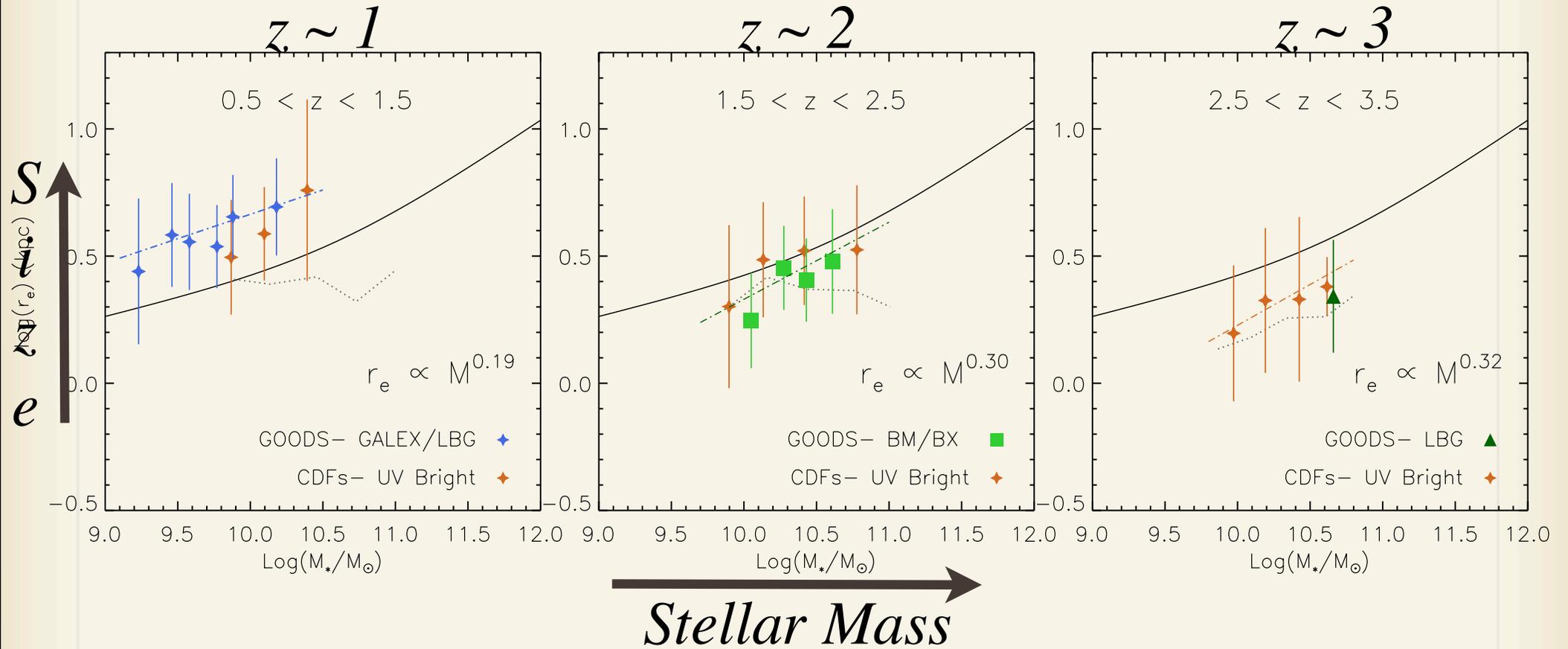
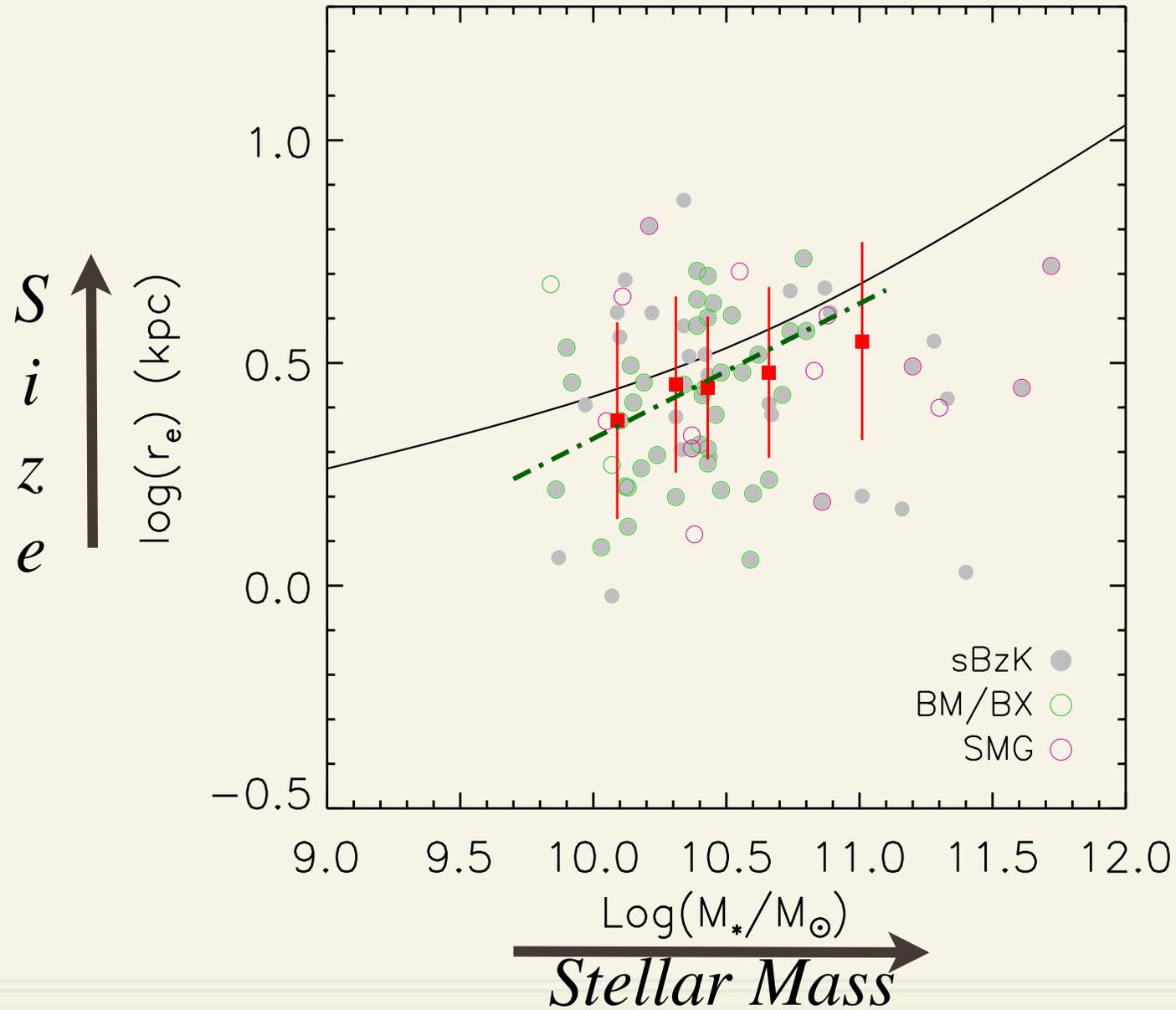


TABLE 2
BEST FITS POWER LAW PARAMETER FOR THE STELLAR MASS - SIZE
RELATION

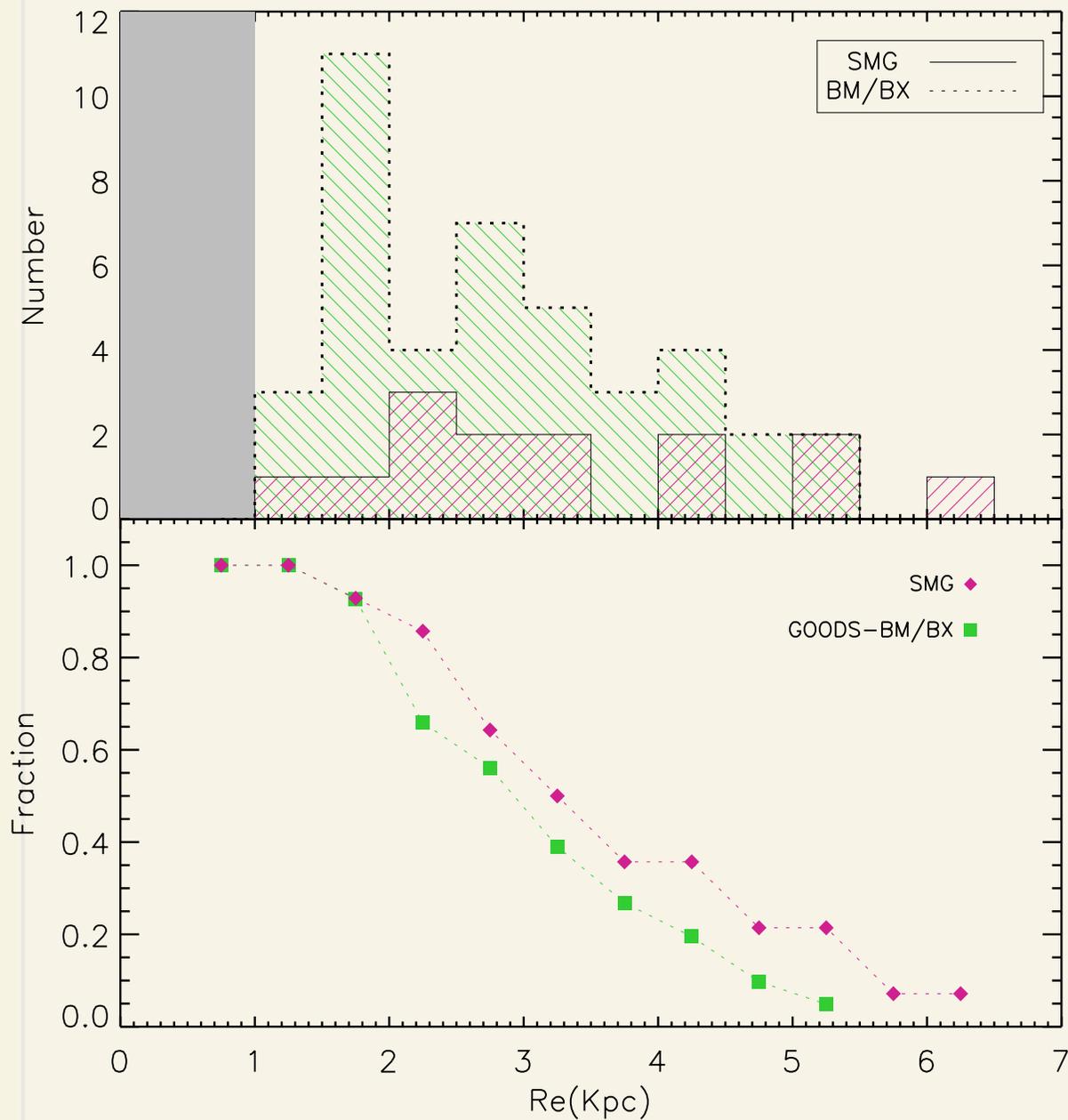
Sample	Redshift	α
GALEX/LBG	$0.6 < z < 1.4$	0.19 ± 0.05
BM/BX	$1.4 < z < 2.7$	0.30 ± 0.06
CDFs-UV Bright	$2.5 < z < 3.5$	0.32 ± 0.06
<i>sBzK</i>	$1.4 < z < 3.2$	0.31 ± 0.09

Note. Power law parameter α is defined as $r_e \propto M^\alpha$.

Stellar mass size relation for star forming *sBzK* galaxies



Sizes of Submm Galaxies



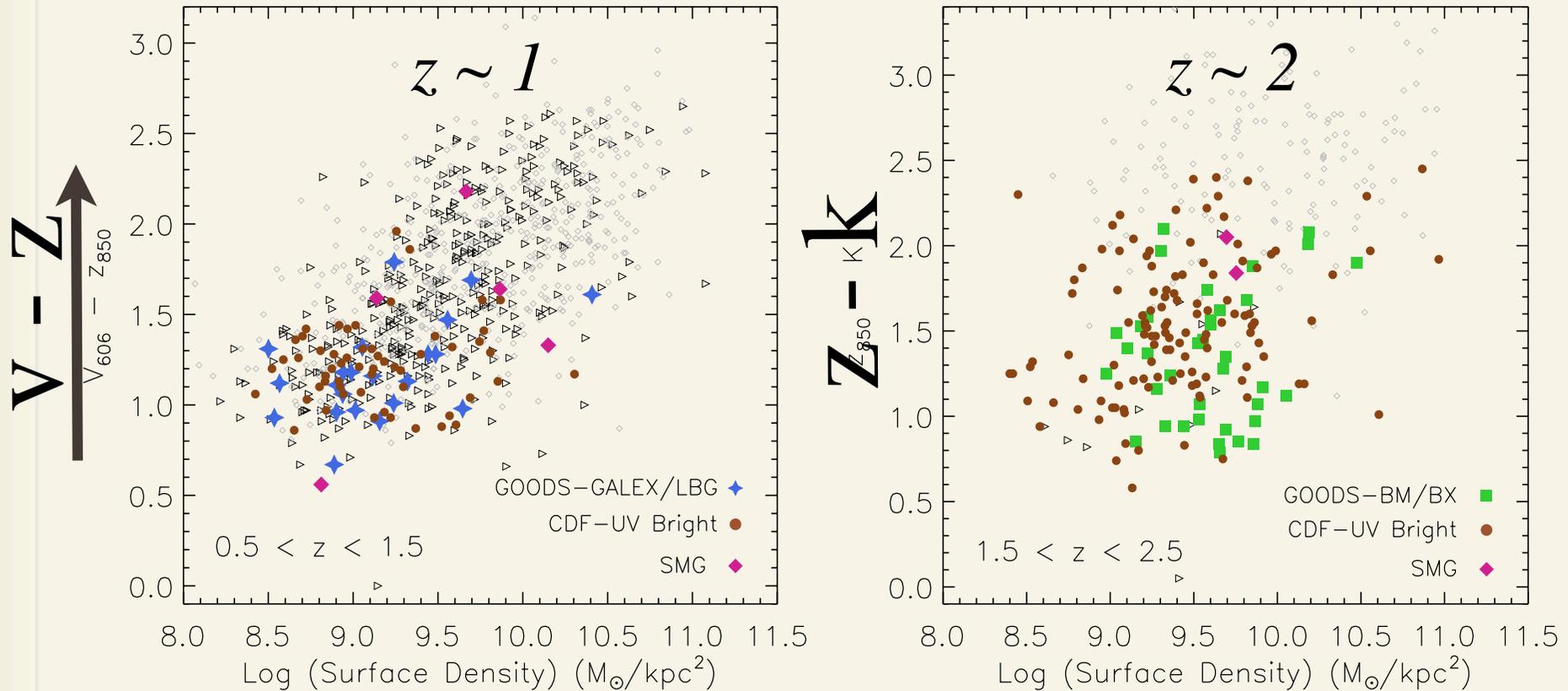
$$10 < \log(M_*/M_\odot) < 11$$

$$\overline{\text{SMGs}} = 2.65 \pm 0.56$$

$$\overline{\text{BM/BX}} = 2.68 \pm 0.19$$

Color - Surface Density

$$10 < \log(M_*/M_\odot) < 11$$

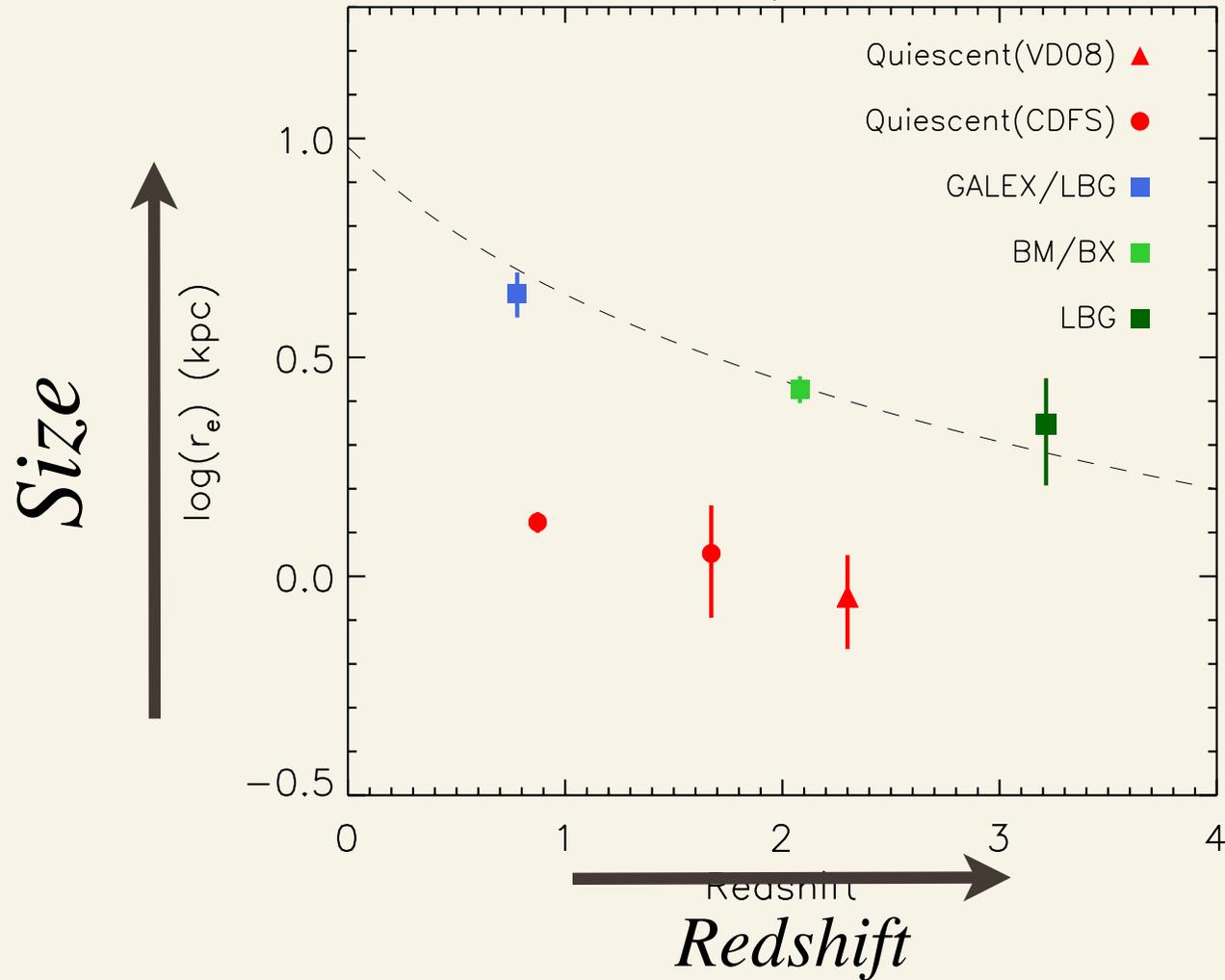


Stellar Mass Surface Density

Conclusion



$10 < \log(M_*/M_\odot) < 11$



UV-bright galaxies evolve as $r_e \propto (1 + z)^{-1.11 \pm 0.13}$

Summary

- ~ *Quantifying the size mass evolution of galaxies without potential uncertainties of photometric redshift.*
- ~ *UV-Bright galaxies evolve strongly with redshift.*
- ~ *UV-Bright galaxies are significantly larger than quiescent galaxies. At the same mass, the median difference is 0.45 ± 0.09 .*
- ~ *The LBG, BM/BX and GALEX/LBG samples show smooth evolution with redshift.*
- ~ *The SMGs have half-light radii similar to UV-Bright galaxies of the same mass.*

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