The CLEF-SSH simulation project: Modeling the SZ cluster population with radiative gas cooling and energy feedback

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The CLEF-SSH simulation project:

(<u>CL</u>uster <u>E</u>volution & <u>F</u>ormation using <u>S</u>upercomputer <u>S</u>imulations with <u>Hydrodynamics</u>)

•CLEF collaboration: IAS, Orsay (FR) - LA2T, Toulouse (FR) -Sussex (UK) using <u>existing</u> French computing facilities.

•Objectives:

Perform large <u>hydro simulations</u> of LSS that include models of <u>radiative cooling</u> and <u>energy feedback</u> to study:
Cluster physics & scaling laws
Map making (e.g. SZ effect)

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•Core Members:

Antonio da Silva, IAS, (PI)
Nabila Aghanim, IAS
Jean-Loup Puget, IAS
Alain Blanchard, Toulouse
Rachida Sadat, Toulouse

Scott Kay, Univ. Sussex
Peter Thomas, Univ. Sussex
Andrew Liddle, Univ. Sussex

•First proposal: To CINES, Montpelier (PI: Antonio da Silva).

•Designed as part of our contribution to the PLANCK WG5 (clusters and secondary anisotropies) simulation effort

The Planck WG5 workpackage breakdown:

5.5.1.1 perform hydro- simulations of individual clusters

5.5.1

5.5.1.2 perform hydro- simulations of large volumes

5.5.1.3 simulate SZ maps at full numerical resolution

5.5.1.4 simulate Planck observations of SZ maps

> 5.5 Physical Simulations

5.5.2.1 model cluster dist. in semi- analytic way

5.5.2 5.5.2.2 evaluate SZ power

spectrum using SZ maps

5.5.2.3 assess feasibility of cluster velocity studies

5.5.6.1 realistic modeling of ionisation history

5.5.6

5.5.6.2 simulate radiative transfer through IGM

5.5.6.3 simulate Planck observations in detail

5.5.6.4 investigate effect of systematics on science 5.5.3.1 construct approx. template maps

5.5.3

5.5.3.2 update using largescale N- body simulations

5.5.3.3 update using largescale hydro simulations

5.5.3.4 include contaminating effects

5.5.3.5 simulate X- ray and optical cluster props.

5.5.3.6 model selection functions using observations

5.5.3.7 develop tools for cluster extraction from data 5.5.5.1 perform N- body simulations of large- scale struct.

5.5.5

5.5.5.2 construct 3- dim cluster catalogues

5.5.5.3 simulate observable cluster signatures

5.5.5.4 apply and test cluster- detection algorithms

5.5.5.5 combine with highres. hydro cluster sims.

5.5.5.6 select, develop, and supply required tools

First simulation: the CLEF run1

Cosmology: Flat Lambda CDM model with:
 ·Ωm=0.3; ΩΛ=0.7; Ωb=0.0486; σ8=0.9; h=0.7
 ·Simulation details:

•L=200 Mpc/h ; 2(428)^3 (156 Million) particles (mgas =7e9Msun/h; mdark=1.4e10Msun/h); Soft.= 20kpc/h

•Radiative cooling and energy feedback (Kay 2004)

•Simulation code:

•A parallel version of GADGET II (Springel, Yoshida & White, 2001), PM-Tree code with SPH.

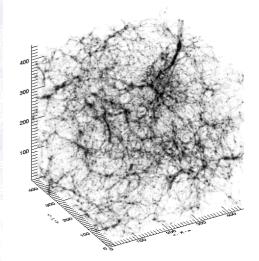
•System resouces:

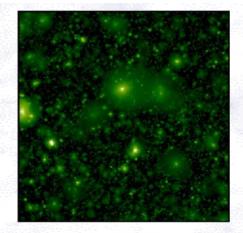
•66000 CPU hours (SGI R14K); 0.5 Tb of data

•Permits to obtain:

•extensive cluster catalogues (4000 objects @ z=0 resolved with at least 1000 particles each).

Maps of 5 deg². In a second step: 10deg²





First glimpse at the CLEF run:

•CLEF collaboration: IAS (FR) - Toulouse (FR) - Sussex (UK) using French system CINES

•Objectives: Perform high resolution simulations to study:

•Cluster physics & scaling laws

•Map making (e.g. SZ effect)

•First run:

•L=200 Mpc/h ; 2(428)^3 particles

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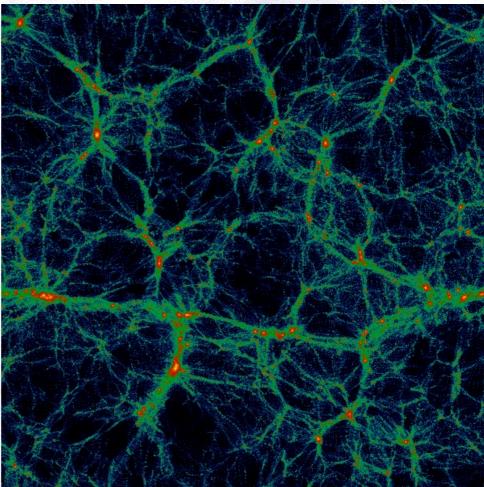
•GADGET (Springel, Yoshida & White, 2001). Tree-SPH code

•Permitted to generate:

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•Maps of 5-10 deg^2

Dark matter



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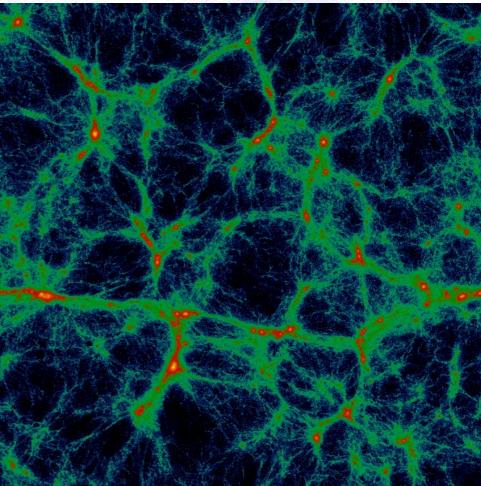
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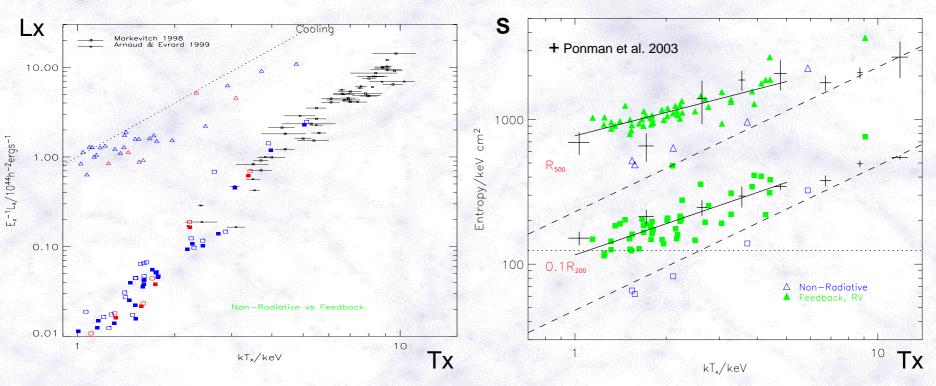
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Baryonic gas



Results from test runs: (fheat=0.1, Sheat=1000 keV cm2)



·2 parameter energy feedback model (S. Kay 2004) : fheat, Sheat

Cold dense gas particles (T<1e5, n>1e-3 cm^-3) have fheat probability of being heated up with an entropy Sheat.

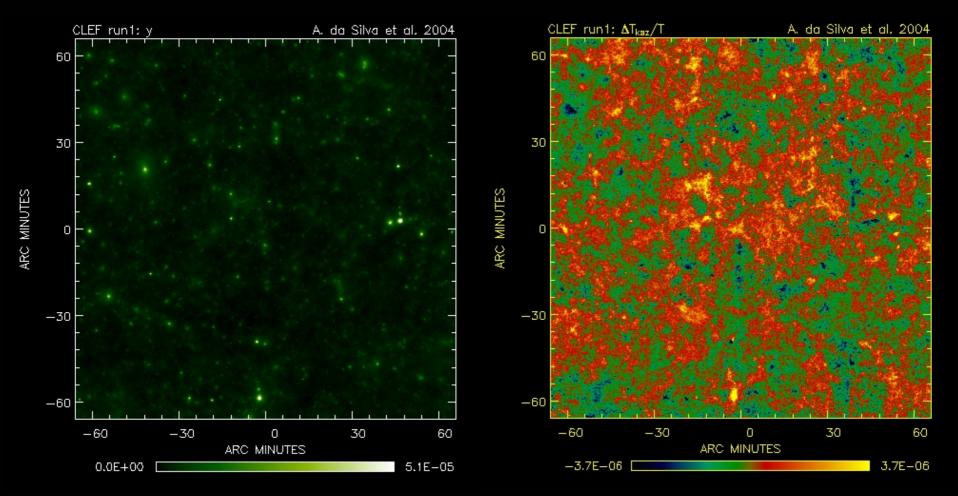
•Draw random O<r<1 (uniform dist.) for each gas particles with T<1e5, n>1e-3 cm^-3

- ·If r<fheat then reheat particle by the fixed entropy Sheat
- ·If r>fheat then convert particle to colisionless baryonic material
- •fheat: Largely controls the amount of cooled gas fraction
- •Sheat: determines de amount of energy feedback

Preliminar results: SZ scaling laws @ z=0

CLEF Y-M results + CLEF: feedback 0 6730: preheating 10-4 ++ da Silva et al 2003 + non-rad. o cooling 10⁻⁵ 10-5 r'żła / (h−1 Mpc)² (h⁻¹ Mpc)² 10-6 10⁻⁶ Y^{int} / 10-7 10-7 10⁻⁸ 10-8 10¹⁴ 1013 1014 1015 10 10 $M_{200} / (h^{-1} M_{sun})$ $M_{200} / (h^{-1} M_{sun})$ $Y \propto M^{5/3} \propto T^{5/2} \propto L_{_X}^{5/4}$ $Y \propto f_{\rm gas} M T_{\rm mw} d_A^{-2} \propto f_{\rm gas} M^{5/3} d_A^{-2} (1+z)$ Non_radiative slope = 1.69Energy feedback slope = 1.95Radiative slope = 1.79Preheating slope = 1.93

Preliminar results: 5 deg^2 SZ maps



ymean = 3.23 E-6; Sigma_y = 1.96 E-6

Sigma_k = 9.65 E-7

Concluding remarks:

Cluster scaling laws from hydrodynamical simulations:

- Radiative gas physics (Cooling and energy feedback) has non negligible impact in cluster scaling relations and the SZ effect.
- The magnitude of these effects is not yet well understood and need to be properly assessed with hydro sims. for preparing and interpreting future SZ/CMB observations, such as PLANCK.
- Larger simulations with high-resolution are most welcome! They permit to derive cluster scaling laws in wider mass ranges, with improved statistics.

Hydro maps:

- High-resolution SZ hydro maps are the most powerful tool for studying selection effects of deep surveys (high-angular resolution), complementarity between different observational strategies,...
- So far these maps have been limited to small sky areas (1 deg), which are too small for large beam experiments as PLANCK
- The CLEF run will permit to obtain a maximum 10deg^2 (unprecedented for hydro simulations with cooling and energy feedback)!