

Stellar Mass Density at $z \sim 7$,
SFR density at $z > 7$,
Use of Lensing to Study High-
Redshift Galaxies

Rychard Bouwens

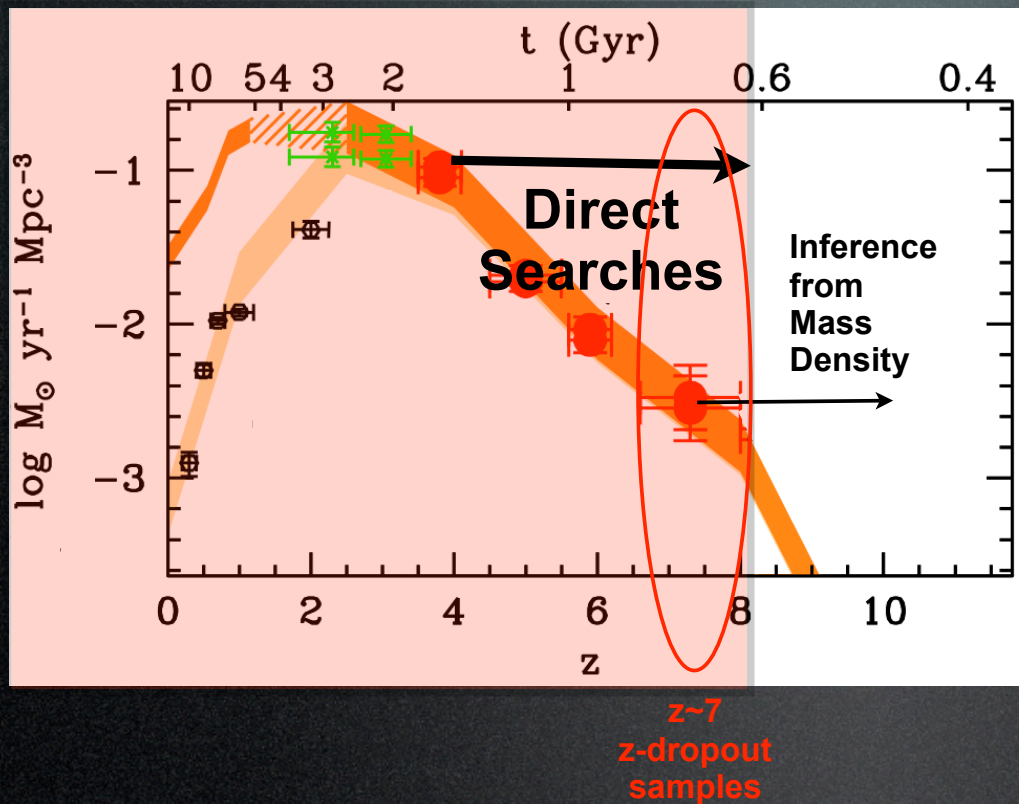
Garth Illingworth, Valentino Gonzalez, Ivo Labbe

Marijn Franx, Larry Bradley, Wei Zheng, Holland Ford

“Lyman-Alpha Universe” -- IAP, Paris

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SFR density at $z > 7$



Challenge:

Difficult at present to determine SFR density at $z > 7$ from direct searches

Solution:

Infer SFR density at $z > 7$ from stellar mass density estimates at $z \sim 7$

Bouwens et al. 2008, 2009; Magnelli et al. 2009; Reddy & Steidel 2009; Schiminovich et al. 2005; and much other work...

Stellar Masses of $z\sim 7-8$ Galaxies

Spitzer data are key for stellar mass estimates



In 2006, we estimated the stellar masses of $z \sim 7$ galaxy candidates using the HUDF data

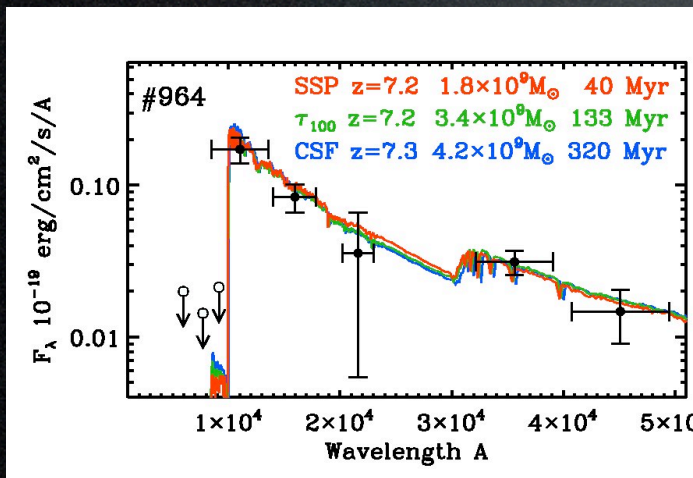
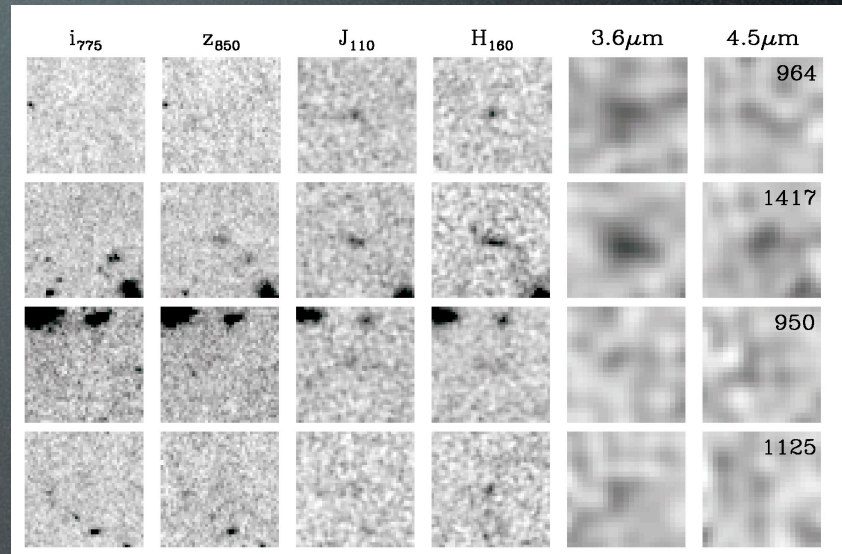
HUDF



z-dropouts

HST/NICMOS
Rest-frame UV

Spitzer/IRAC
Rest-frame Optical



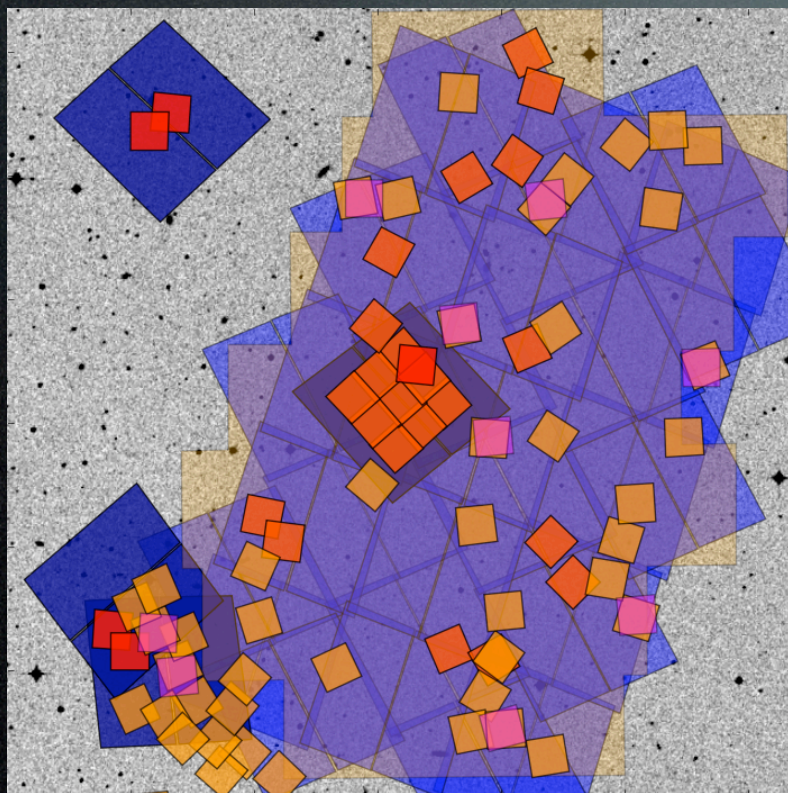
Stellar Masses of $0.3 - 1.0 \times 10^{10} M_{\odot}$

Ages of $\sim 50-200$ Myr

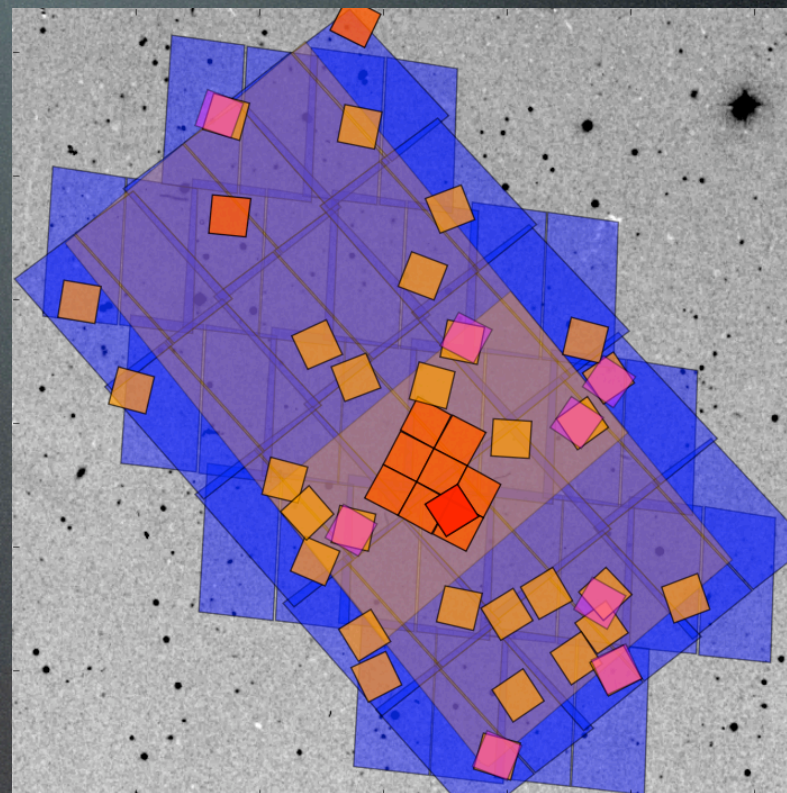
Stellar Mass Density at $z \sim 7.4$ ($> 0.3 L^*$) is
20-60% of $z \sim 5-6$ values

Very wide area, deep near-IR data available now to expand searches for $z \sim 7$ galaxies

$\geq 80 \text{ arcmin}^2$ of Deep ($J_{110} \sim H_{160} \geq 26.5 \text{ AB mag}$) NICMOS coverage



CDF-S GOODS



HDF-N GOODS

15 candidate $z \sim 7$ z-dropouts (12 shown here)

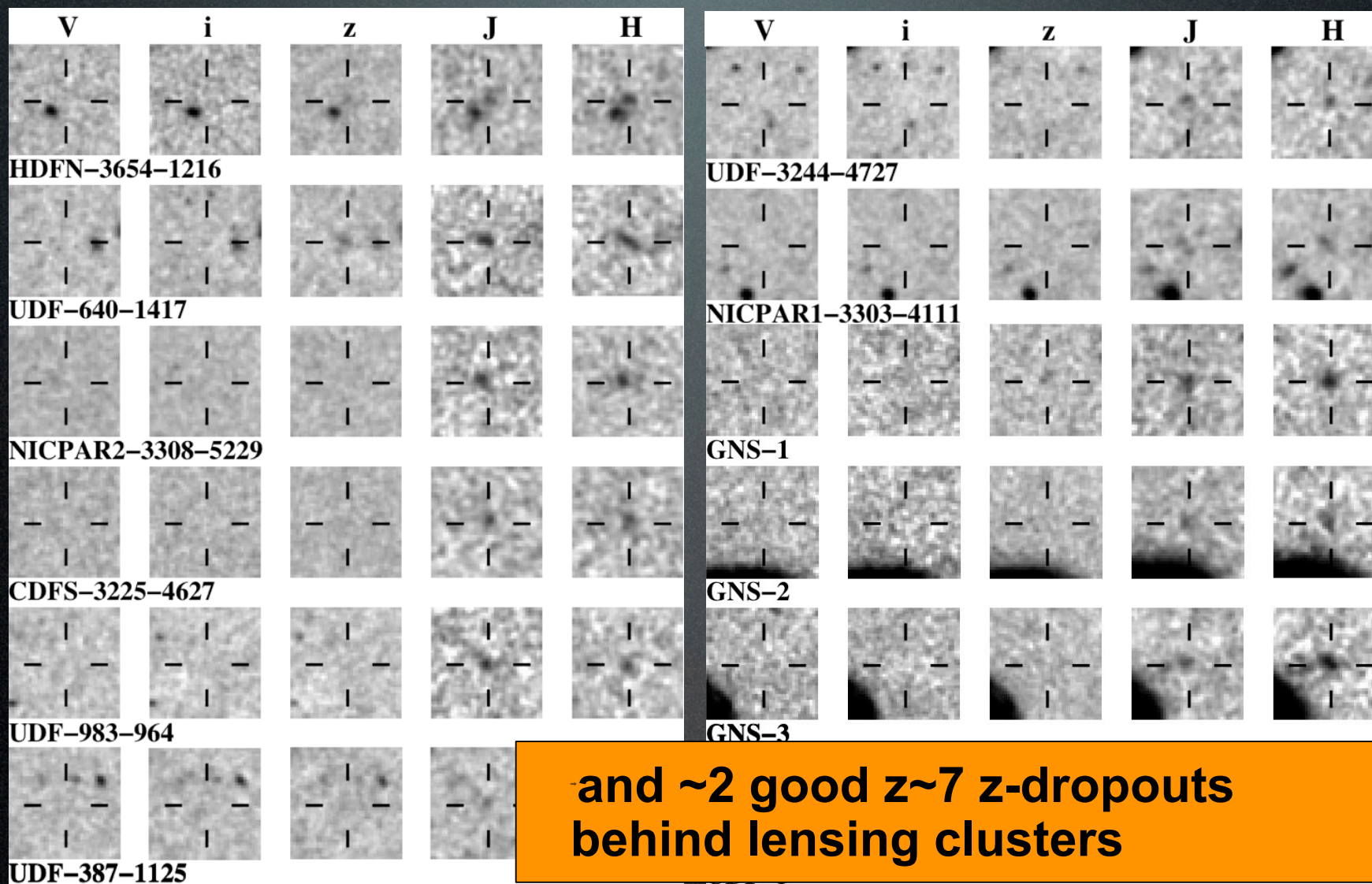
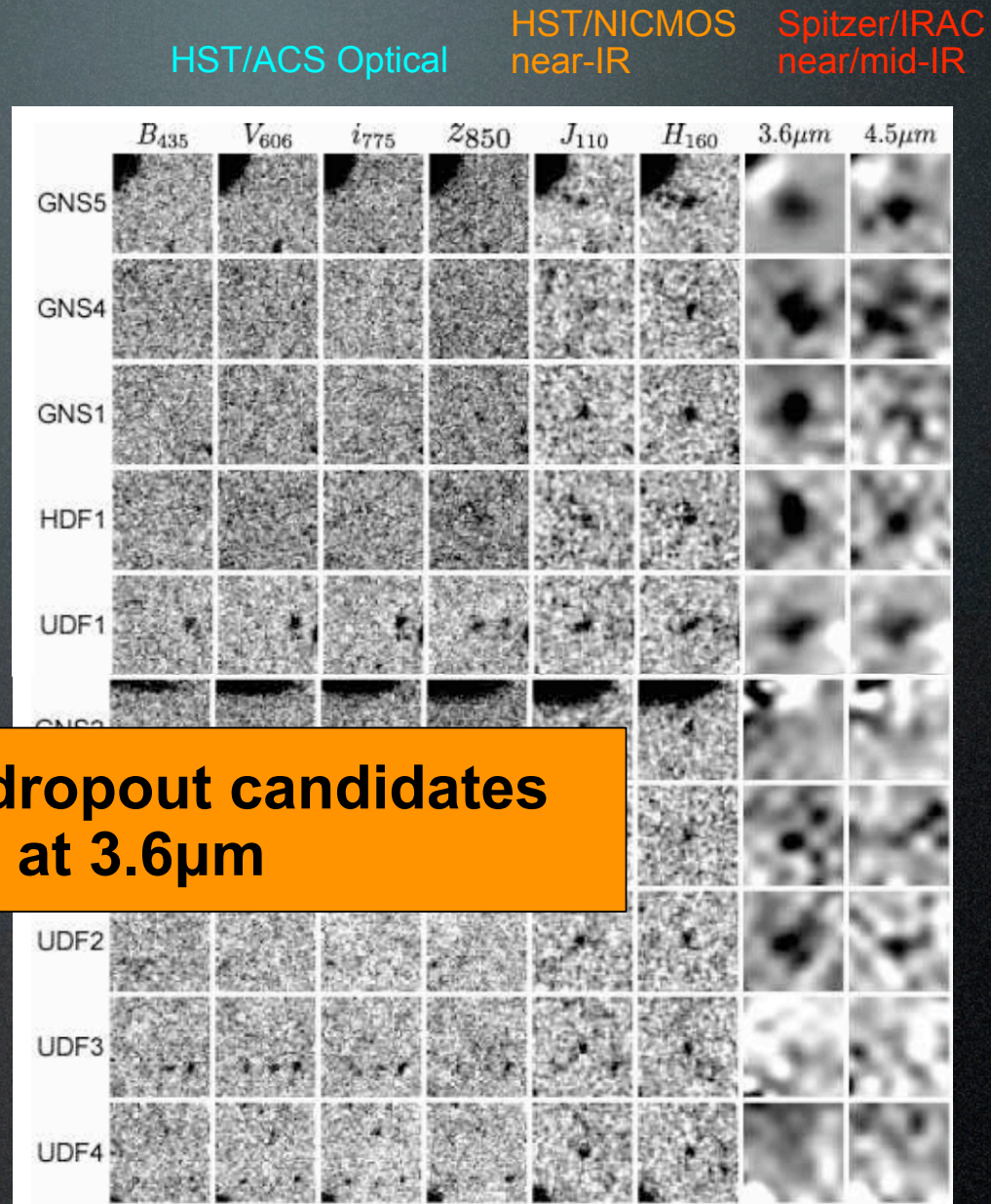
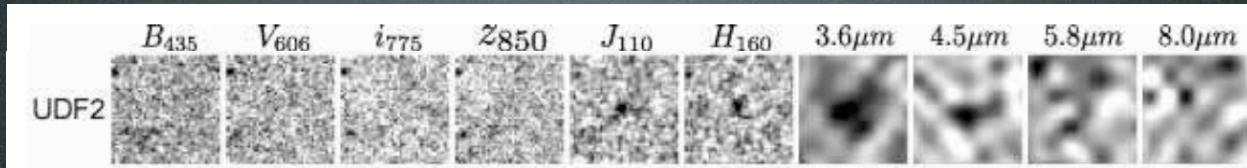


Image cutouts of new sample of z-dropouts

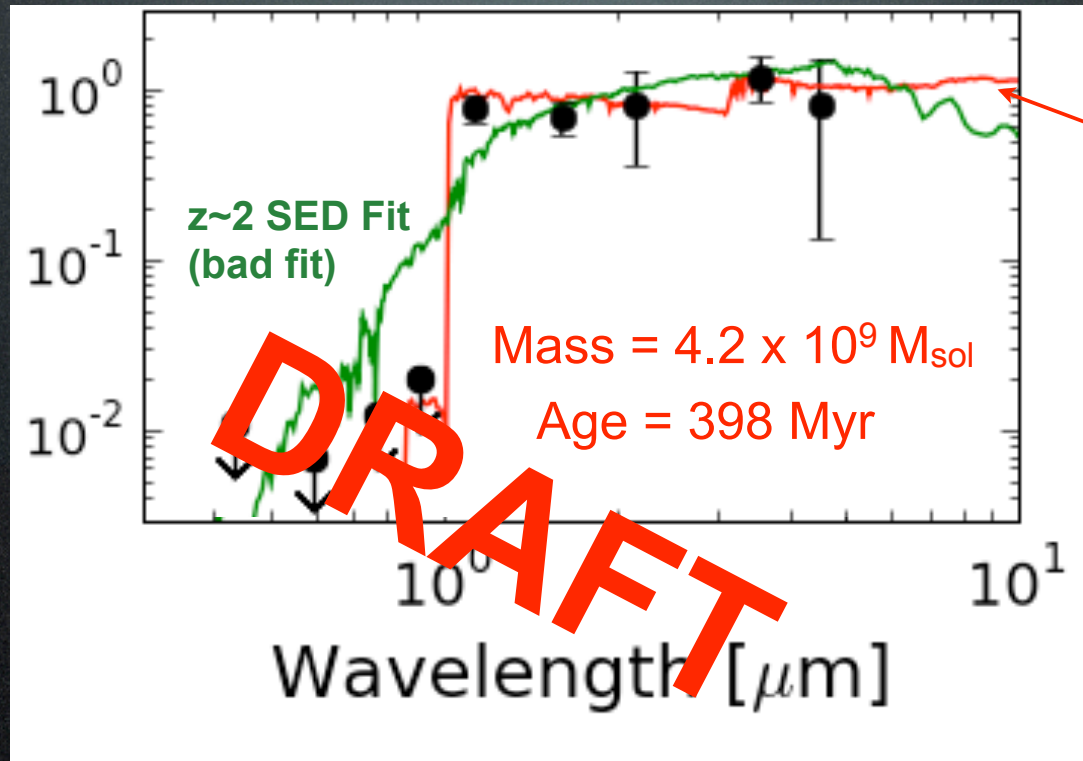


> 50% of z~7 z-dropout candidates clearly detected at 3.6 μm

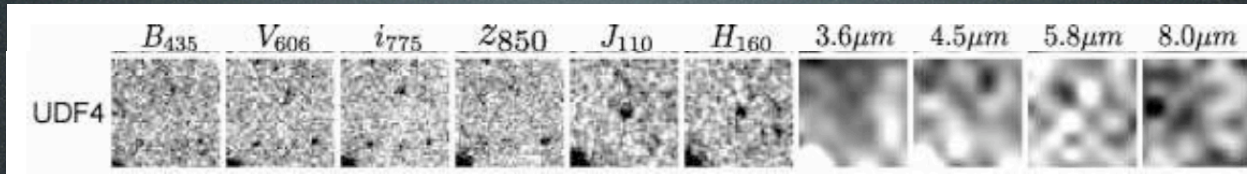
SED Modelling of $z \sim 7$ Galaxy Candidates



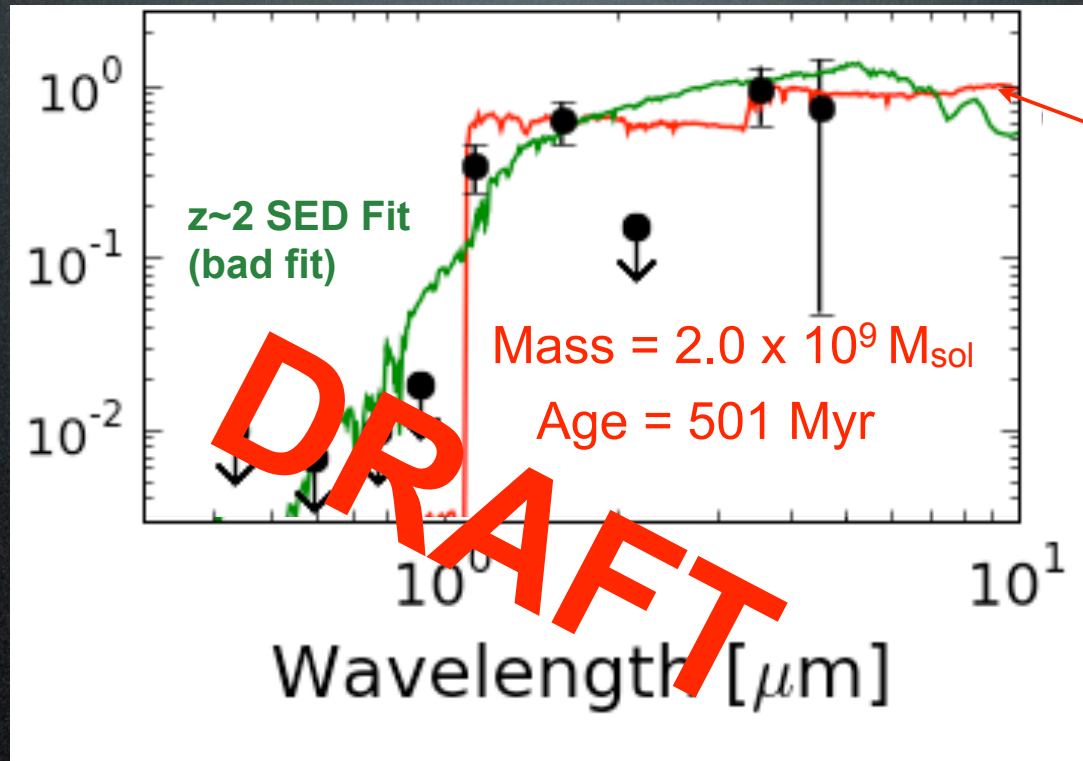
Flux
 $10^{30} \text{ ergs s}^{-1}$
 $\text{cm}^{-1} \text{ Hz}^{-1}$



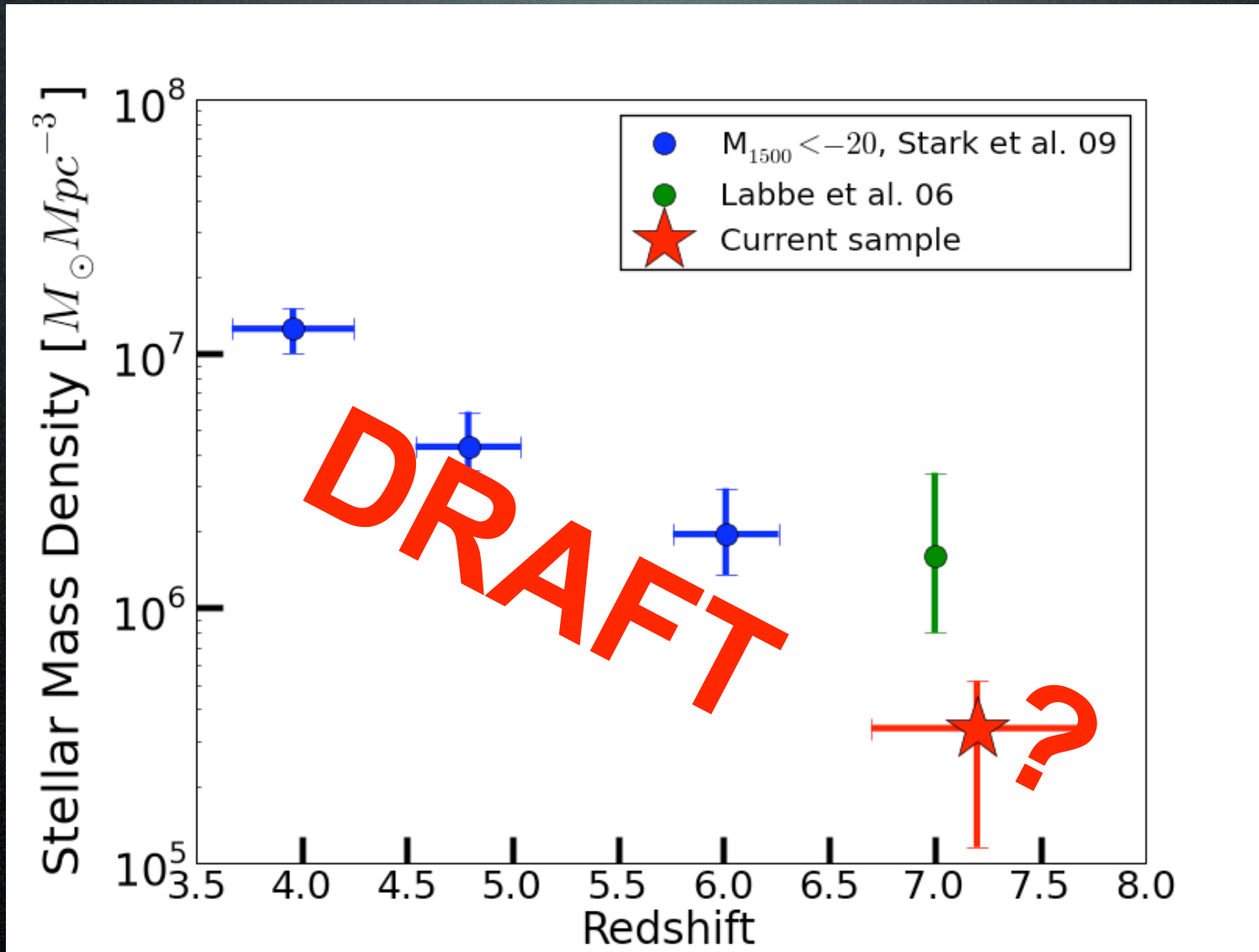
SED Modelling of $z \sim 7$ Galaxy Candidates



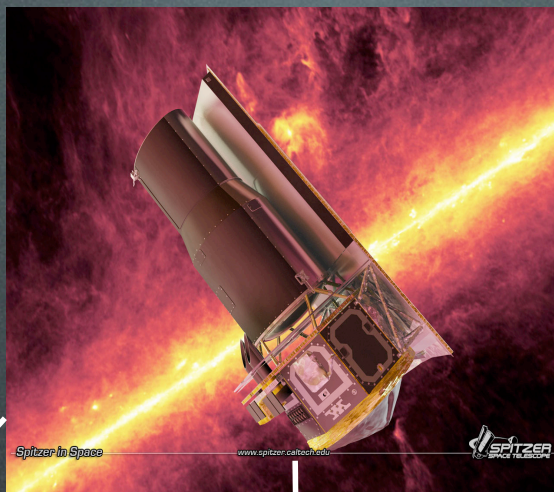
Flux
 $10^{30} \text{ ergs s}^{-1}$
 $\text{cm}^{-1} \text{ Hz}^{-1}$



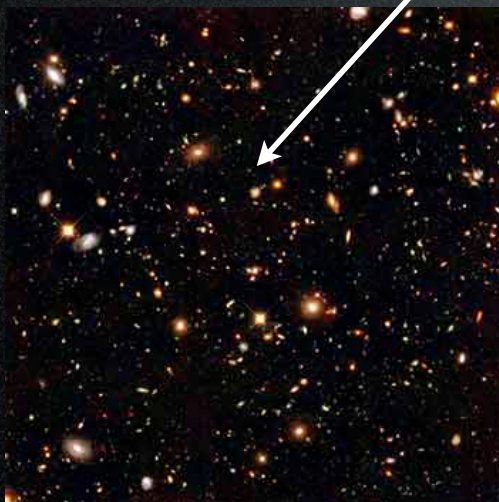
Stellar Mass Density



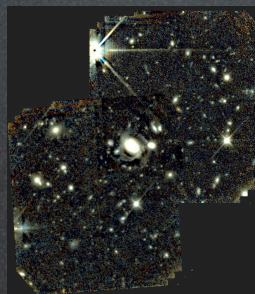
Probing further down the mass function with deeper Spitzer data



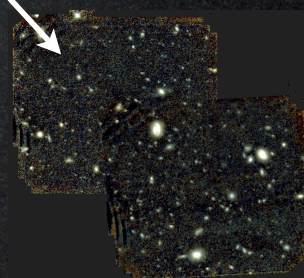
HUDF



HUDF05-1

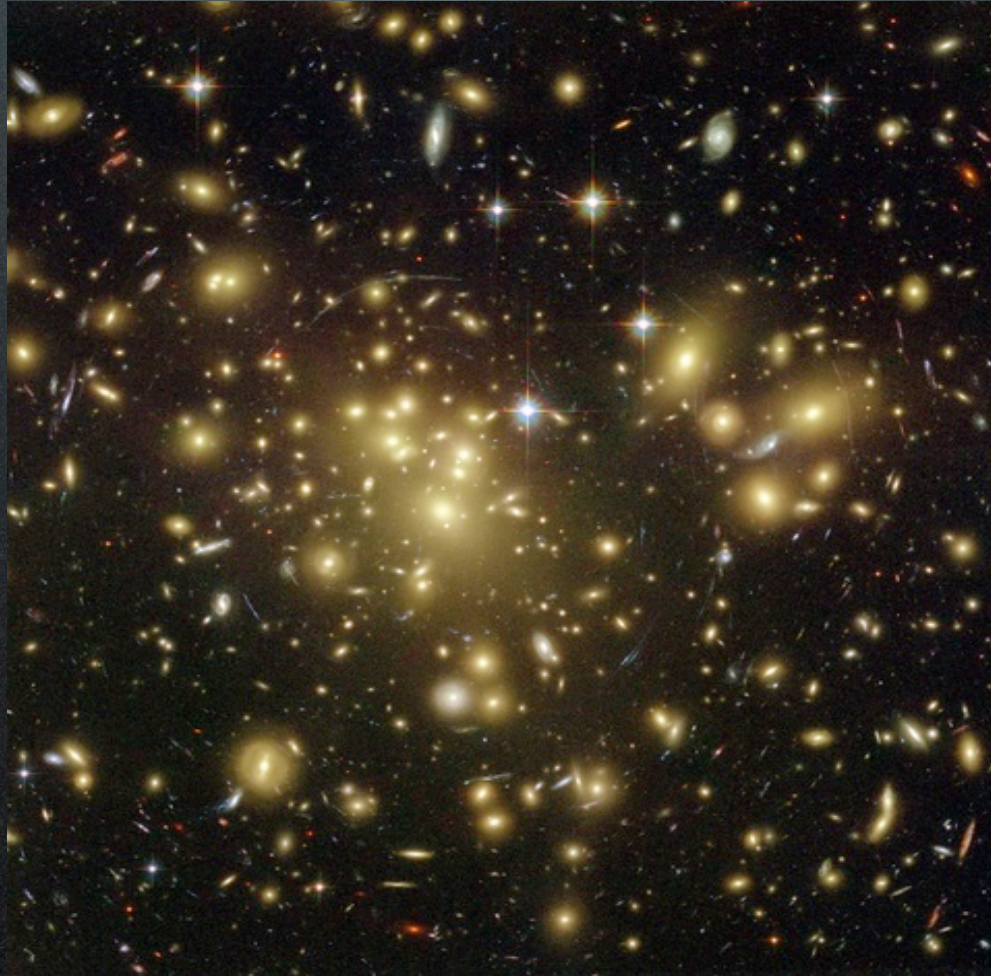


HUDF05-2



$z \geq 4$ galaxies from gravitational lensing

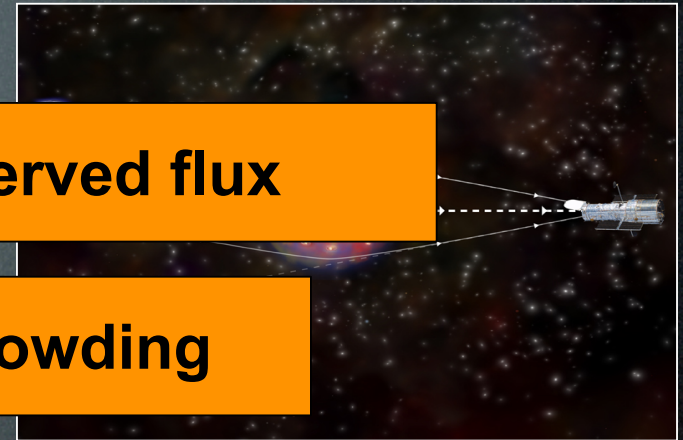
Abell 1689



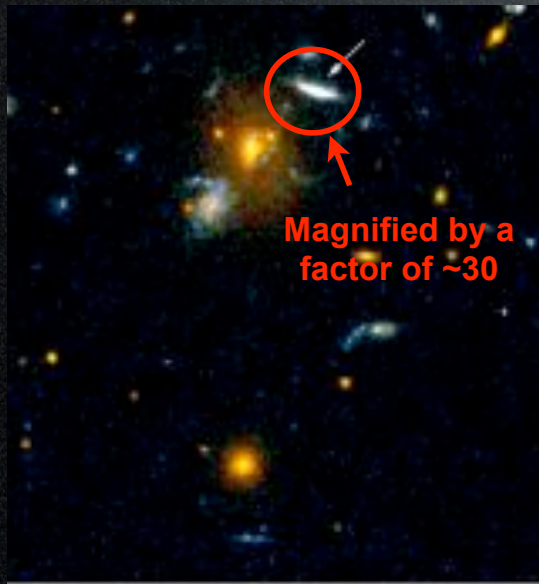
Using gravitational lensing by clusters to look at mass in lower luminosity galaxies

- Significantly enhance observed flux

- Significantly reduce crowding

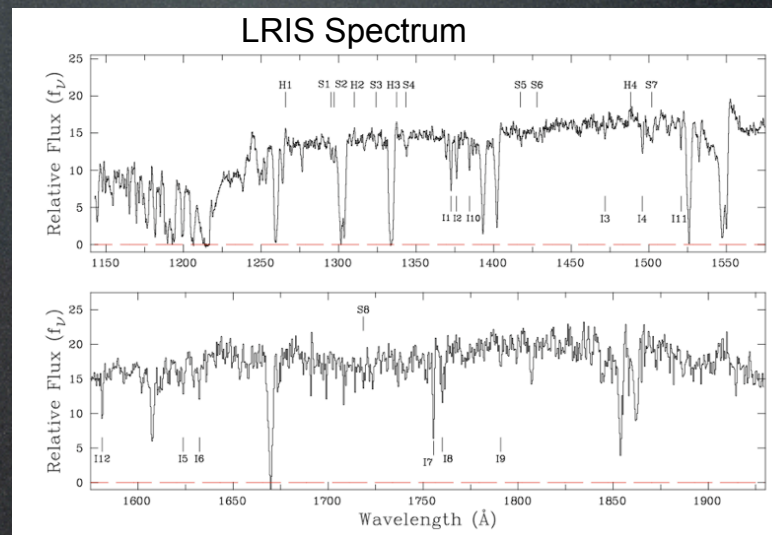


Massive galaxy clusters have certainly proven useful for magnifying the fluxes of faint galaxies that would have too low of fluxes for detailed studies, e.g., cB58 at $z=2.72$



Magnified by a factor of ~ 30

The Distant Galaxy MS 1512-cB58 (HST + WFPC2)

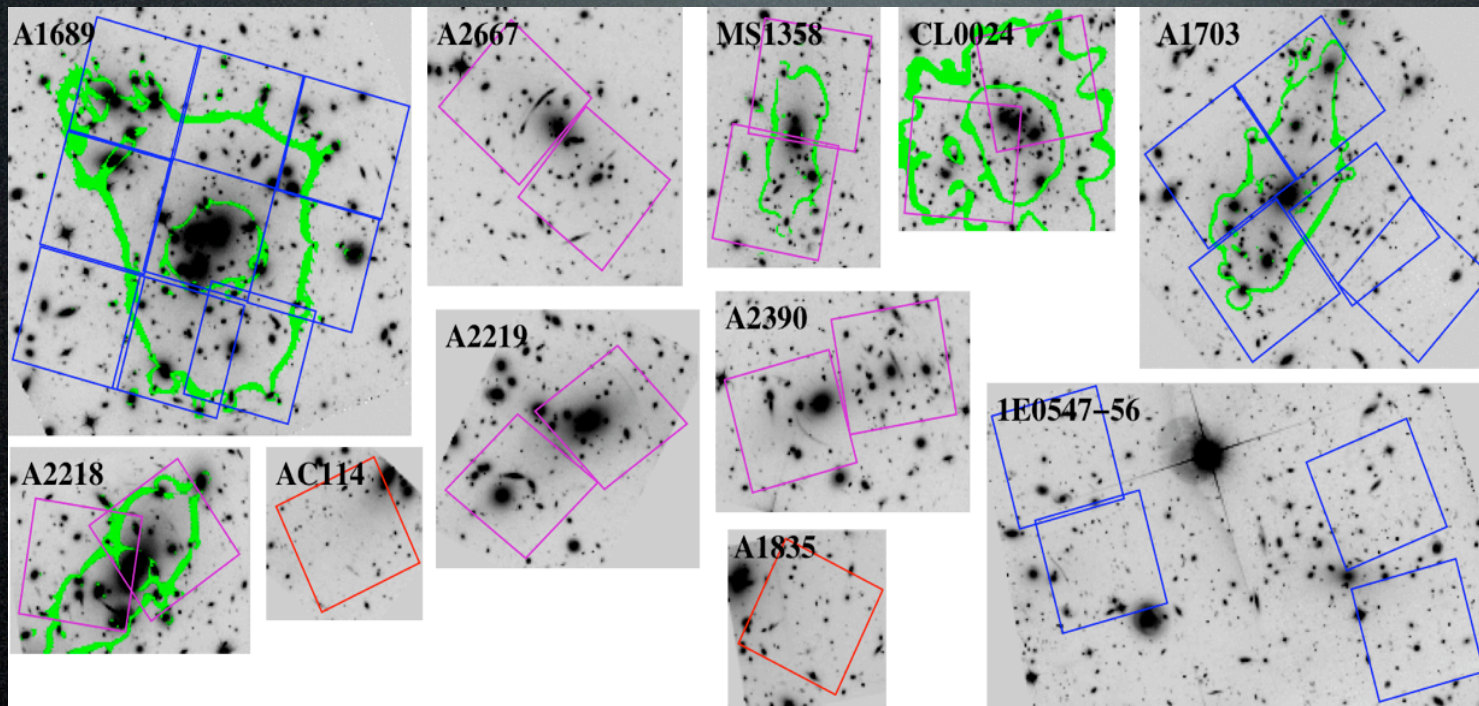


Substantial magnification of the flux from cB58 allows us to study it in great detail

Searching for Lower Luminosity Galaxies that have been highly magnified by massive galaxy clusters

Cluster Search Fields:

-- 23 arcmin² search area (11 clusters)



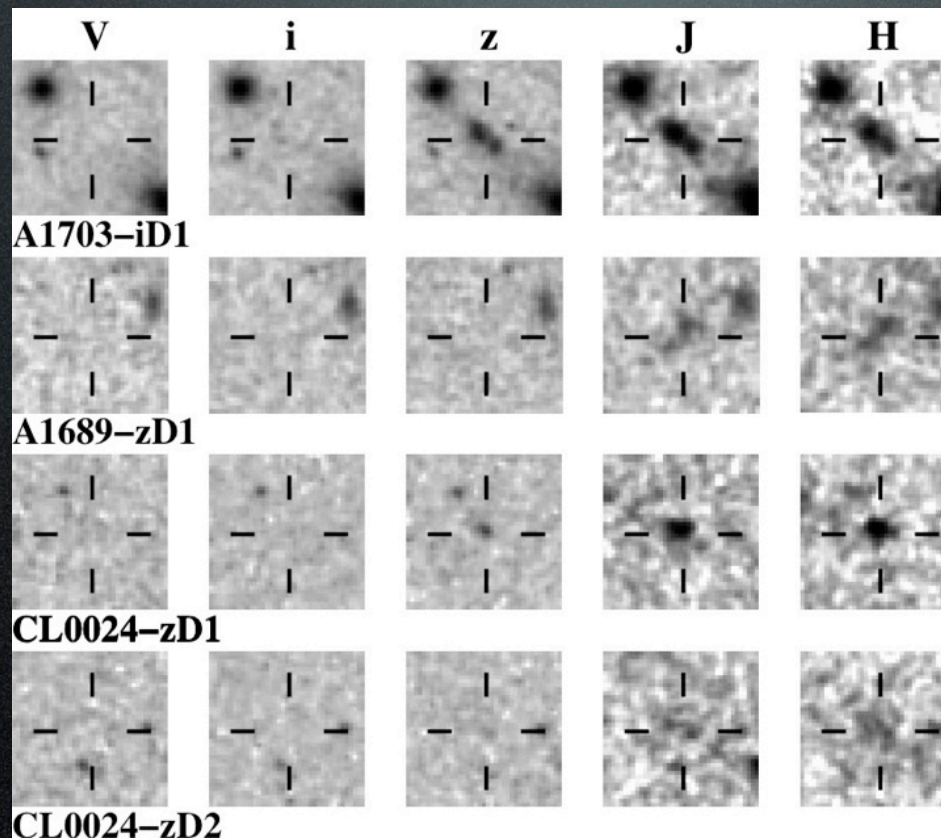
Bouwens et al. 2009; Zheng et al. 2009; see also Richard et al. 2008

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Searching for Lower Luminosity Galaxies that have been highly magnified by massive galaxy clusters

Cluster Search Fields:

-- 4 bright $z\sim 6-7$ candidates found + Kneib et al. 2004 $z\sim 6.5$ candidate



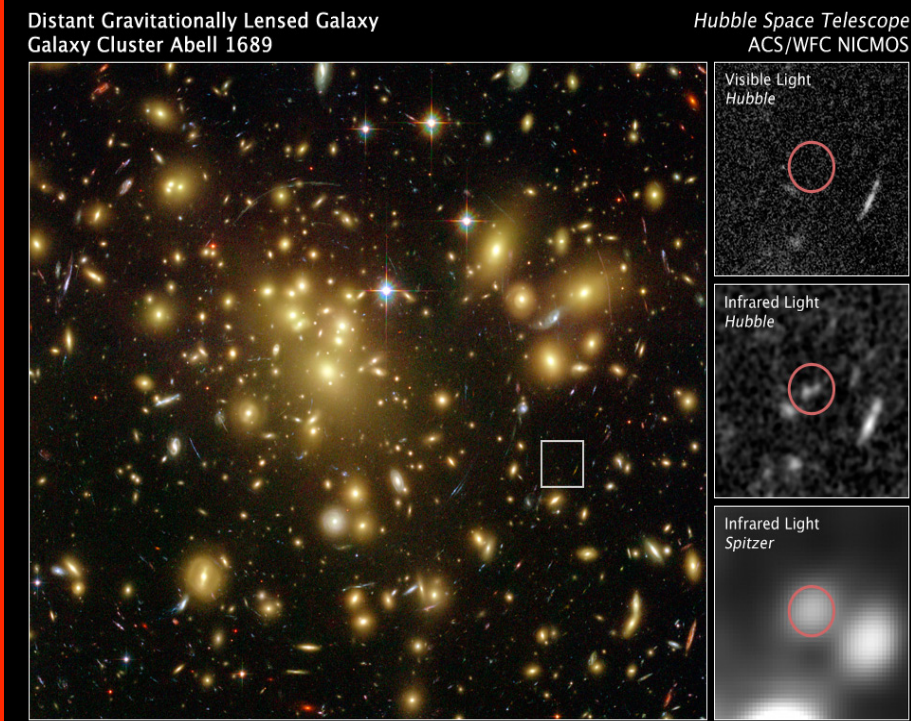
4 other $z\geq 7$ candidates (but which do not have deep enough optical data to be sure)

Bouwens et al. 2009; Zheng et al. 2009; see also Richard et al. 2008

Searching for Lower Luminosity Galaxies that have been highly magnified by massive galaxy clusters

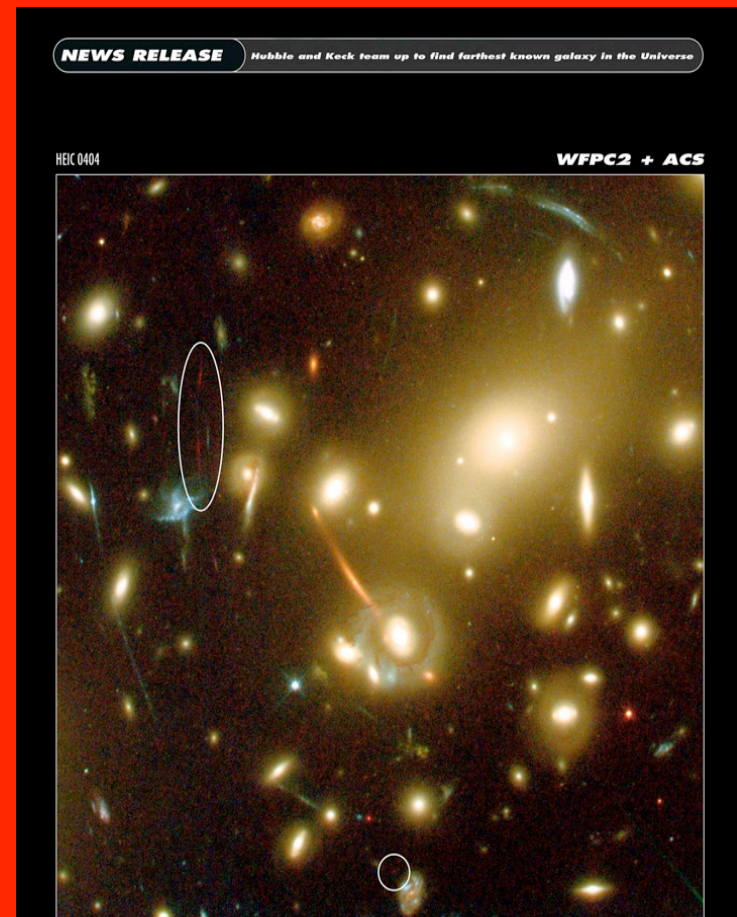
Cluster Search Fields:

-- 4 bright $z\sim 6-7$ candidates found +
 $z\sim 7.6$ candidate (Bradley et al. 2008)



4 other $z\geq 7$ candidates (but which do not have
Bouwens et al. 2009; Zheng et al. 2009

$z\sim 6.5$ candidate (Kneib et al. 2004)



HUBBLE SPACE TELESCOPE

ESA, NASA, J.-P. Kneib (Observatoire Midi-Pyrénées) and R. Ellis (Caltech)

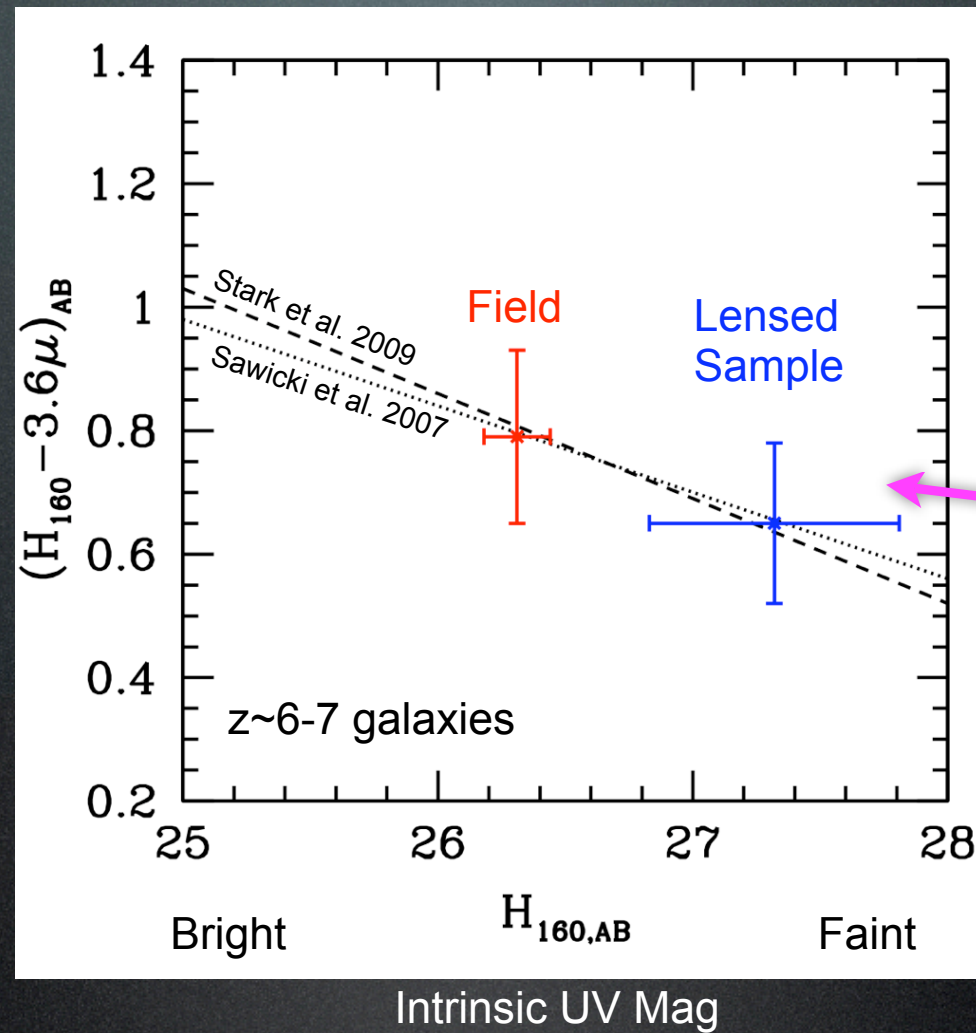


How do lower luminosity lensed galaxies differ from brighter field galaxies in terms of their Balmer Breaks?

Old
High
M/L

Balmer
Break

Young
Low
M/L



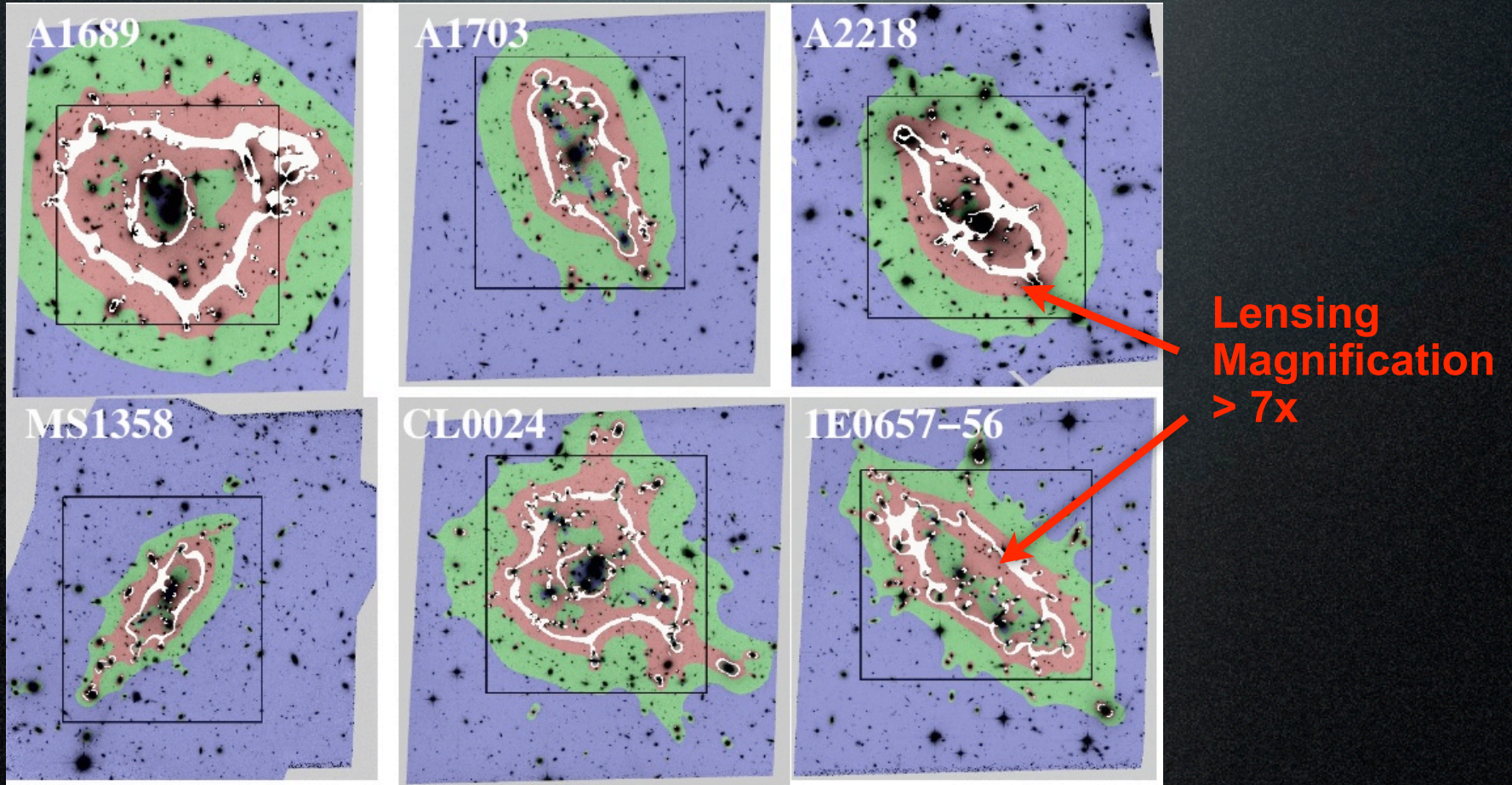
Fainter
Galaxies
Have Lower
M/L ratios

Not only search for
lensed galaxies at $z \sim 6-8$

But also

Search for lensed galaxies at
 $z \sim 4-6$ behind galaxy clusters...

Six clusters with deep HST optical data + lensing models
(reaching 27.5 AB mag)



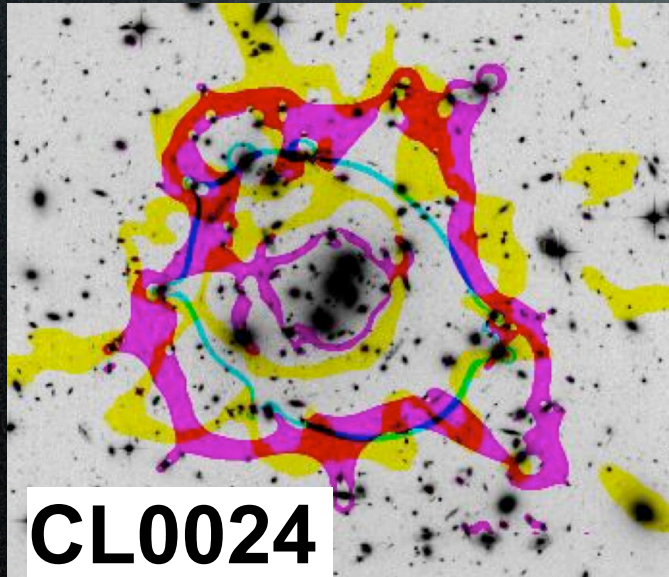
Models: Limousin et al. 2007, 2009; Eliasdottir et al. 2008; Franx et al. 1997; Zitrin et al. 2009; Bradac et al. 2007

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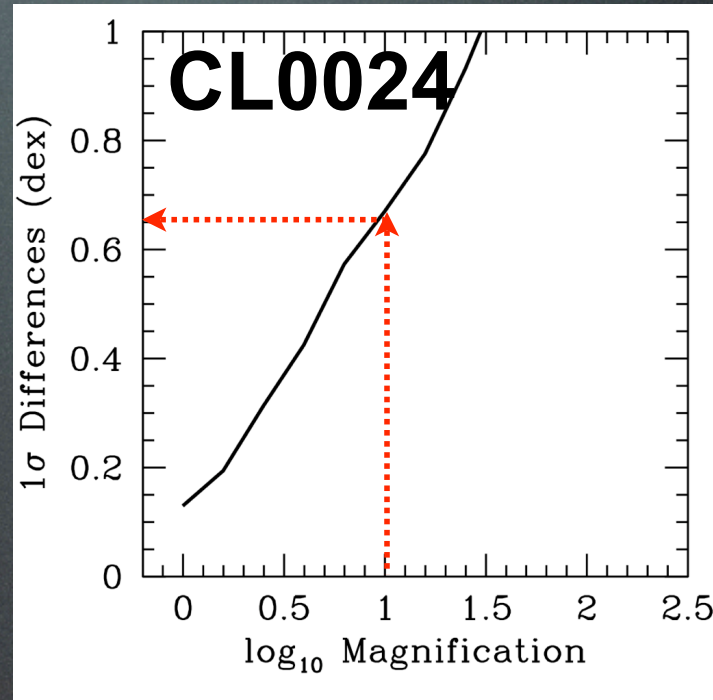
How well can we use lensing
by clusters to find highly
magnified ($> 7x$) galaxies?

Note that we consider all galaxies lensed by $>$ factor of 7 together to be conservative

Model-to-model Dispersion
in magnification maps

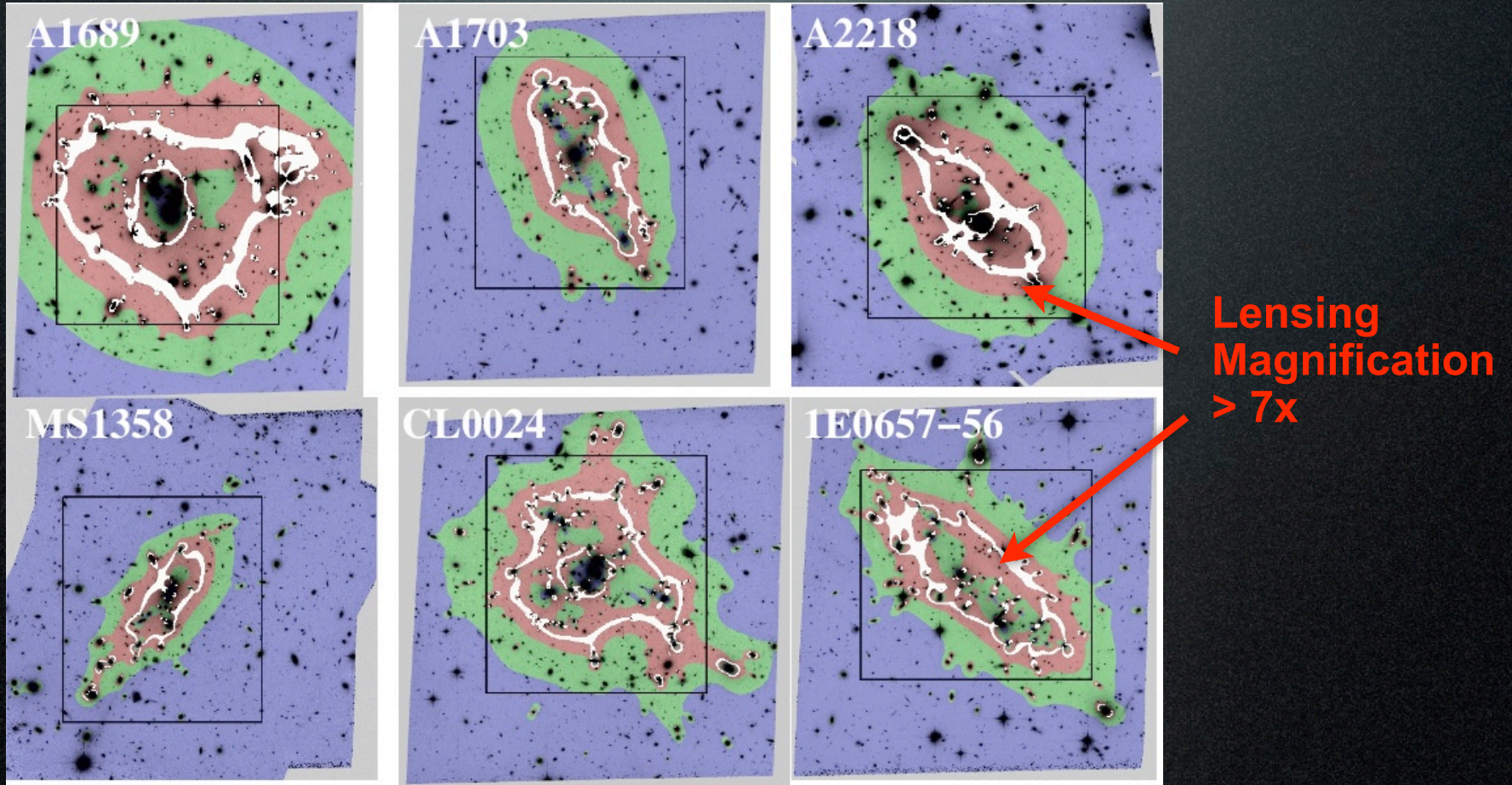


few sets of arcs
weak lensing constraints



Factor of ~4
uncertainties
at magnification
factors of ~10....

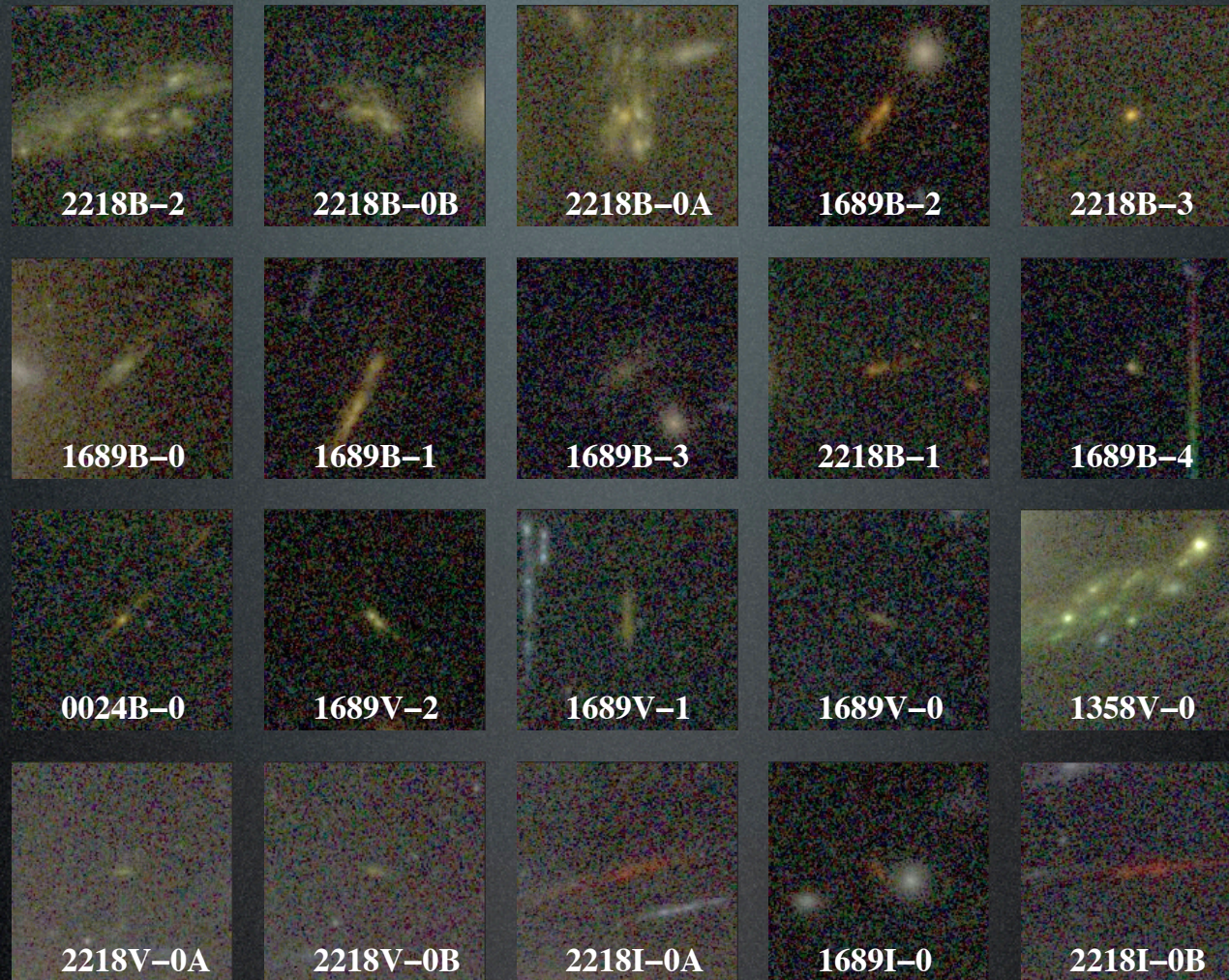
Six clusters with deep HST optical data + lensing models
(reaching 27.5 AB mag)



Models: Limousin et al. 2007, 2009; Eliasdottir et al. 2008; Franx et al. 1997; Zitrin et al. 2009; Bradac et al. 2007

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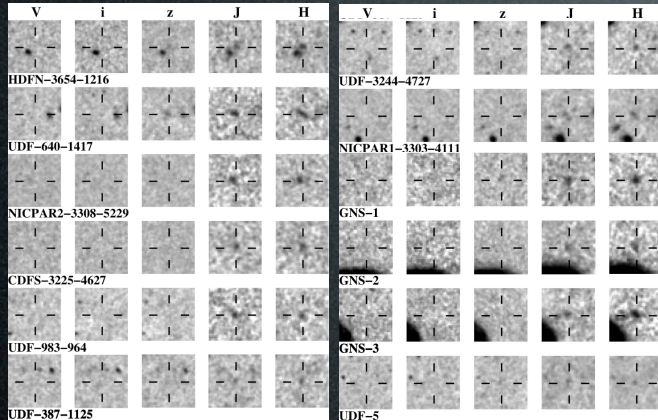
Sample of 20 Highly Magnified ($>7\times$ magnification) $z\sim 4-6$ Galaxies:



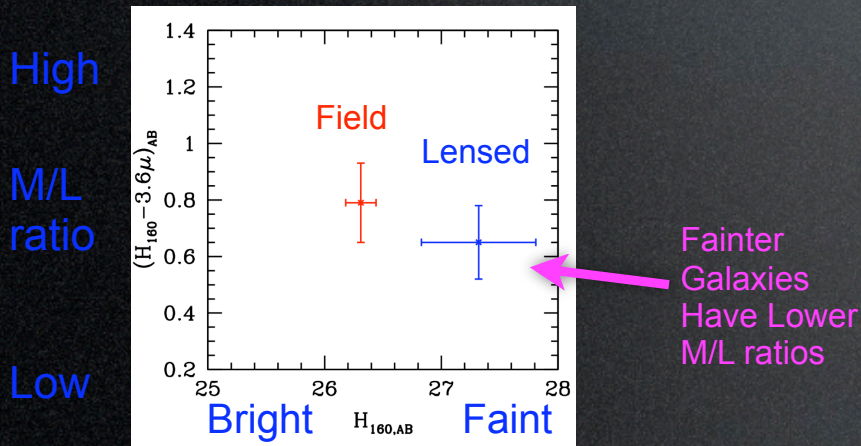
11 $z\sim 4$ galaxies,
6 $z\sim 5$ galaxies,
3 $z\sim 6$ galaxies

Summary

With current (~15 object) z-dropout selections, we can begin deriving much more robust estimates of stellar mass density at $z \sim 7$

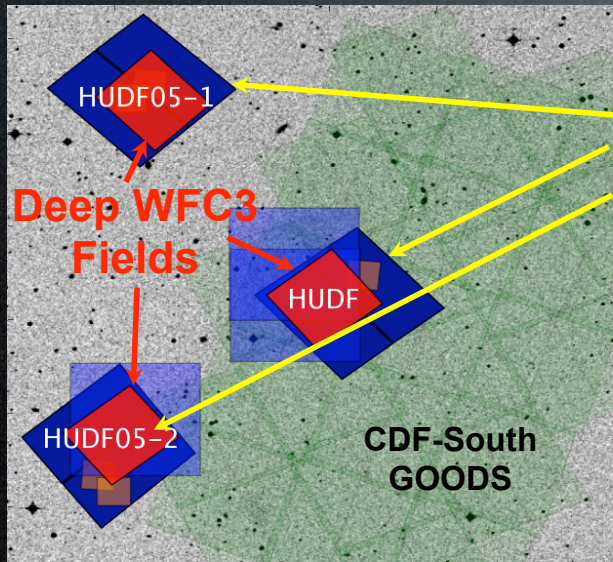


Gravitational Lensing by galaxy clusters is extremely useful for detailed studies of very faint $z \sim 4-8$ galaxies, but may be challenging for conducting LF studies



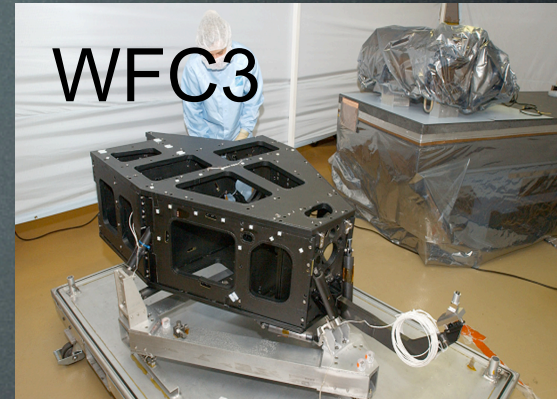
The Future

HUDF09 WFC3/IR program



192 WFC3/IR orbits:
96 orbits / 1 field
48 orbits / 2 fields

Will reach ~29 AB mag in near-IR (1.05, 1.25, 1.60 microns)
Should find 50-100 $z \geq 7$ galaxies



**Deepest
optical data**

