

# DYNAMIC WITH EUCLID GLOBULAR CLUSTERS

*What will Euclid measure concerning the GCs?* Lançon et al. 2021



Stellar mass based on luminosity



Projected position in the sky



Projected distance to the host galaxy



Projected size (it will be difficult to estimate)



Photometric redshift



Number per galaxy



Environment: isolated GCs? Properties of the host galaxy?



Age/Metallicity (it will be difficult to estimate)

# DYNAMIC WITH EUCLID GLOBULAR CLUSTERS

**EUCLID STRENGTH:  
STATISTICS!!**

*New constraints on? Improve our understanding of?*



Nature of dark matter



Nature of globular cluster



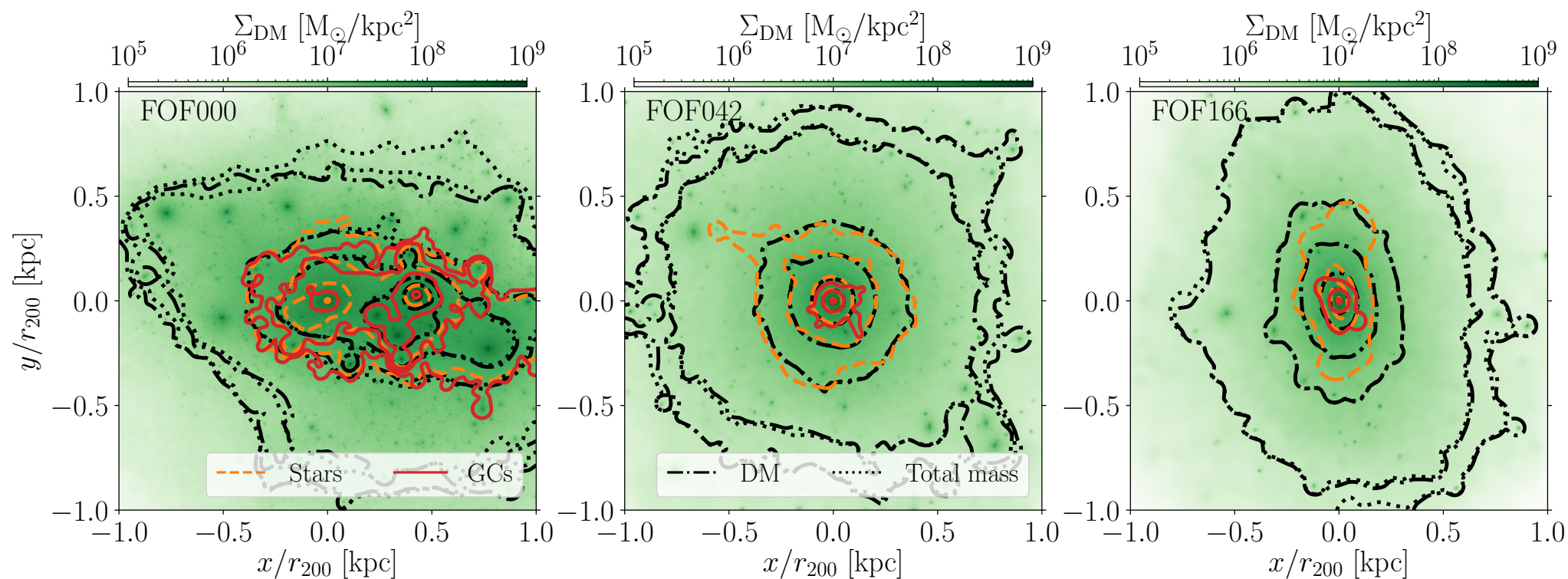
Galaxy formation model

# DYNAMIC WITH EUCLID GLOBULAR CLUSTERS

*Constraining the nature of dark matter*

 **Shape of host dark matter halos**  
from  
*the stellar surface brightness maps*  
*the number density maps of GCs*


*Reina-Campos et al. 2022*

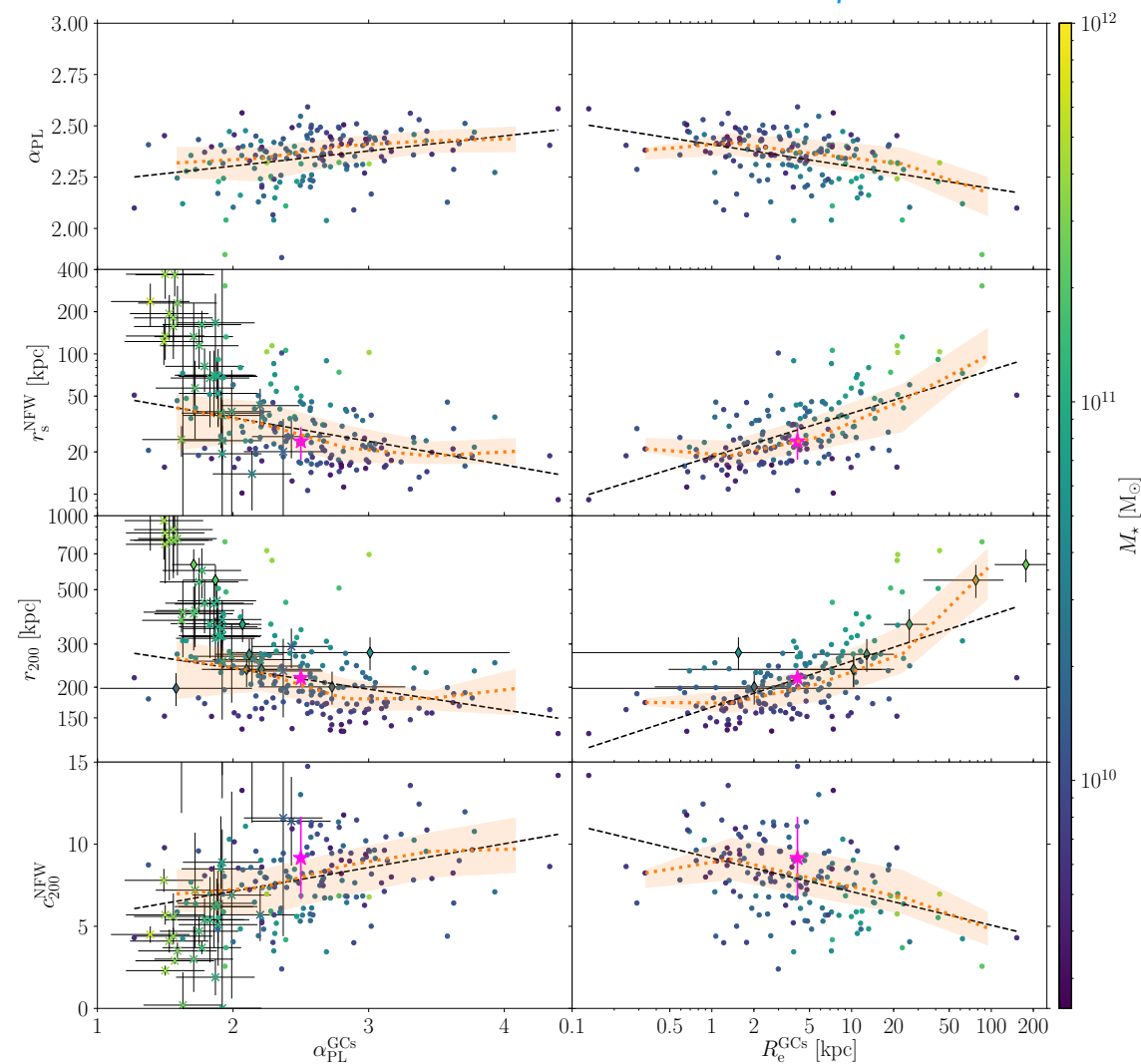


# DYNAMIC WITH EUCLID GLOBULAR CLUSTERS

Reina-Campos et al. 2021

Constraining the nature of dark matter

 Structural parameters of their host dark matter halos (slope, scale radius, virial radius, concentration) from *the projected radial profiles of GCs*



# DYNAMIC WITH EUCLID GLOBULAR CLUSTERS

Constraining the nature of dark matter

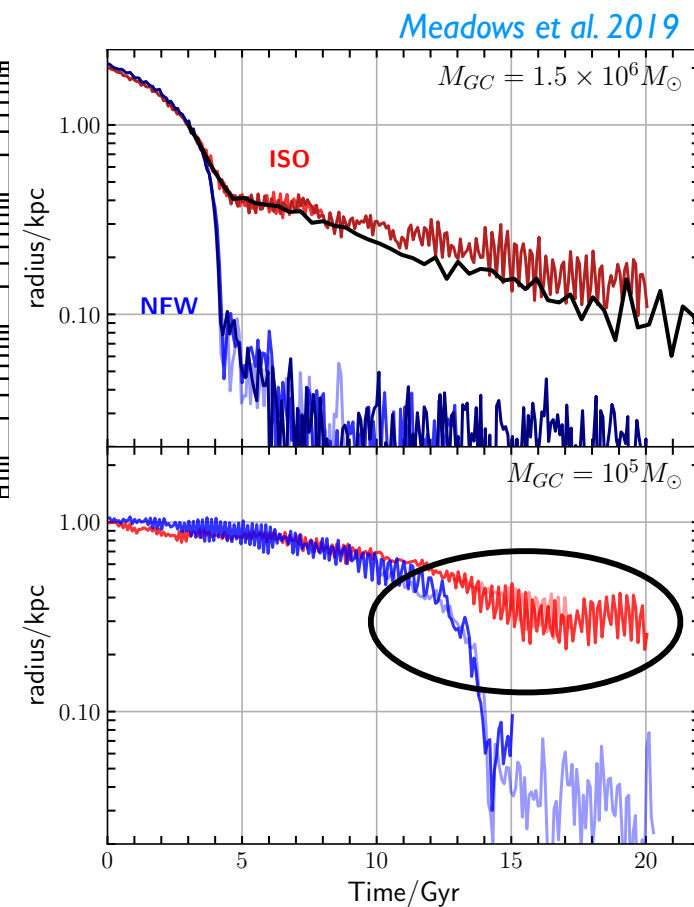
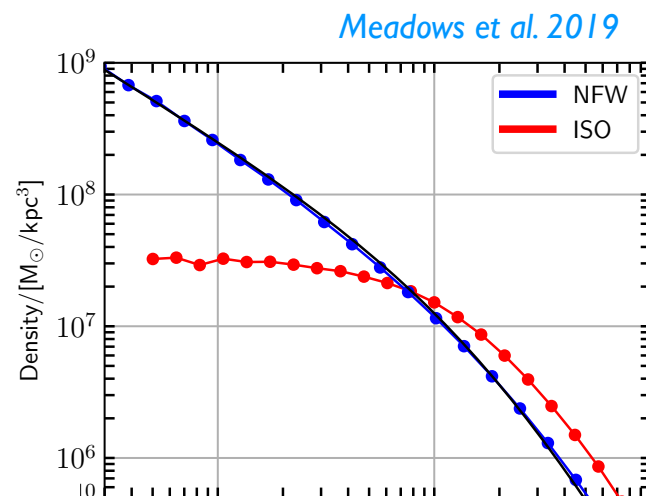
🍊 Host dark matter halo profile by resolving the Timing problem:

« The infall time of GCs can be much shorter than the Hubble time in a NFW halo, so why we observe them? »

from  
the spatial distribution of GCs

Revealing core stalling effect or not!

Cole et al. 2012  
Boldrini et al. 2019  
Sanchez-Salcedo 2022  
Bilek et al. 2021



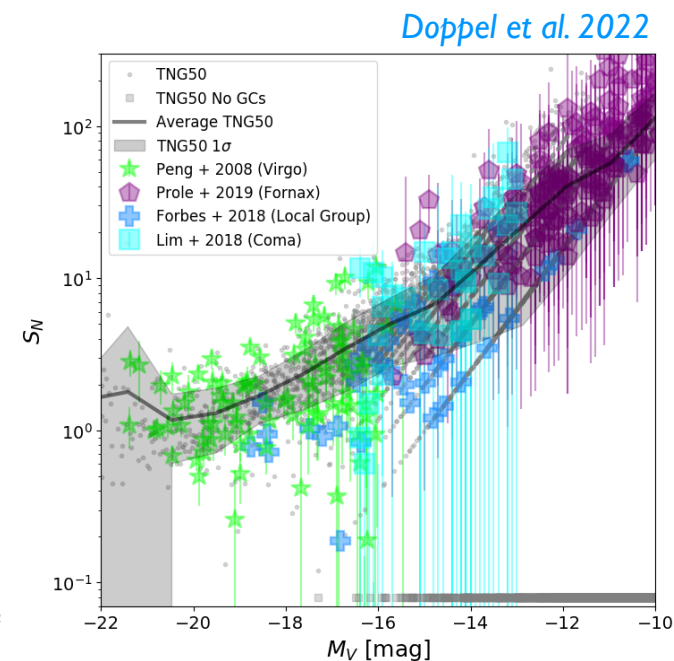
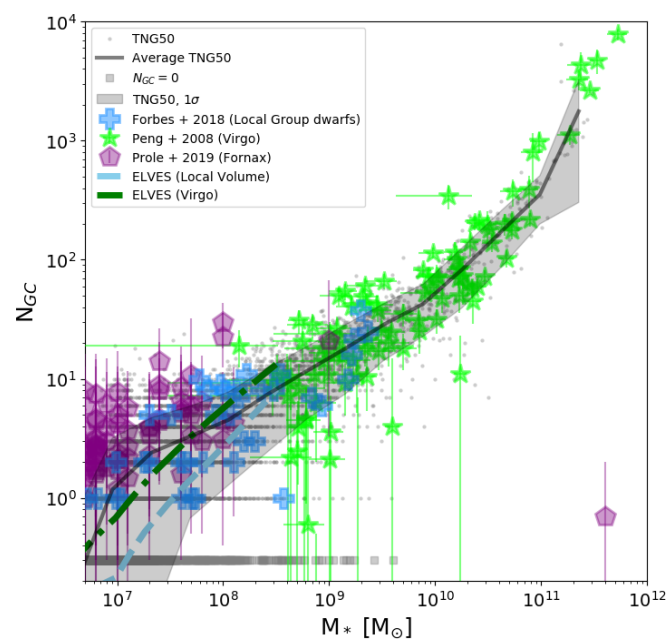
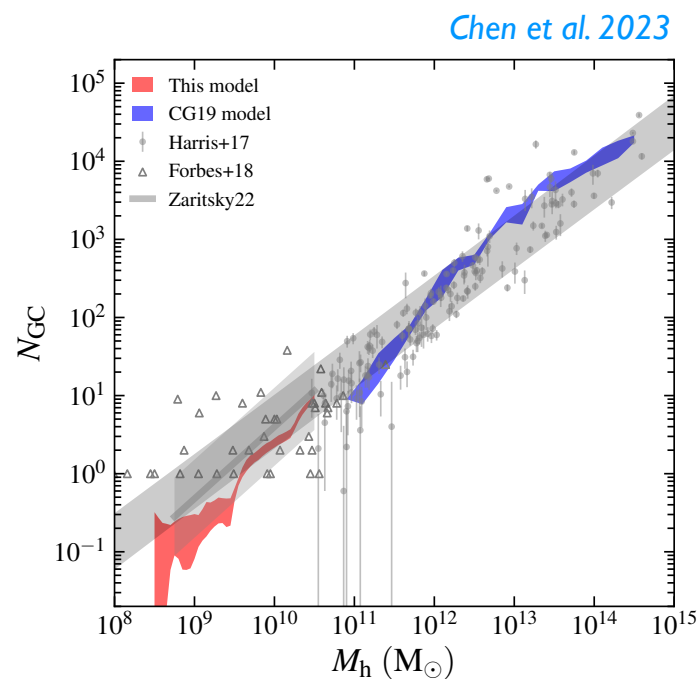
# DYNAMIC WITH EUCLID GLOBULAR CLUSTERS

Constraining our galaxy formation model (& origins of GCs)

🍌 Distinguish dark matter theories and feedback mechanisms from

*the number and specific frequency of GCs especially for dwarf galaxies*

$$S_N = N_{GC} 10^{0.4(M_V+15)}$$



# DYNAMIC WITH EUCLID GLOBULAR CLUSTERS

*Constraining the nature of globular cluster*

 Distinguish their formation mechanism:

- in gas clouds
- in dark matter substructure
- both

From

*their environment (Isolated GCs? Close to dwarfs? At different redshift?)*

*Sameie et al. 2022*

*Vitral & Boldrini et al. 2021*

# DYNAMIC WITH EUCLID GLOBULAR CLUSTERS

*Proposals for specific GC targets*



GC velocity

**JWST**



Profile for close GCs?

**JWST**