

Hydrogen in SNe Ib

Eddie Baron

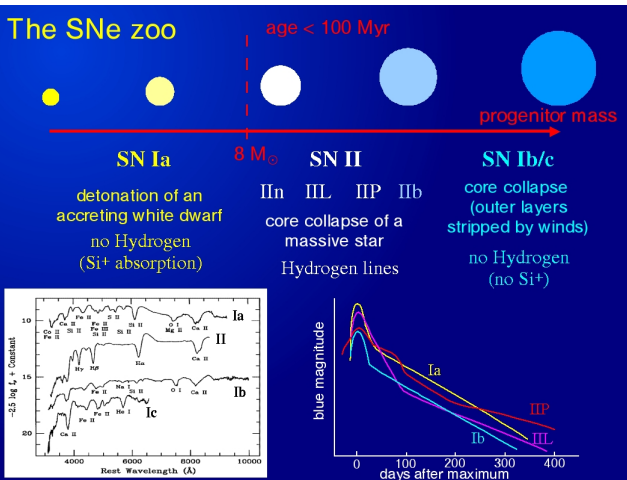
University of Oklahoma, USA



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- Usual Suspects
 - David Branch
 - Peter Hauschildt
- Good Undergraduates
 - Wesley Ketchum
 - Spencer James

SN Taxonomy

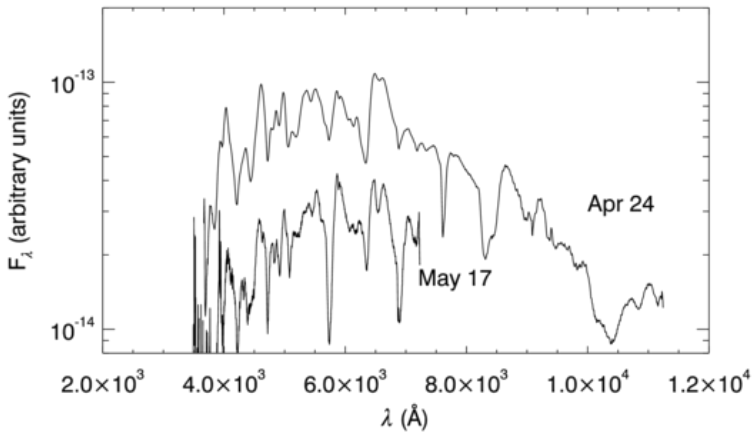


3 Classes of Ixb SNe

- Most if not all in Binary Systems
- SNe IIb $M_H \sim 0.3 M_{\odot}$
 - Progenitor: Stable Case C, Critical mass $\sim 0.3 M_{\odot}$ Roche lobe shrinks
- Hydrogen Rich Ib $M_H \sim 0.1 M_{\odot}$
 - Progenitor: Stable Case C, followed by wind loss? (Posiadlowski et al.)
- Hydrogen Poor Ib $M_H \sim 0.001 M_{\odot}$
 - Progenitor: UnStable Case C, Common Envelope Formation (Posiadlowski et al.)
 - Progenitor: Rotation + Nonconservative Case A or B, Followed by Wind loss (Yoon et al.)

SN IIb

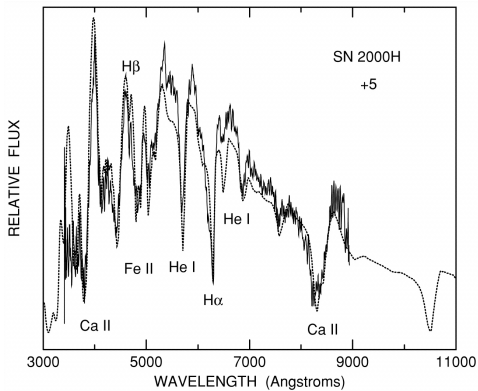
SN 1993J: Days +25, 48



- Highly parametrized
- Good for line IDs
- Can't determine abundances or masses

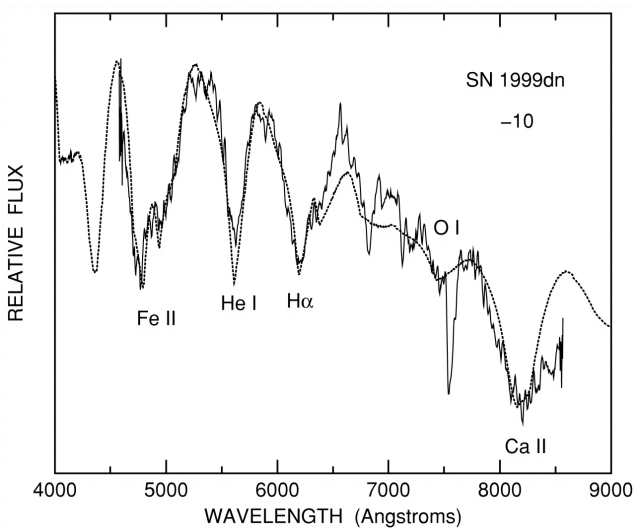
Hydrogen Rich SN Ib

SN 2000H



Hydrogen in Typical SN Ib?

SN 1999dn



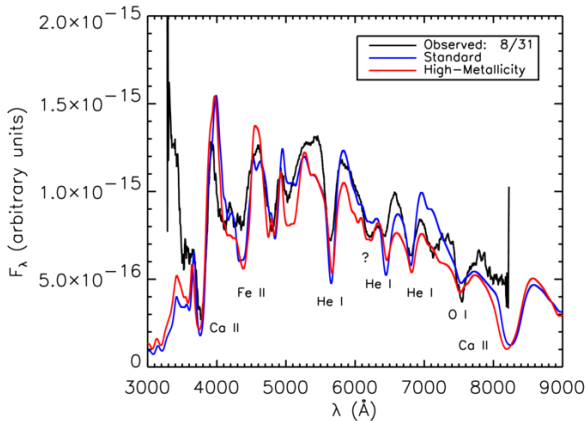
- PHOENIX: Generalized Stellar Atmospheres Code
- Try to use “best physics”

Hydrogen Free Models

- Parametrized Model
- Powerlaw Density
- Helium Core
- Vary Metallicity

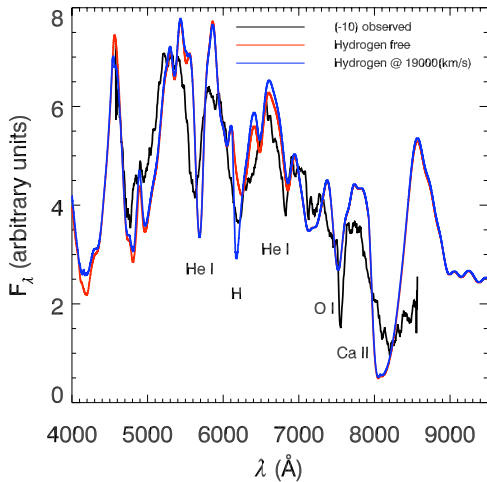
SN 1999dn

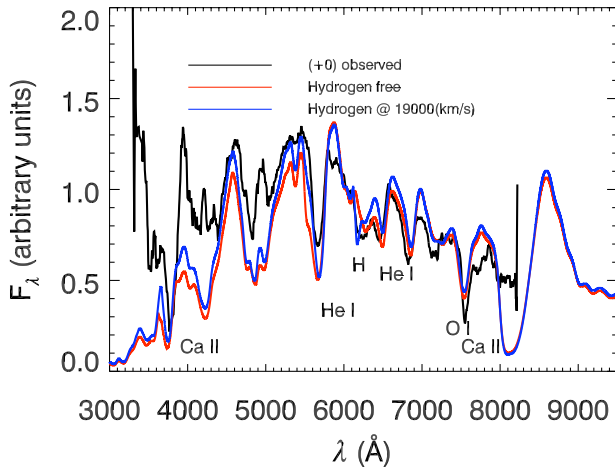
Maximum Light

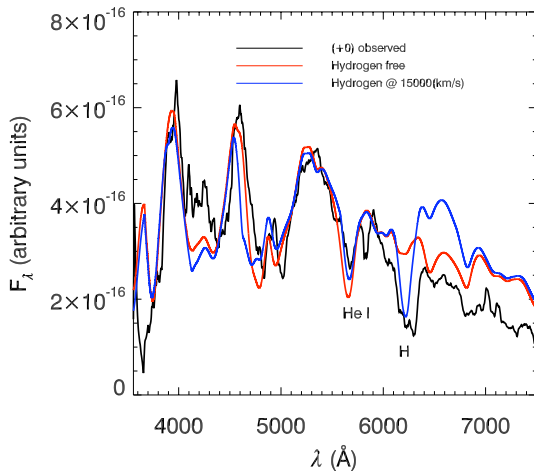


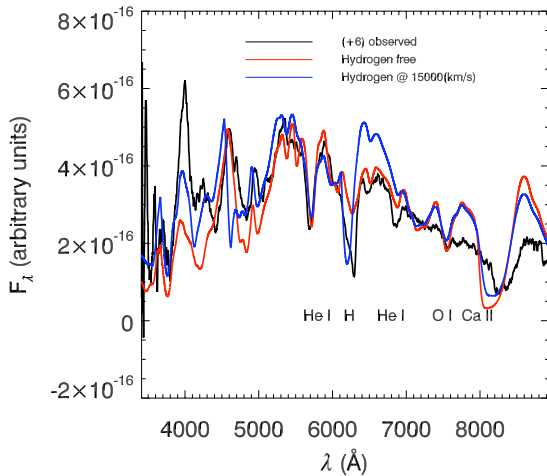
Hydrogen Skin Models

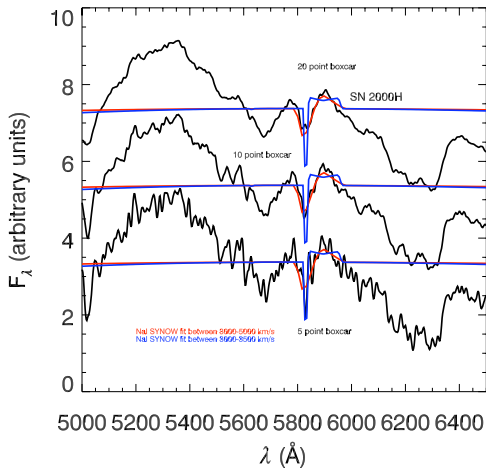
- Parametrized Model
- Solar Compositions
- Powerlaw Density
- Helium Core with hydrogen (solar) skin
- Self-consistent











- Hydrogen is common in SNe Ib
- May be two classes of SNe Ib
 - Class I: Closely Related to SNe IIb
 - Class II: Very thin hydrogen envelope
 - Class III: Hydrogen free?