

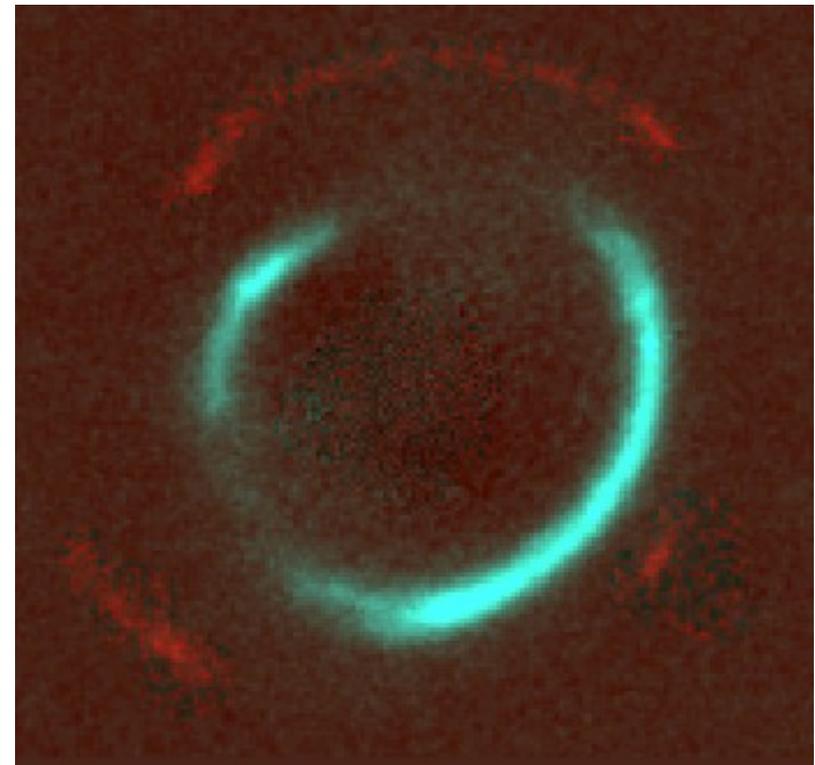
Two rings of light in the galaxy SDSSJ0946+1006

Raphaël Gavazzi



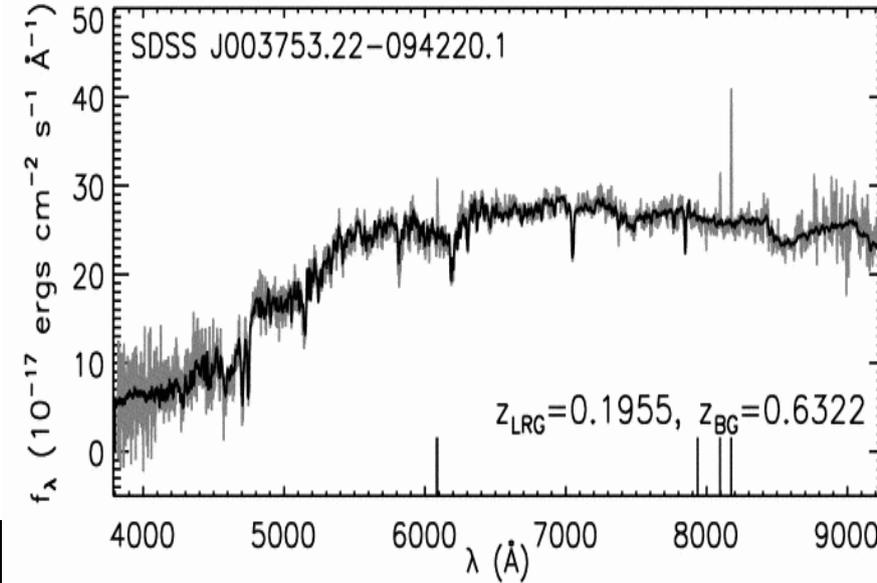
UC Santa Barbara

Collab. T. Treu, P. Marshall, & SLACS team

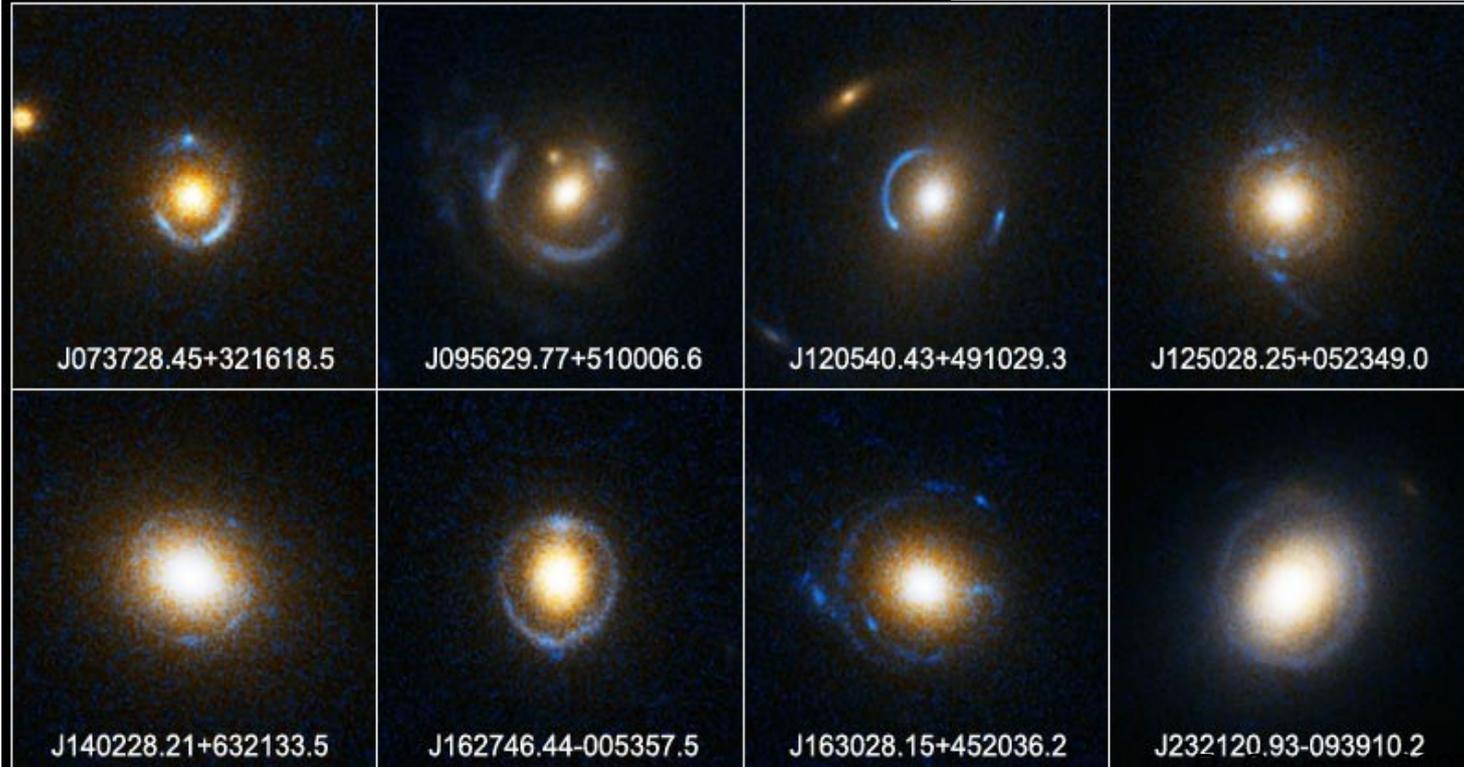


The Sloan Lens ACS Survey (SLACS)

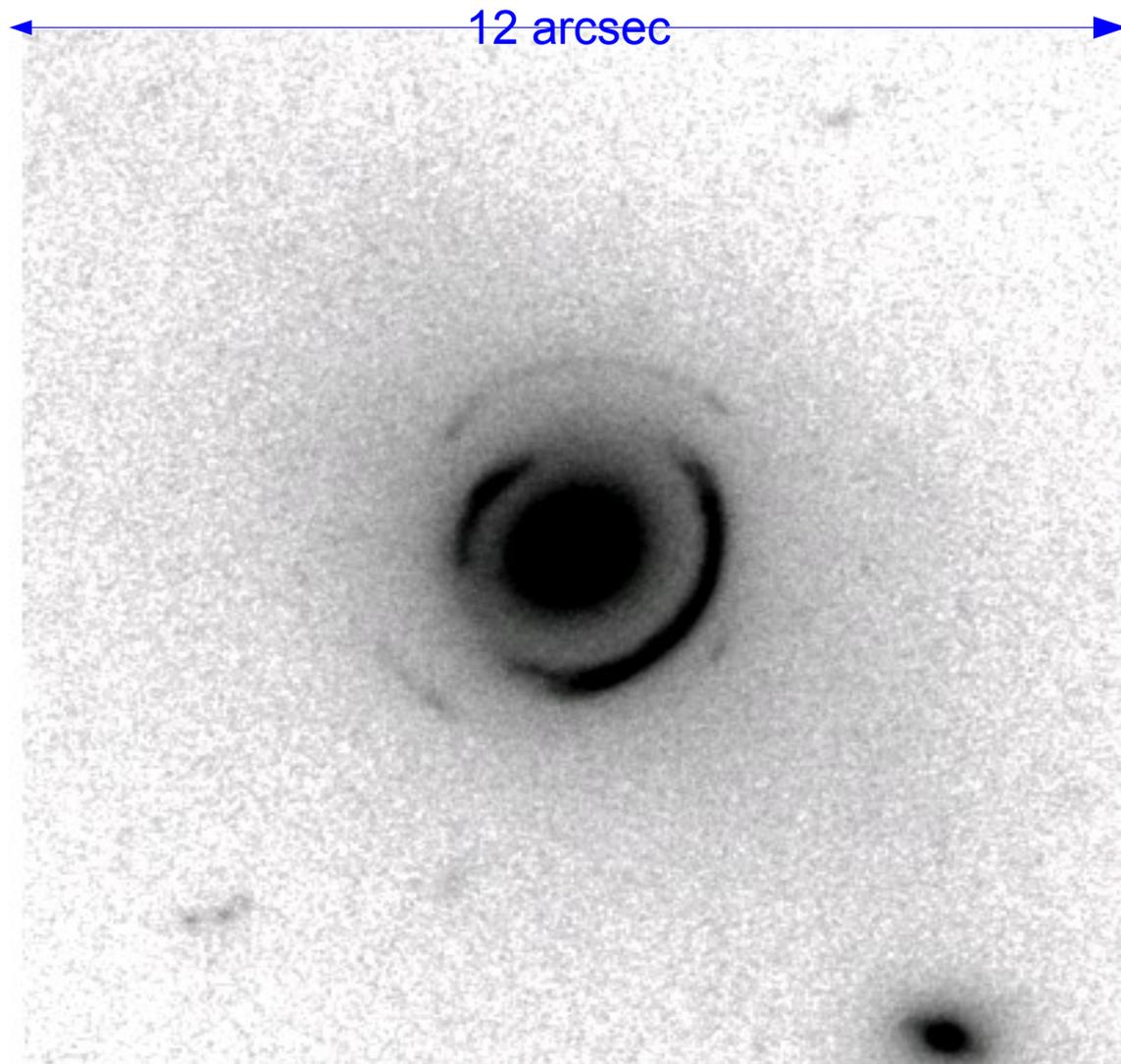
- *Spectroscopic selection in SDSS database: spurious emission lines at $z_s > z_l$*
- *HST follow-up imaging for confirmation and accurate modeling*
- *So far, ~97 new massive lens galaxies ($z_l=0.06-0.4$, $\sigma > 200 \text{ km/s}$, $z_s=0.3-1.1$)*



Einstein Ring Gravitational Lenses



SDSSJ0946+1006



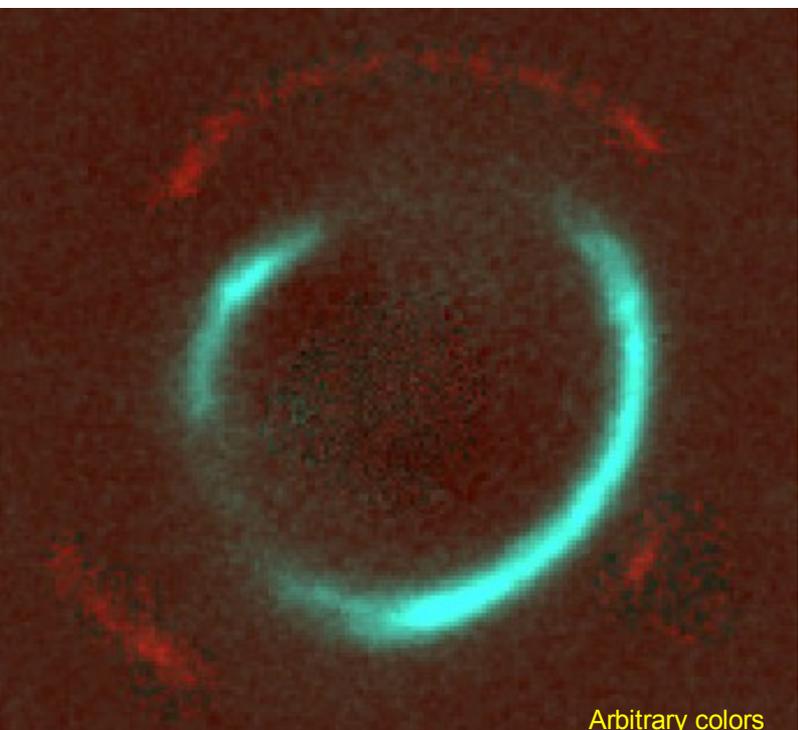
HST/ACS imaging: 2000sec in F814W

SDSS fiber spectroscopy:

$$z_{\text{lens}} = 0.222$$

$$z_{\text{source1}} = 0.609$$

central velocity dispersion: 285 ± 25 km/s



Second source redshift ? unknown!

8 hours LRIS@Keck additional longslit spectroscopy unable to provide z_{source2} . Very faint ($I_{AB} \sim 23 \text{ arcsec}^{-2}$) and no obvious emission lines from 3500 to 8400Å.

SDSSJ0946+1006, the jackpot?

First double galaxy-scale strong lensing event as part of just 50 SLACS systems with deep ACS/F814W imaging!

How likely? $\sim 1/200$ ellipticals is a strong lens would suggest $\sim 1/200$ lensing ellipticals is a double lens...

The abundance of such multiple events tells us about the pop. of faint ($I > 25$) high z sources.



"Well, would you believe it? I won last night's lottery jackpot! "



The jackpot, not just a curiosity....

Double ring tells us about:

- Mass density profile of the lens? YES!
- Cosmography? UNFORTUNATELY NOT
- Mass of the inner ring? APPROXIMATELY...

Lens modeling

Total (DM+stars) density profile

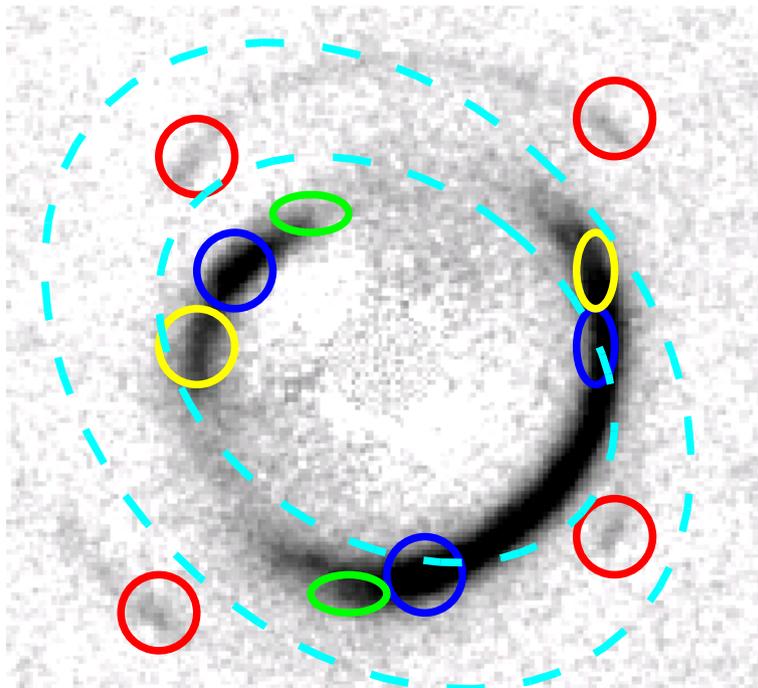
$$\kappa(\vec{r}, z_s) = \frac{b^n}{2} \left(x^2 + y^2/q^2 \right)^{-n/2} \frac{D_{ls}}{D_{os}}$$

$$b = 4\pi(\sigma_v/c)^2 = \left(\frac{\sigma_v}{186.2 \text{ km s}^{-1}} \right)^2$$

$b, n, q, \text{PA}, z_{s2}$

+ external shear $(\gamma_{\text{ext}}, \text{PA}_{\text{ext}})$

} 7 free parameters!

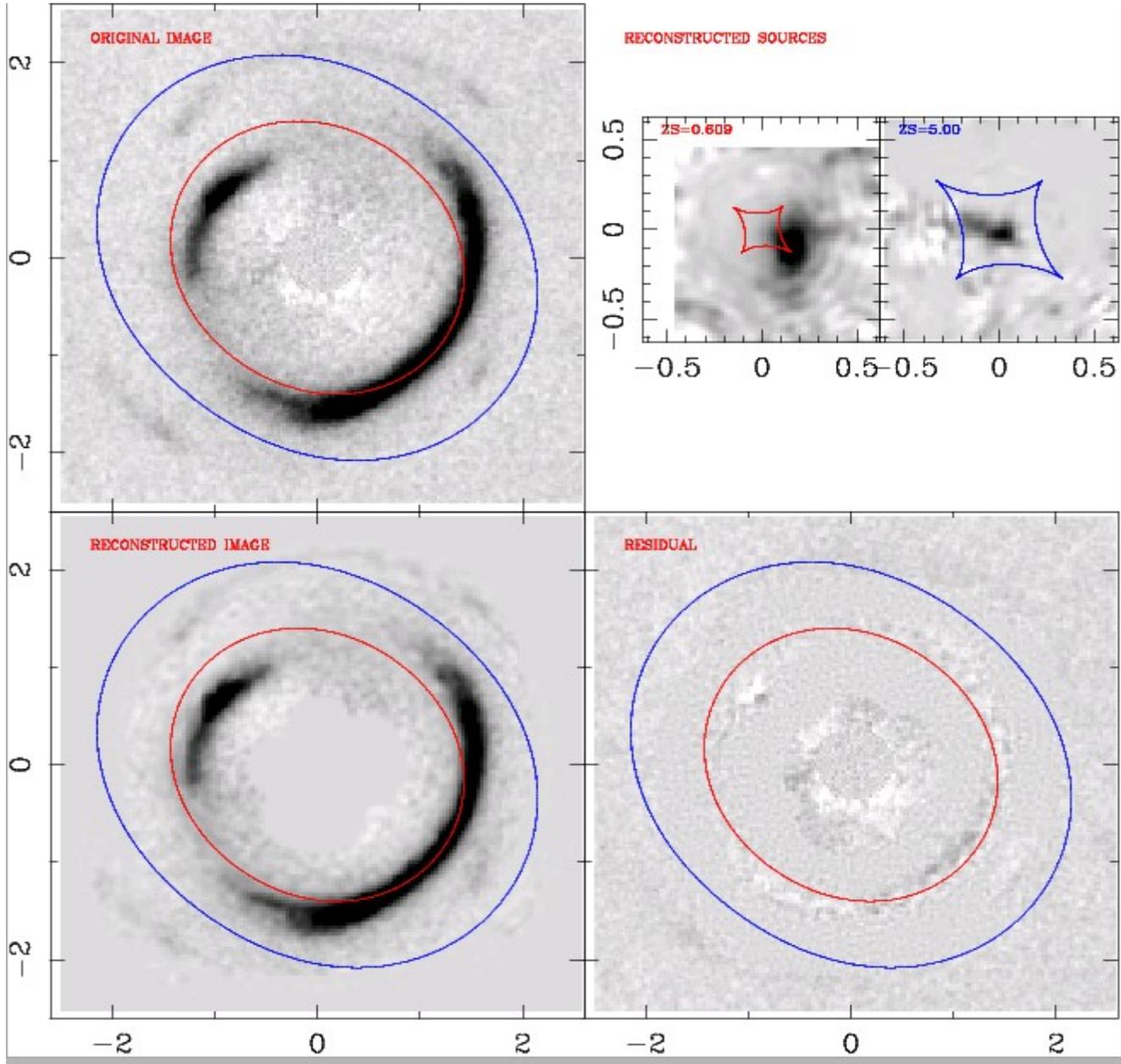


Preliminary analysis:

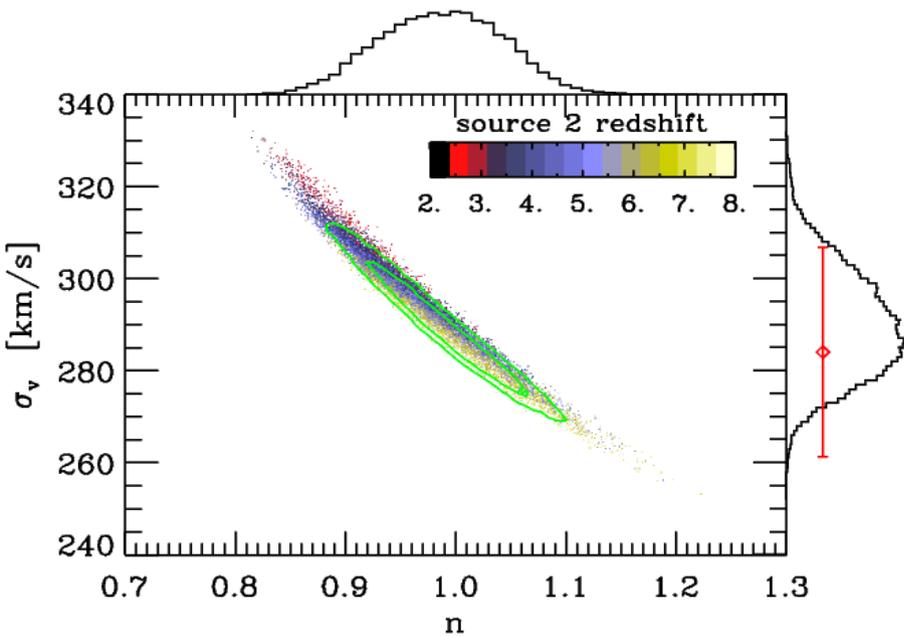
- guess the location of critical lines
- identify conjugate points
- point source χ^2 inversion

Easy and gives fair weight to very dim source2 (1:36 in flux wrt source1)

Confirmation with pixelized source inversion of each source.



Some constraints

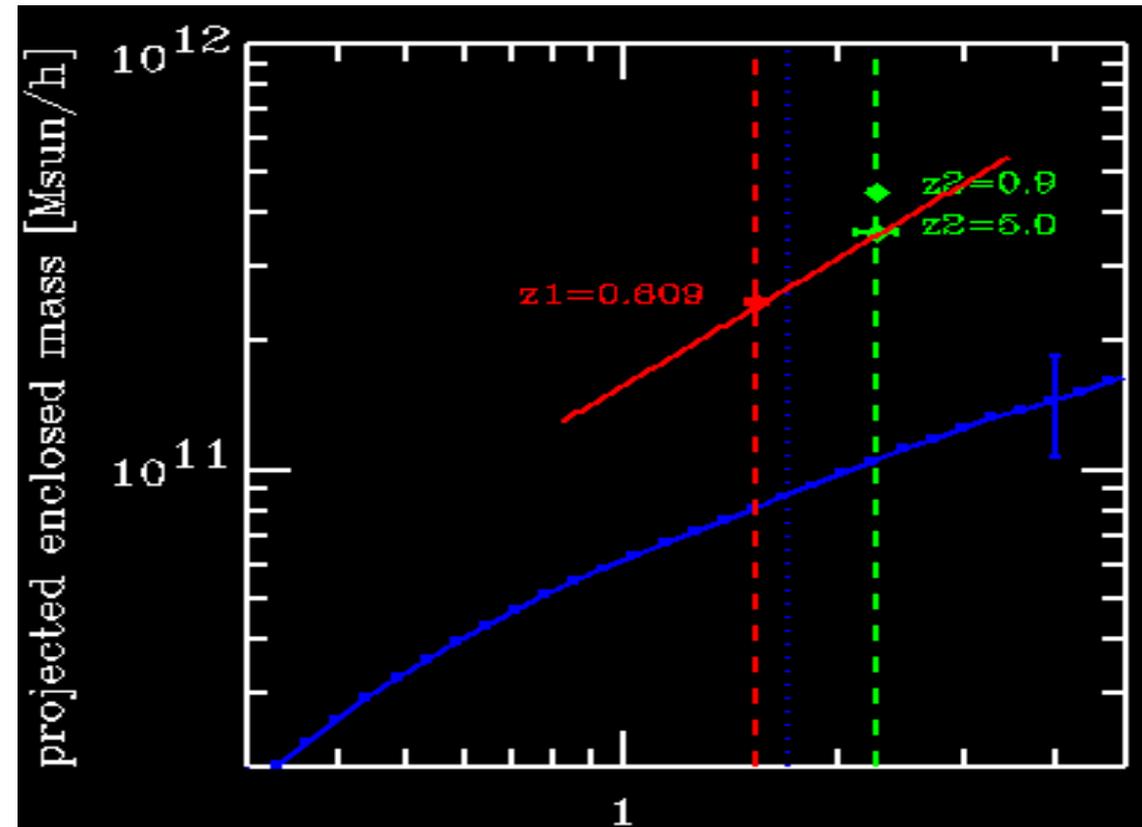
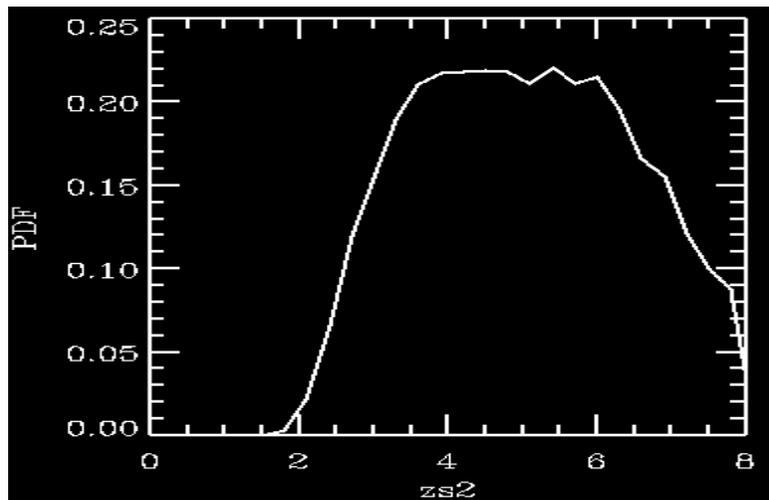


Velocity dispersion 289 ± 11 km/s consistent with SDSS spectro

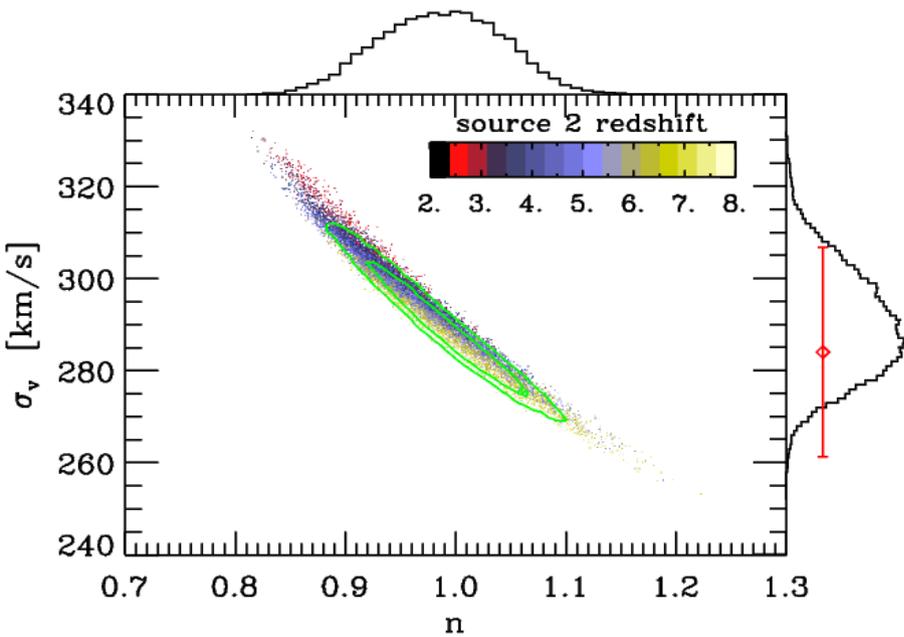
Slope \sim isothermal $n = 0.99 \pm 0.06$.

Source2 redshift loosely constrained but > 3

Fdm($< R_{\text{eff}}$) $\sim 68\%$



Some constraints



Let's push as high as possible stellar M/L_V to reproduce ring1

... then falls short for ring2!!

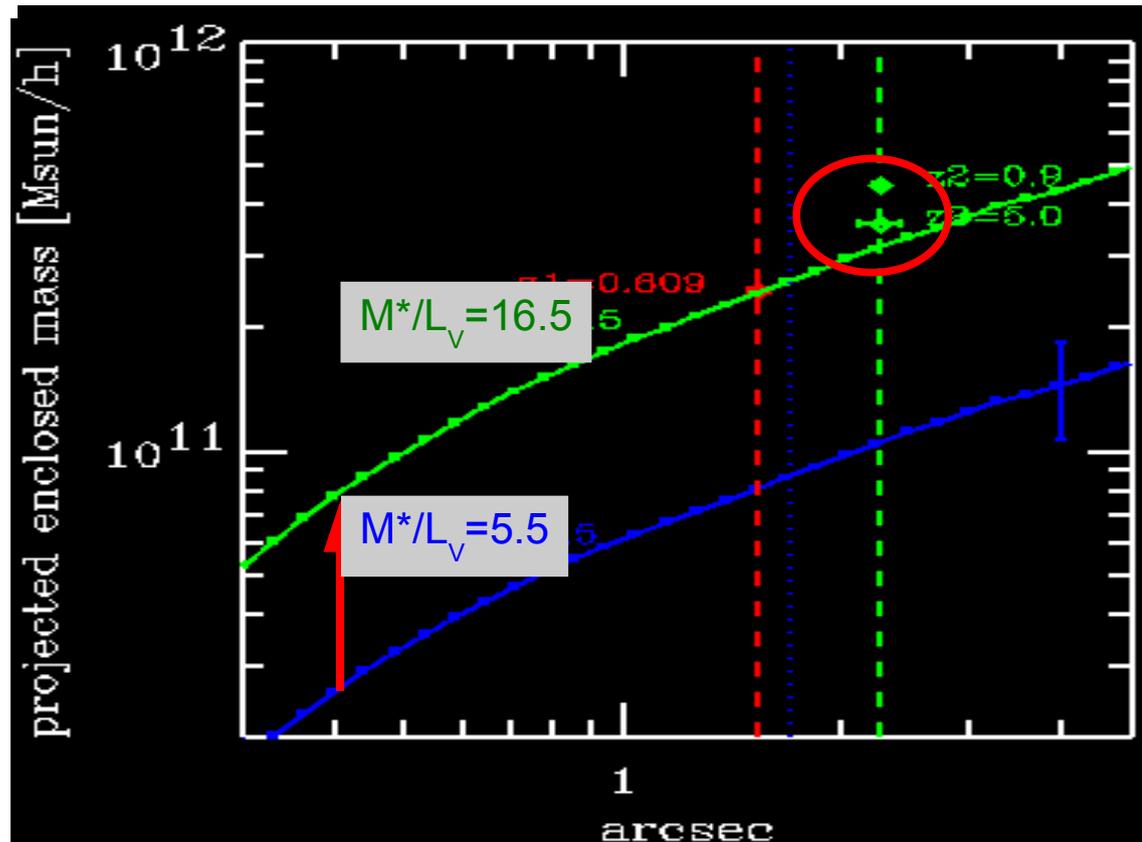
We need extra mass (a halo?!)

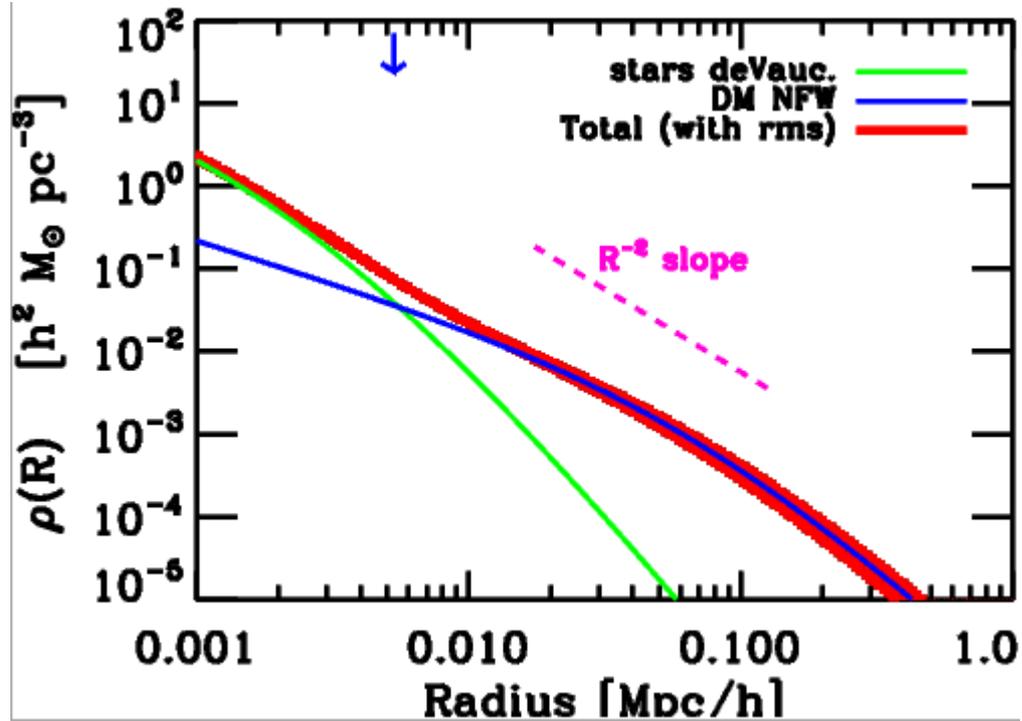
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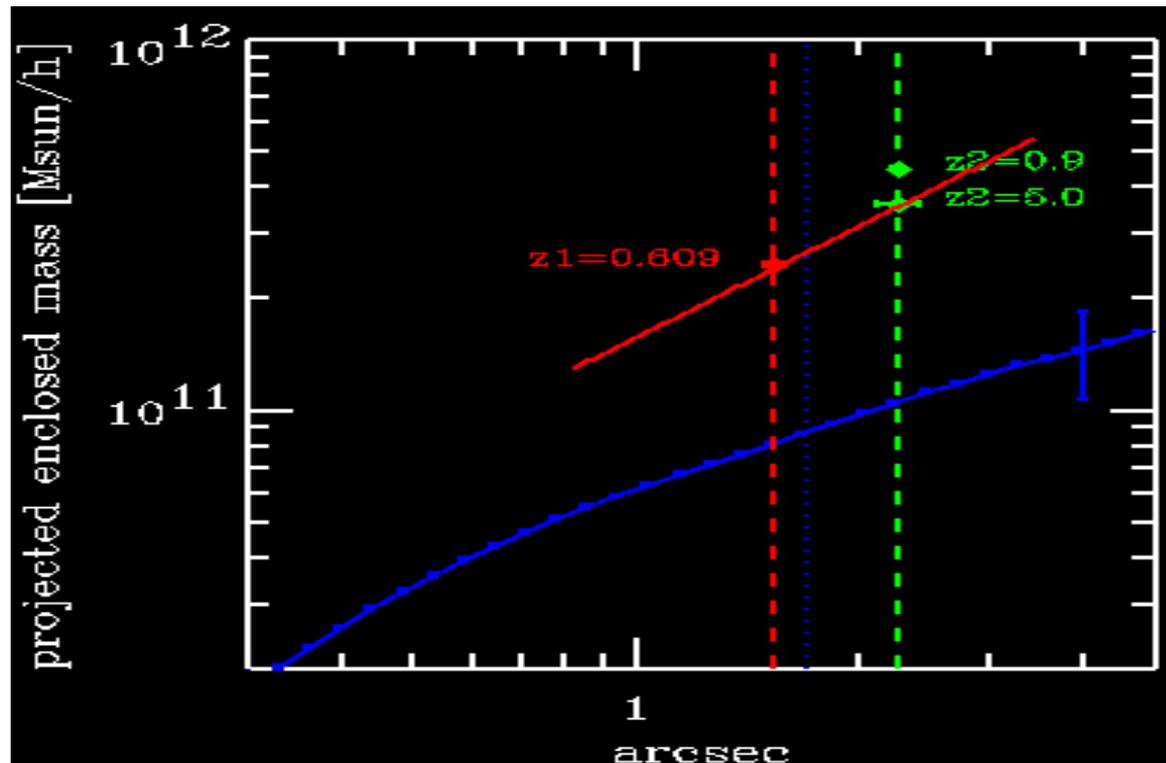
$F_{dm}(<R_{eff}) \sim 68\%$

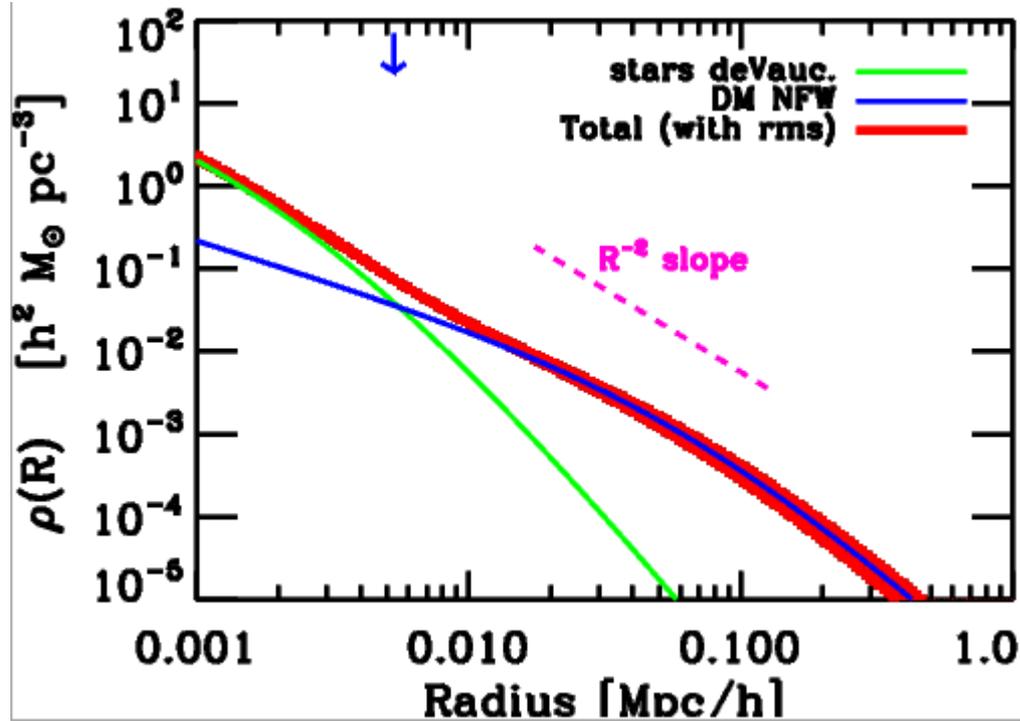




Results from Strong & Weak lensing analysis : Average 3D density profile.

Gavazzi et al 07

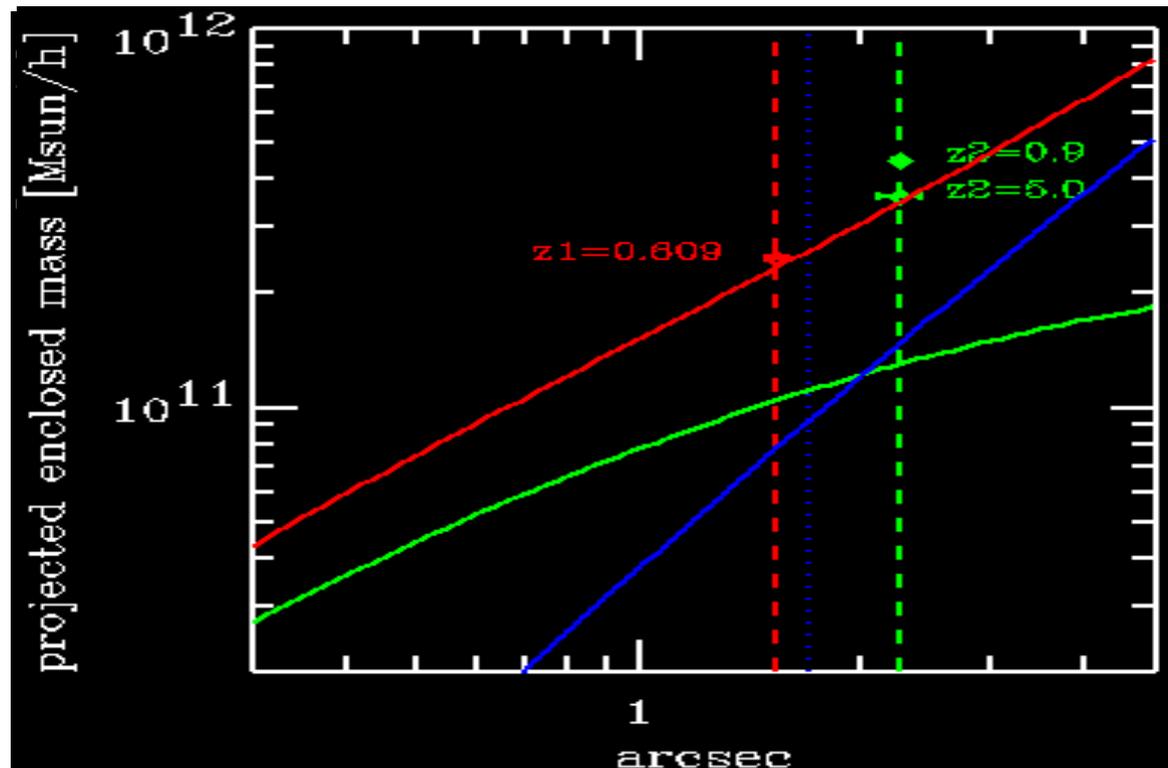




Results from Strong & Weak
lensing analysis : Average 3D
density profile.

Gavazzi et al 07

Excellent
agreement
when using
S+WL results
SLACS once
rescaled to the
actual lens σ_v



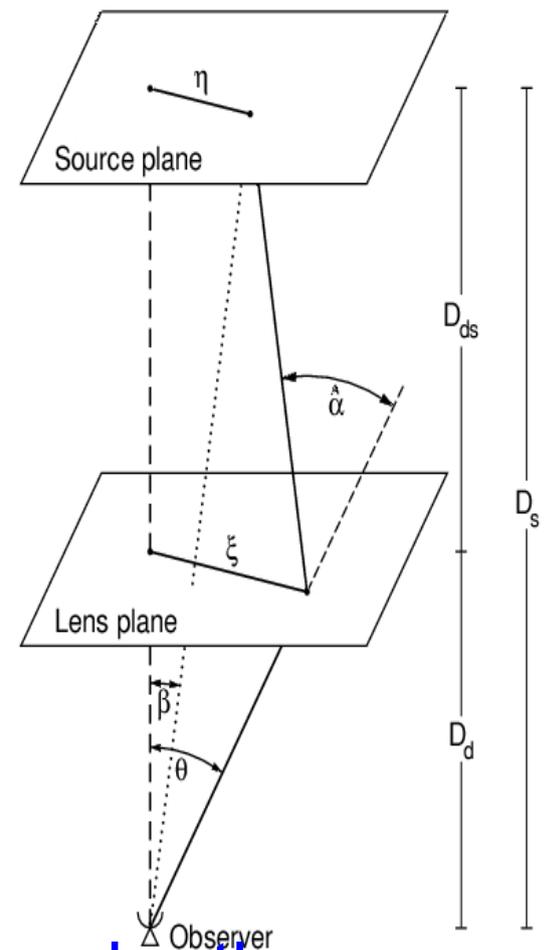
The jackpot, conclusions so far....

- Successful model: it's a double gravitational lens!
- Bulge-halo conspiracy confirmed with **strong lensing only** (isothermal profile)
- Standard light+NFW ok
- What you see is what you get: Strong Lensing needs DM
(true for all SLACS single lenses with sensible M^/L)*

COSMOGRAPHY?

$$\vec{\beta} = \vec{\theta} - \vec{\alpha} \equiv \vec{\theta} - \vec{\nabla} \psi(\vec{\theta}).$$

$$\psi(\vec{\theta}) = \frac{2 D_{ls} D_{ol}}{c^2 D_{os}} \phi(\vec{\theta}).$$



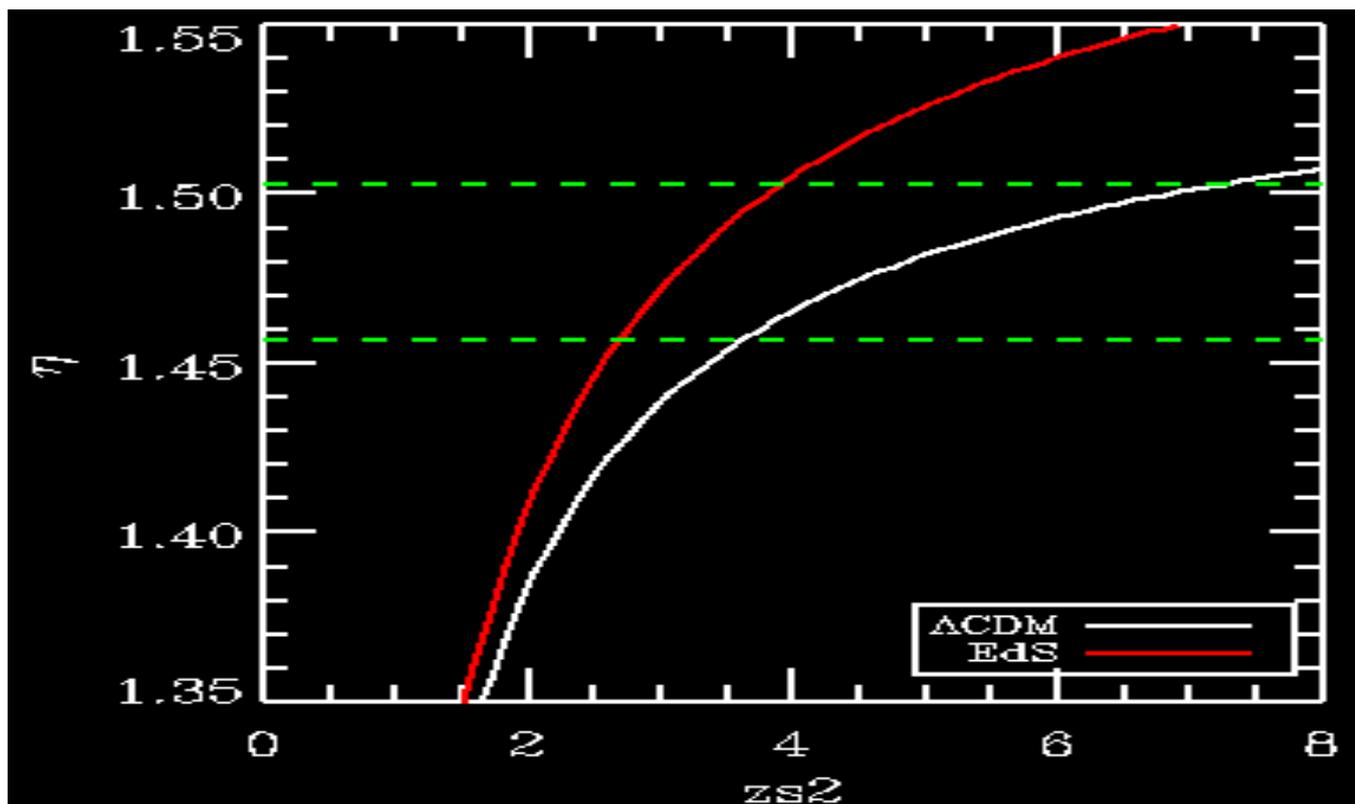
If the potential is known, then we can probe the geometry of the universe.

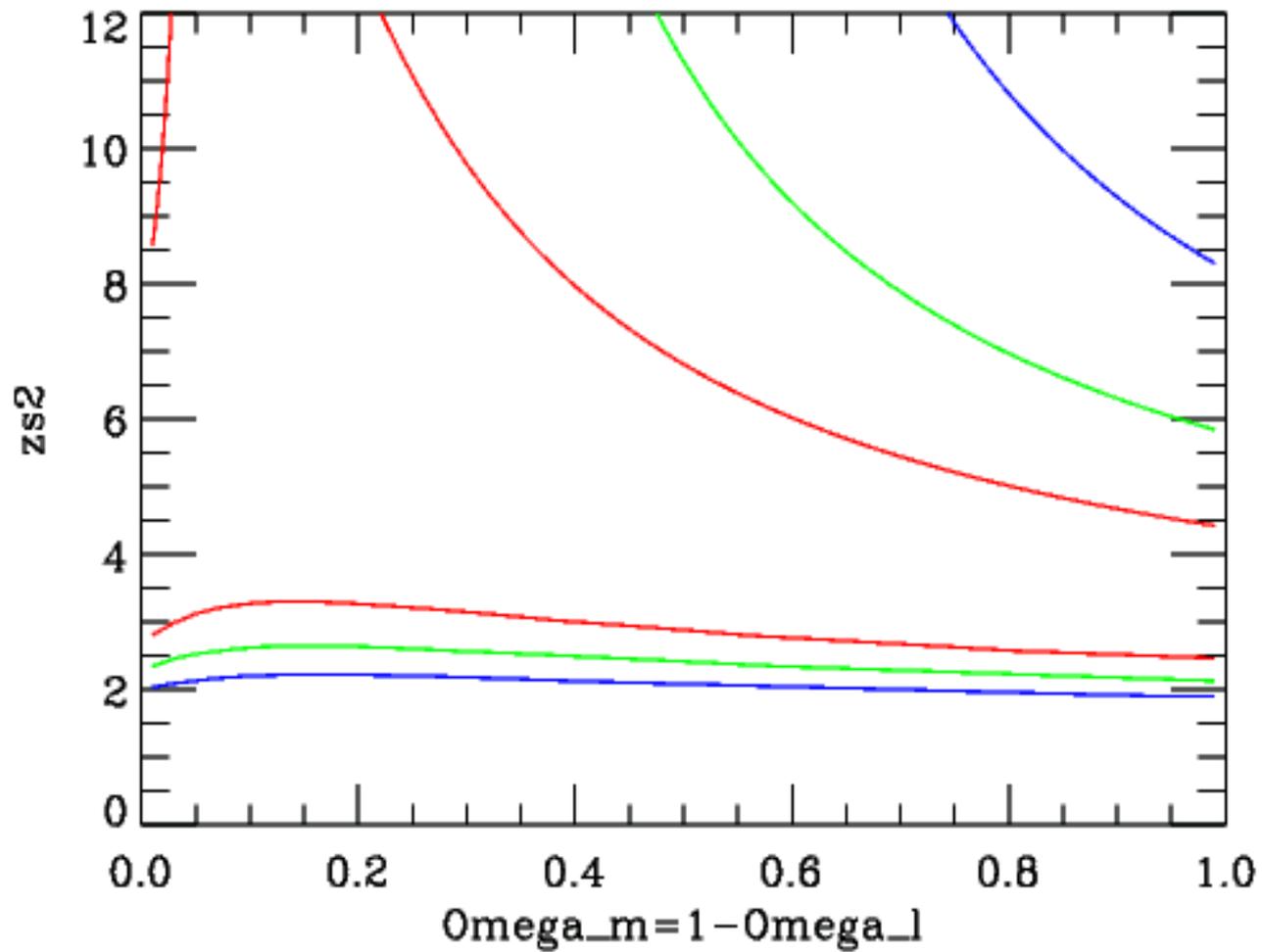
For SDSSJ0946+1006:



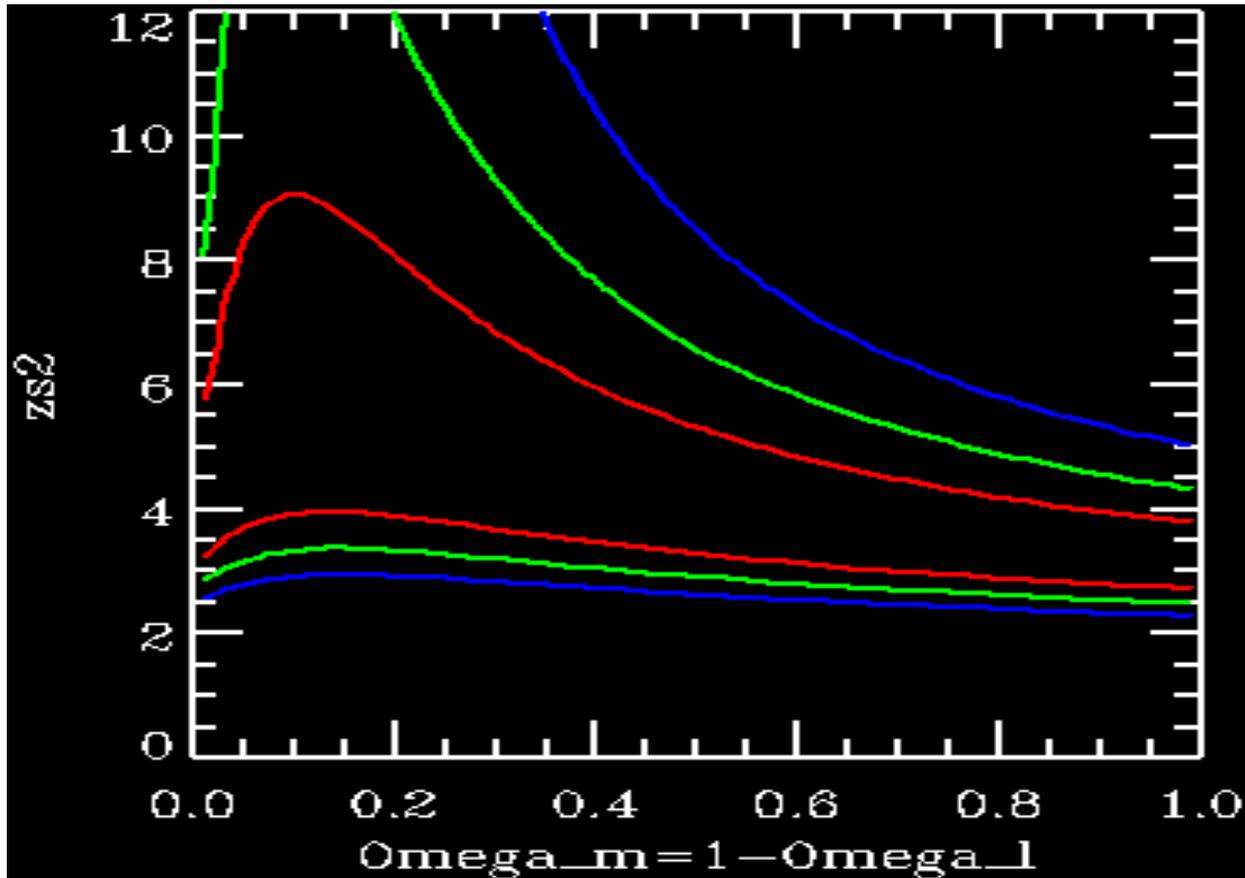
we can measure $\alpha(z_{s1})/\alpha(z_{s2})$: Ratio of distance ratios : a function of cosmological parameters?

Lens model (after marginalisation of ALL unknown params.) gives a PDF for η =ratio of observed Einstein Radii = $D_{ls}/D_s(z_2) / D_{ls}/D_s(z_1)$





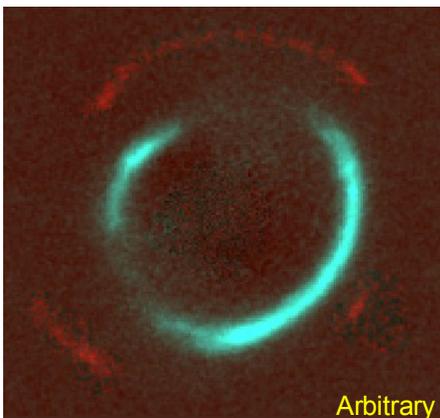
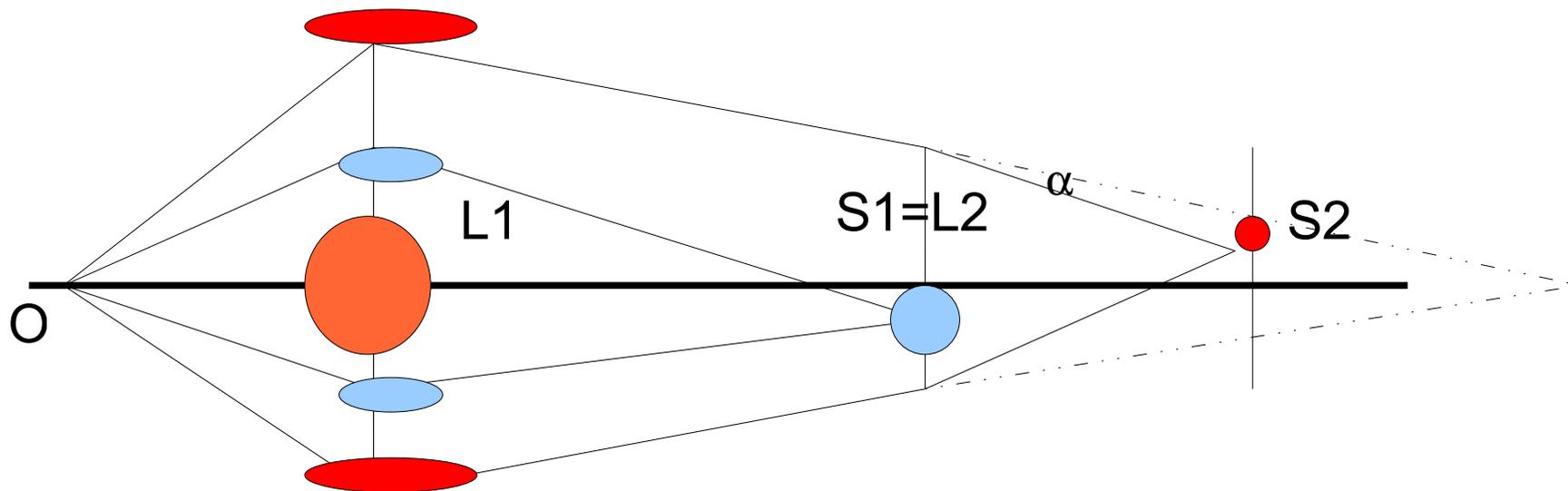
So much for "precision cosmology"...



May be improved a bit with more realistic errors... but...

An ideal optical bench?

Consider the mass contained in source1 that may add extra focusing to source2. That's even more complicated!

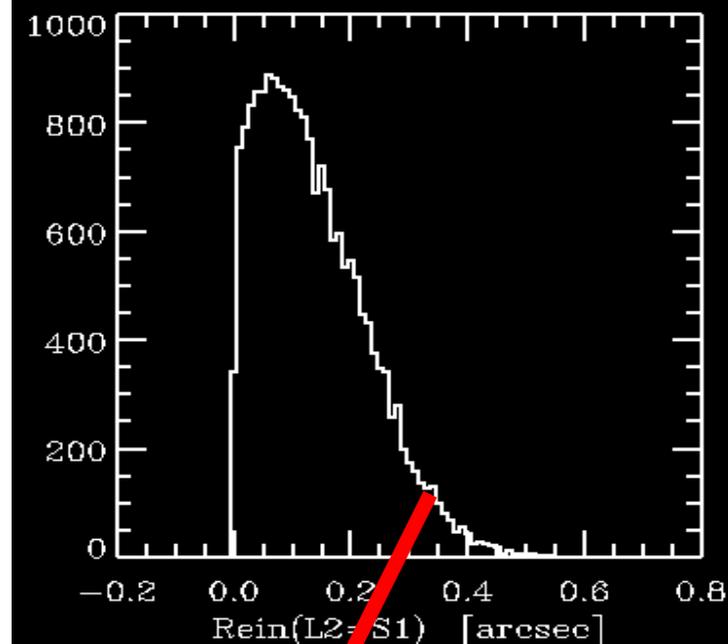


Arbitrary colors

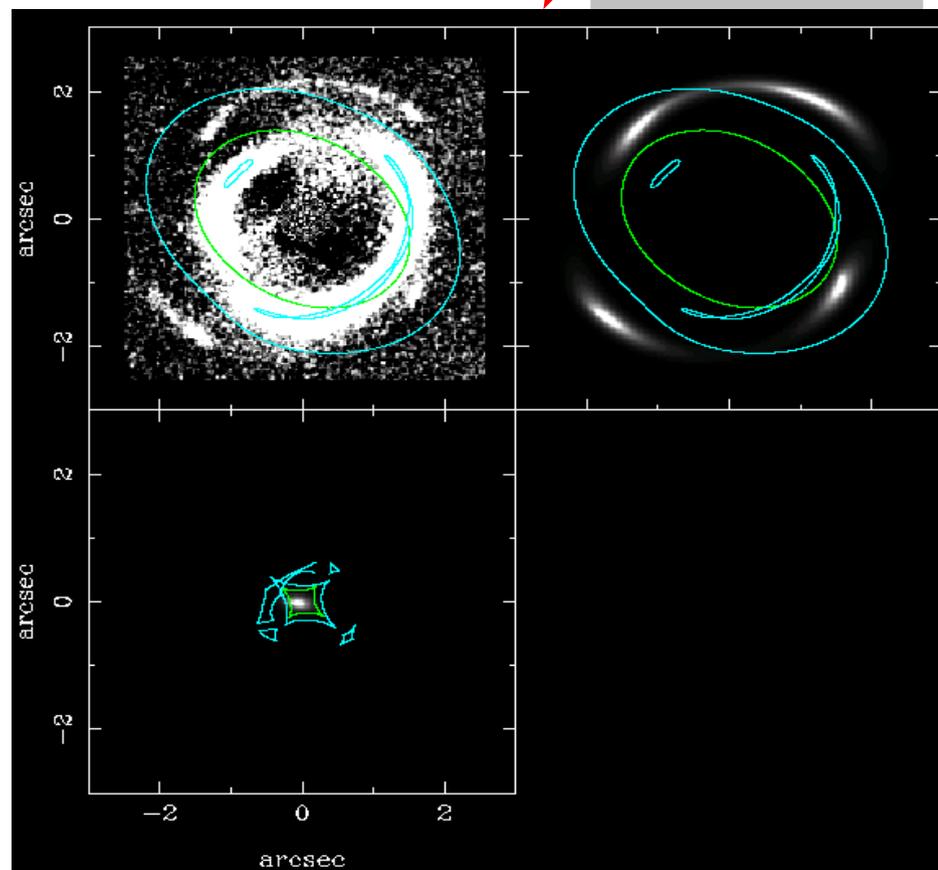
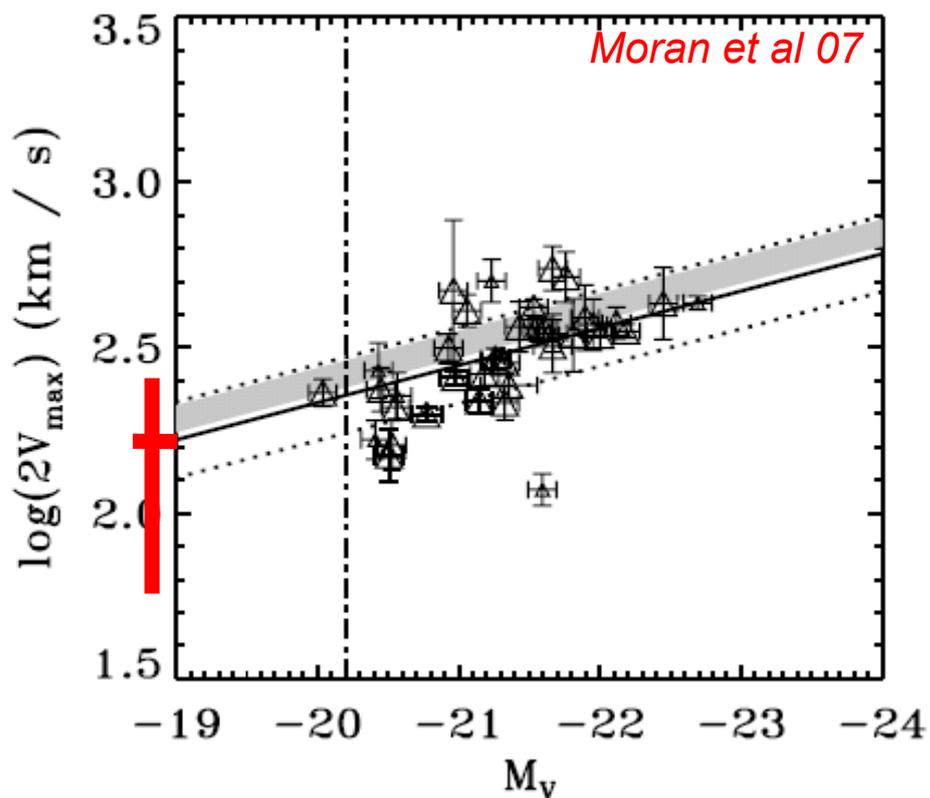
But everybody's noise is somebody's signal...

Positive mass in L2 SIS marginally favored, consistent with rotation velocity $V=90^{+30}_{-60}$ km/s.

Given unlensed luminosity $M_V=-19.0$, falls well onto $z\sim 0.5$ Tully-Fisher relation. (FBCNELLGs #2)



Extreme case



Conclusions

- SDSSJ0946+1006 is a great surprise from SLACS
- Statistics of such events: at least 1 in the sky! We may learn about faint distant sources.
- Mass profile very well probed (normalization and slope) even without z_{s2} and **again close to isothermal and need for DM.**
- **Cosmology poorly constrained with such an unusual optical bench!**
- **Source1 should also be seen as Lens2 for which we get a sensible V_{rot} .**

