

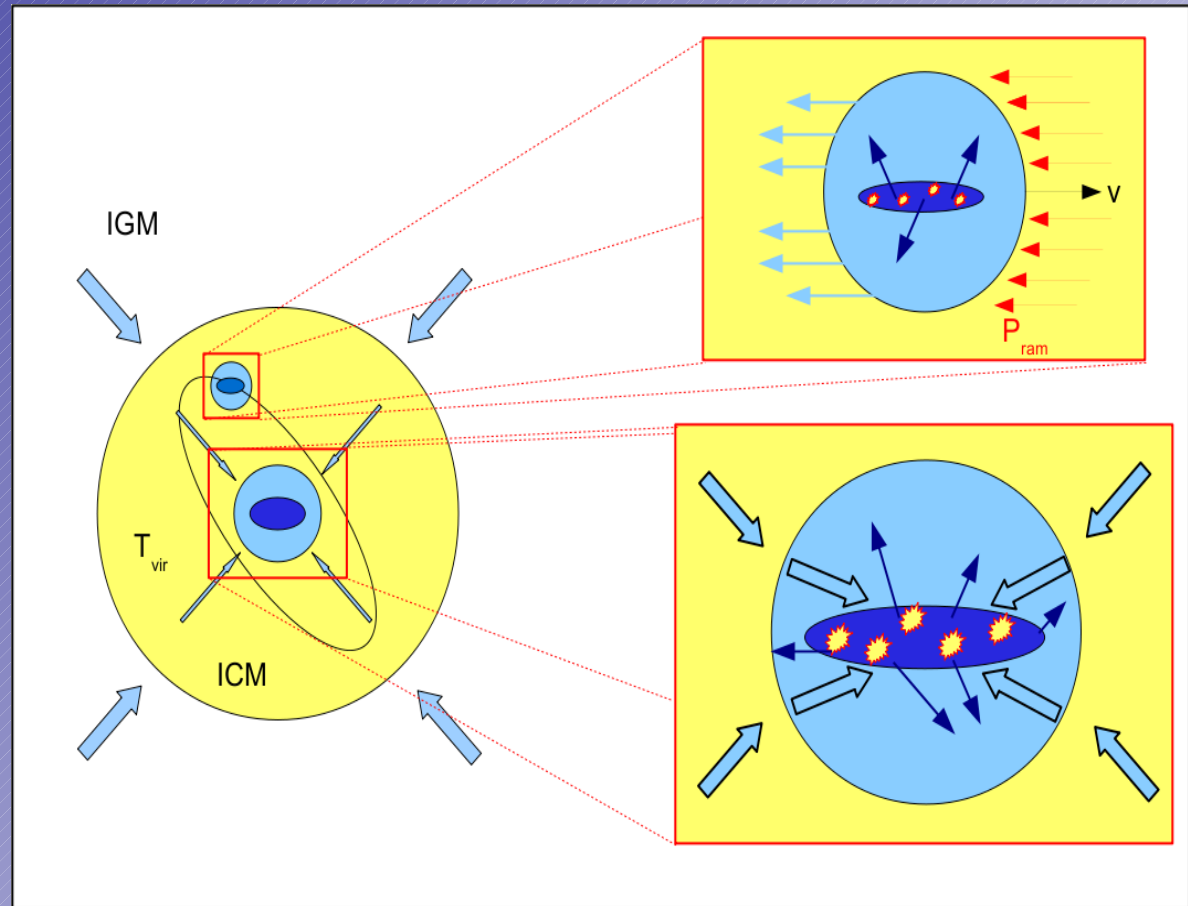
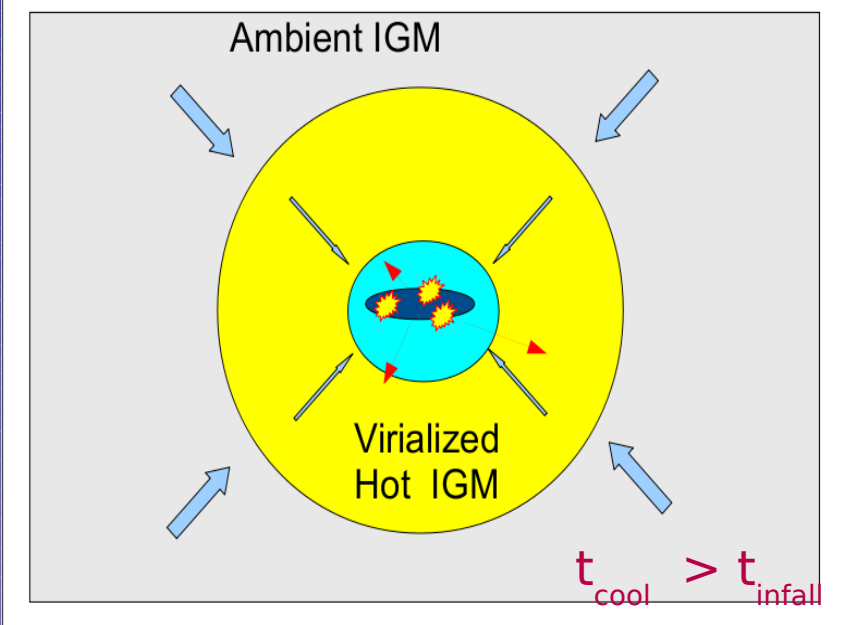
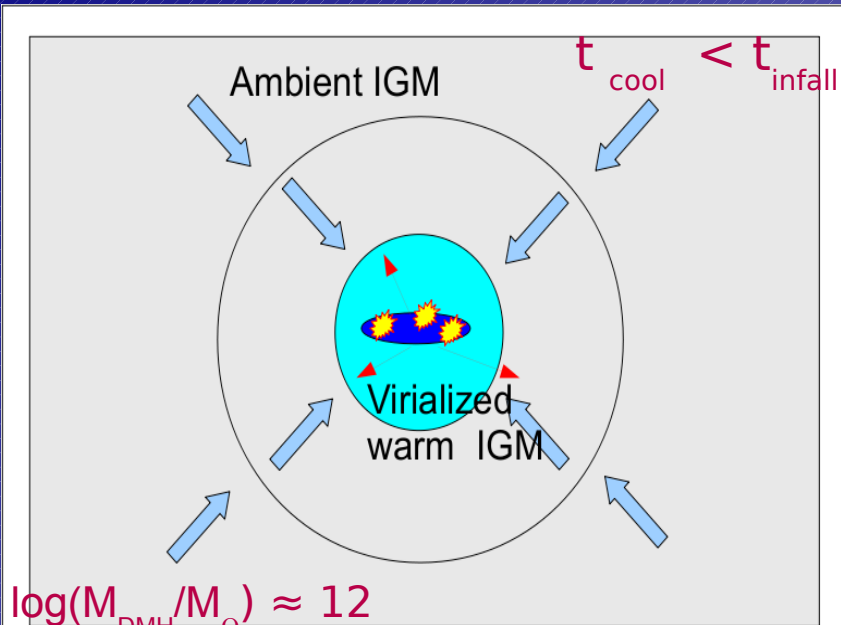
# Gas-Fuelling in the Group Environment - The Star Formation Activity of Satellite Spiral Galaxies

Meiert Grootes (ESA/ESTEC)

With

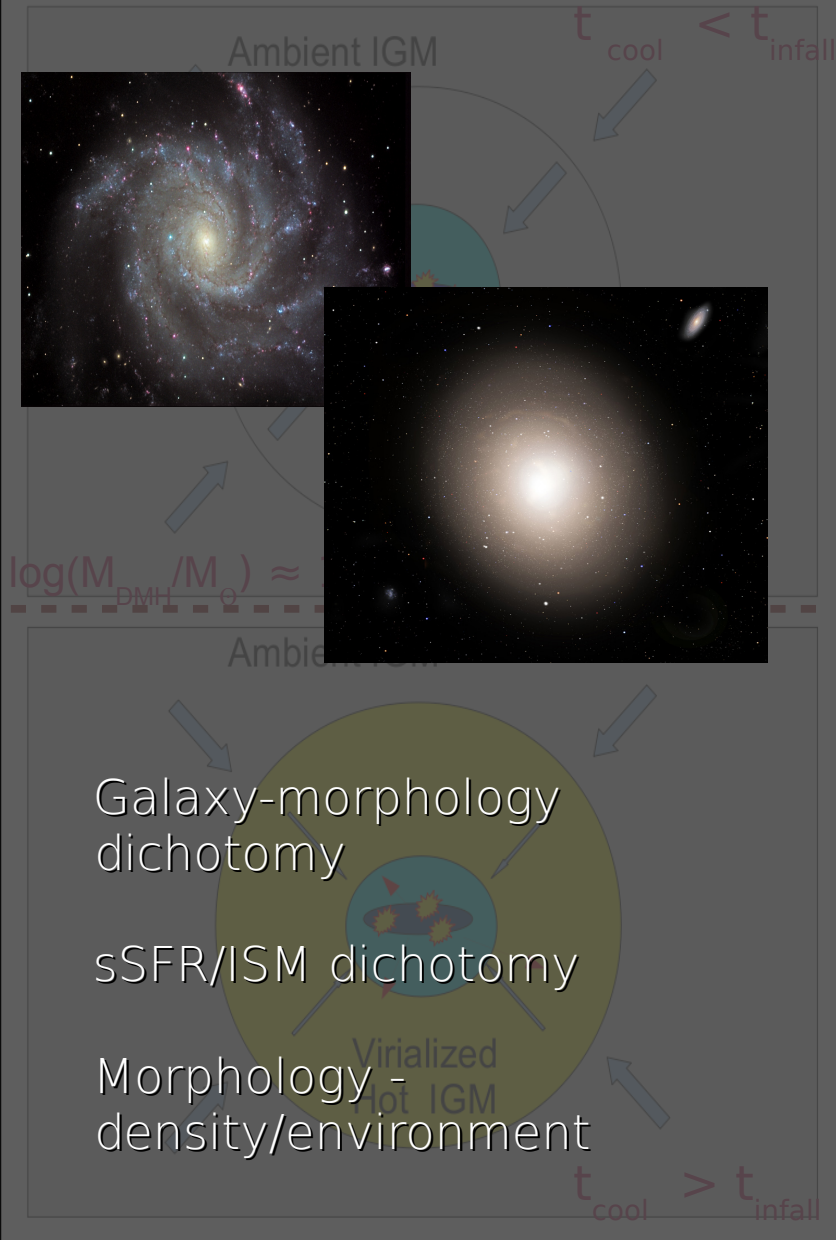
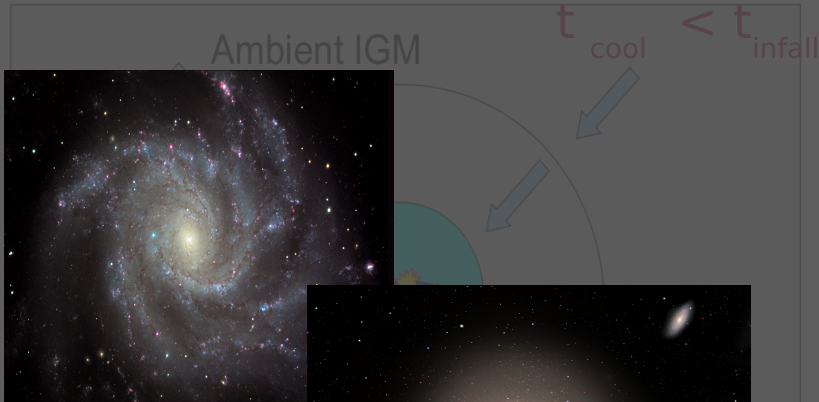
Richard Tuffs (MPIK), Cristina Popescu (UCLan), Jochen Liske (Uni Hamburg), & GAMA

# Galaxy Fuelling and the Gas-Cycle



# Constraining Galaxy Fuelling and the Gas-Cycle ?

A selection of difficulties



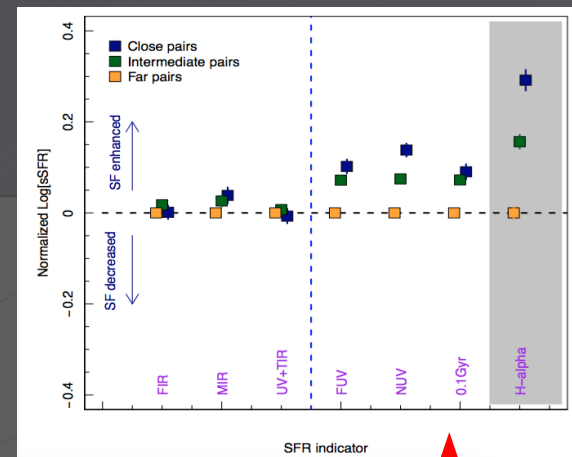
Galaxy interactions & sSFR

Mergers/close pairs

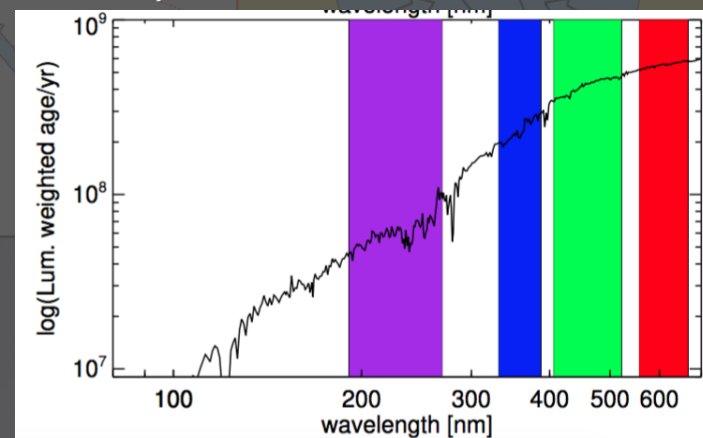
(E.g. Alatalo 2016, Bitsakis 2016, Davies 2015)

Need proxy for gas content → SFR

SFR timescale  $\ll \tau_{\text{dyn}}$  ( $\sim 1$  Gyr)



NUV  
u  
g  
r

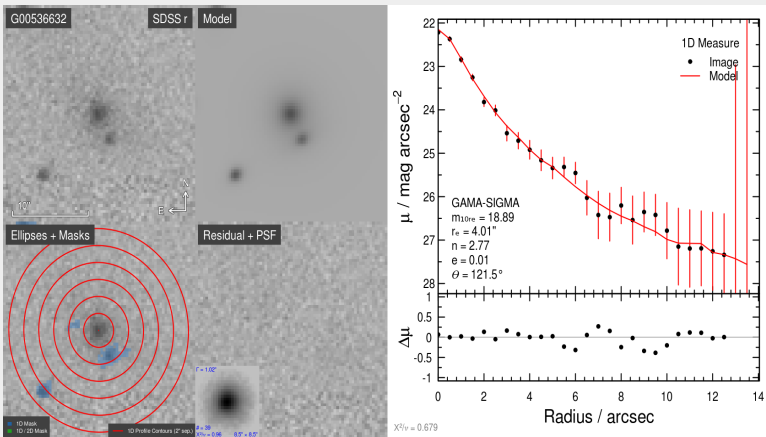


# Constraining the Galaxy Fuelling and the Gas-Cycle ?

## ~~A selection of difficulties~~ GAMA's Solutions

Complete and SFR unbiased selection of disk dominated galaxies (Grootes+2014)

GAMA structural measurements and broadband photometry



Kelvin+2011

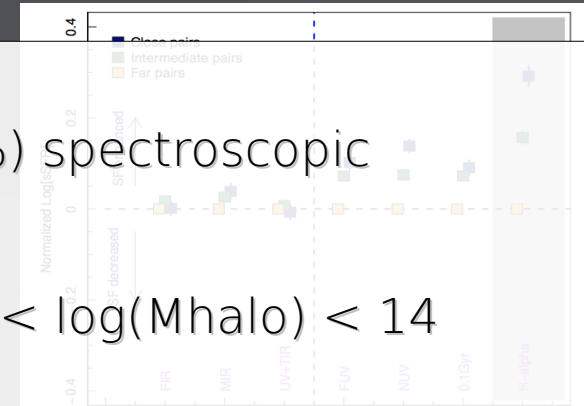
GAMA's high (> 98%) spectroscopic completeness

Galaxy interactions &

SS Group catalogue  $12 < \log(M_{halo}) < 14$

M Exclude (close; 50 kpc/h 1000km/s) pairs

(E.g. Alatalo 2016, Bitsakis 2016, Davies 2015)

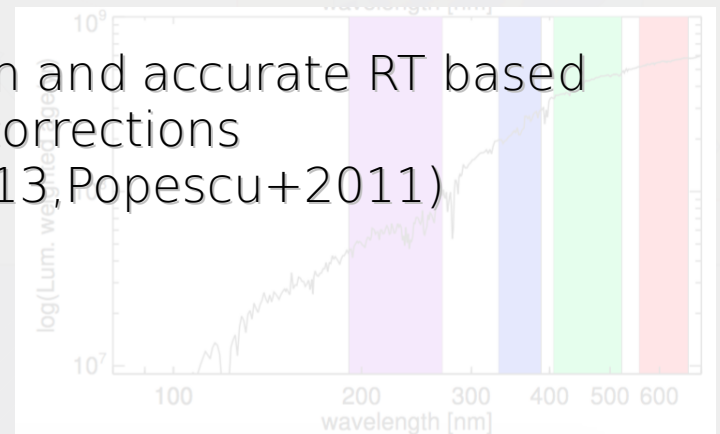


SFR timescale  $\ll \tau_{dyn}$  ( $\sim 1$  Gyr)

Integrated NUV (GALEX;  $\sim 10^8$  yr)

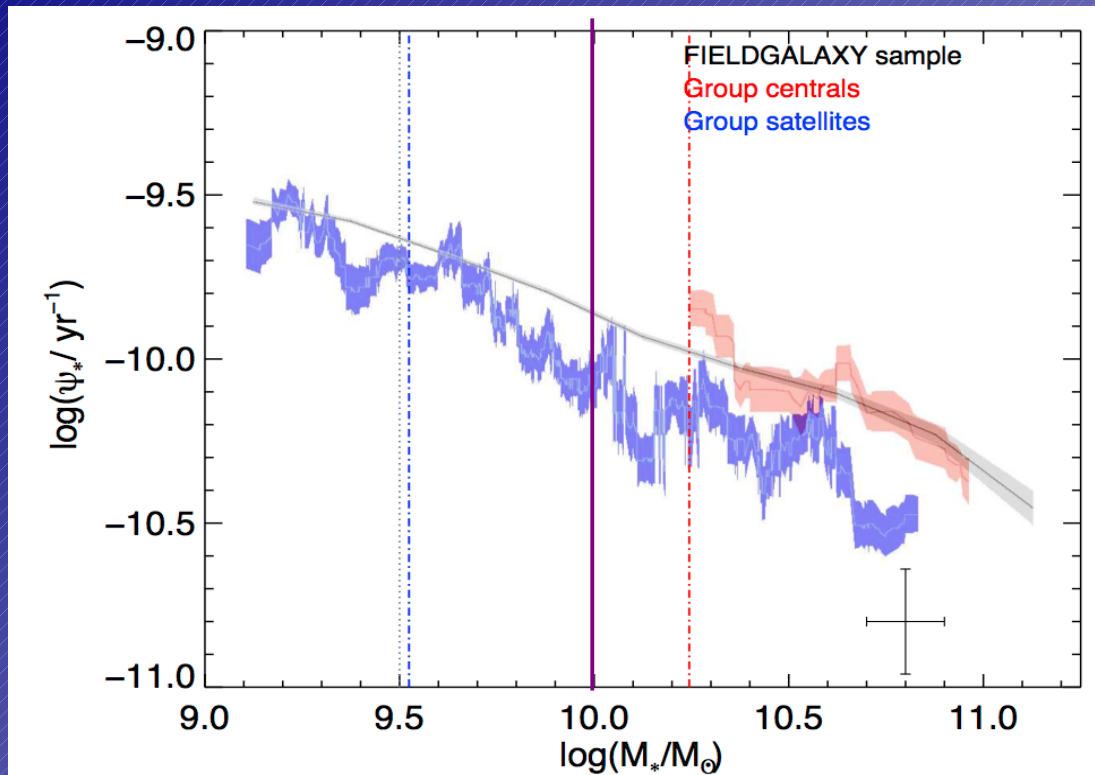
High precision and accurate RT based attenuation corrections

(Grootes+2013, Popescu+2011)



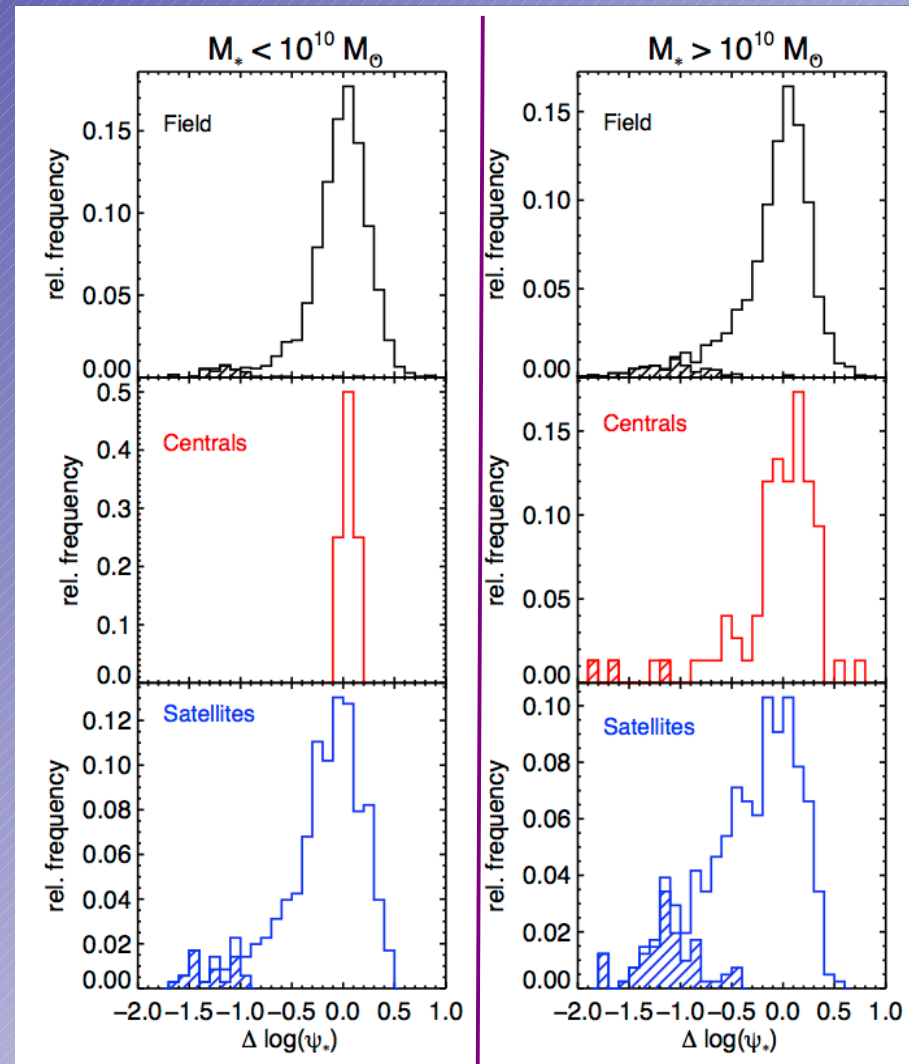


# Impact of the Environment: Satellite & Central Disk-dominated Spirals



disk-dominated fraction only decreases by 40% w.r.t field

Large fraction of disks have spent Gyrs as satellites



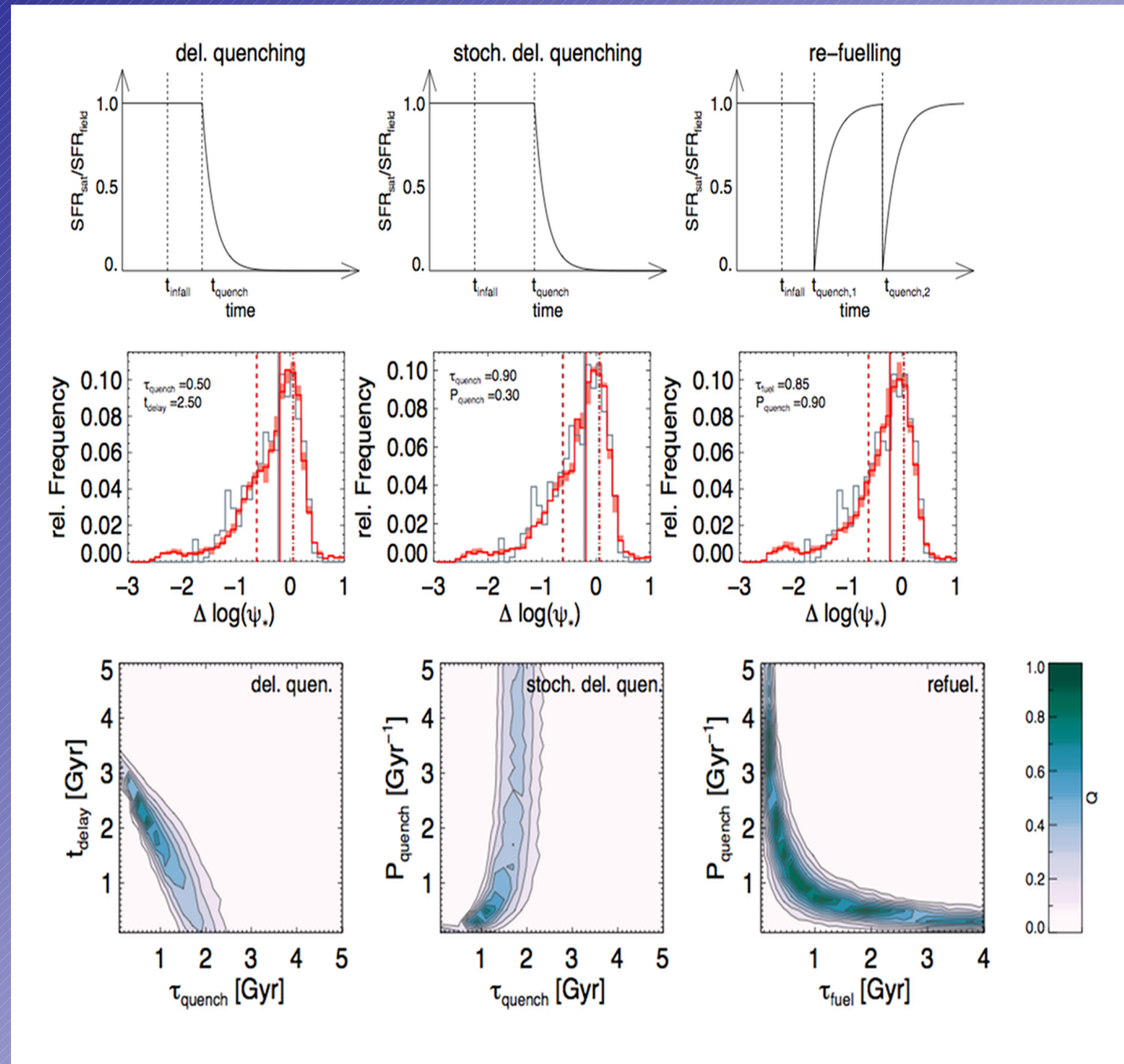
# Prolonged Satellite Star Formation and External Gas Reservoirs

All models favour extended SF at the level of field galaxies and rapid quenching

Rapid cycle of gas with in-&outflows > SFR implied

Reservoirs of gas extraneous to galaxy/ISM needed to fuel SF.

Mass dependence of dsSFR and reservoir and stripping considerations favour IHM as source



## Conclusions / Implications

Gas-fuelling on-going in (disk) satellites, largely independent of environment

Important/dominant contribution of fuelling from group intra-halo medium

Implies two phase IHM (hot/warm & cold phase)

This accretion of (zero net angular momentum) gas poses possible mechanism to transform morphology

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Morphology plays important role in determining SF activity

Color-density arises from morphology-density

# Conclusions / Implications

Come have a look at my poster and check out the imminent paper on arXiv

Also check out the poster w.r.t the effect of central AGN on star formation in satellite galaxies

