

Properties of Galaxies hosting Gamma-Ray Bursts

Sandra Savaglio (Max-Planck-Institut für extraterrestrische Physik, Garching)



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XXVI IAP Annual Colloquium 2010, Paris
"Progenitors and environments of stellar explosions"

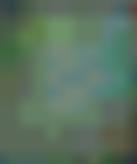
Long-duration GRB: CC SN

Short-duration GRB: NS/BH Merger

GRB 011121

redshift $z = 0.362$

3.98 billion years ago



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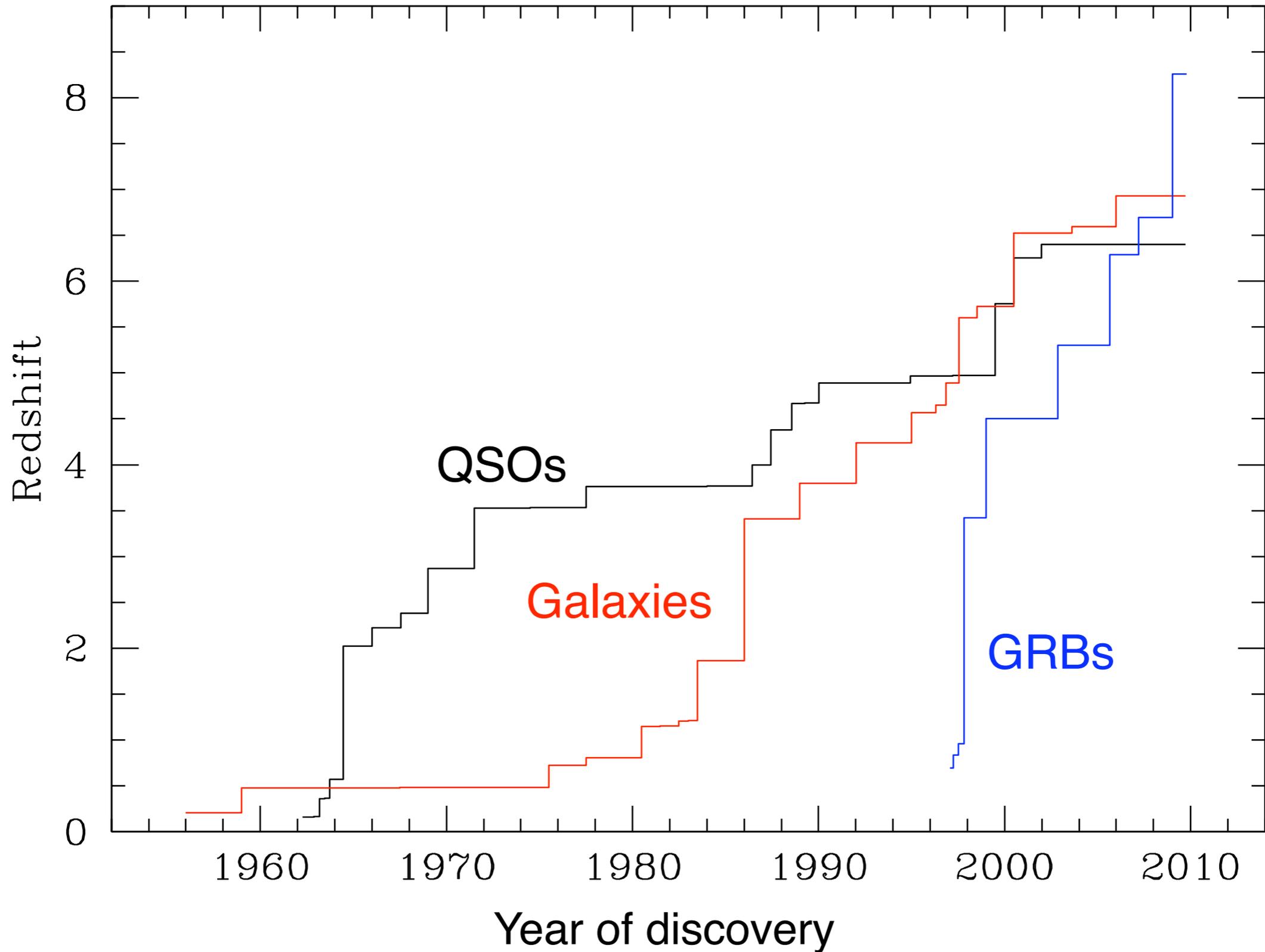
Rates (Long):

1/1000: GRB/CC-SN

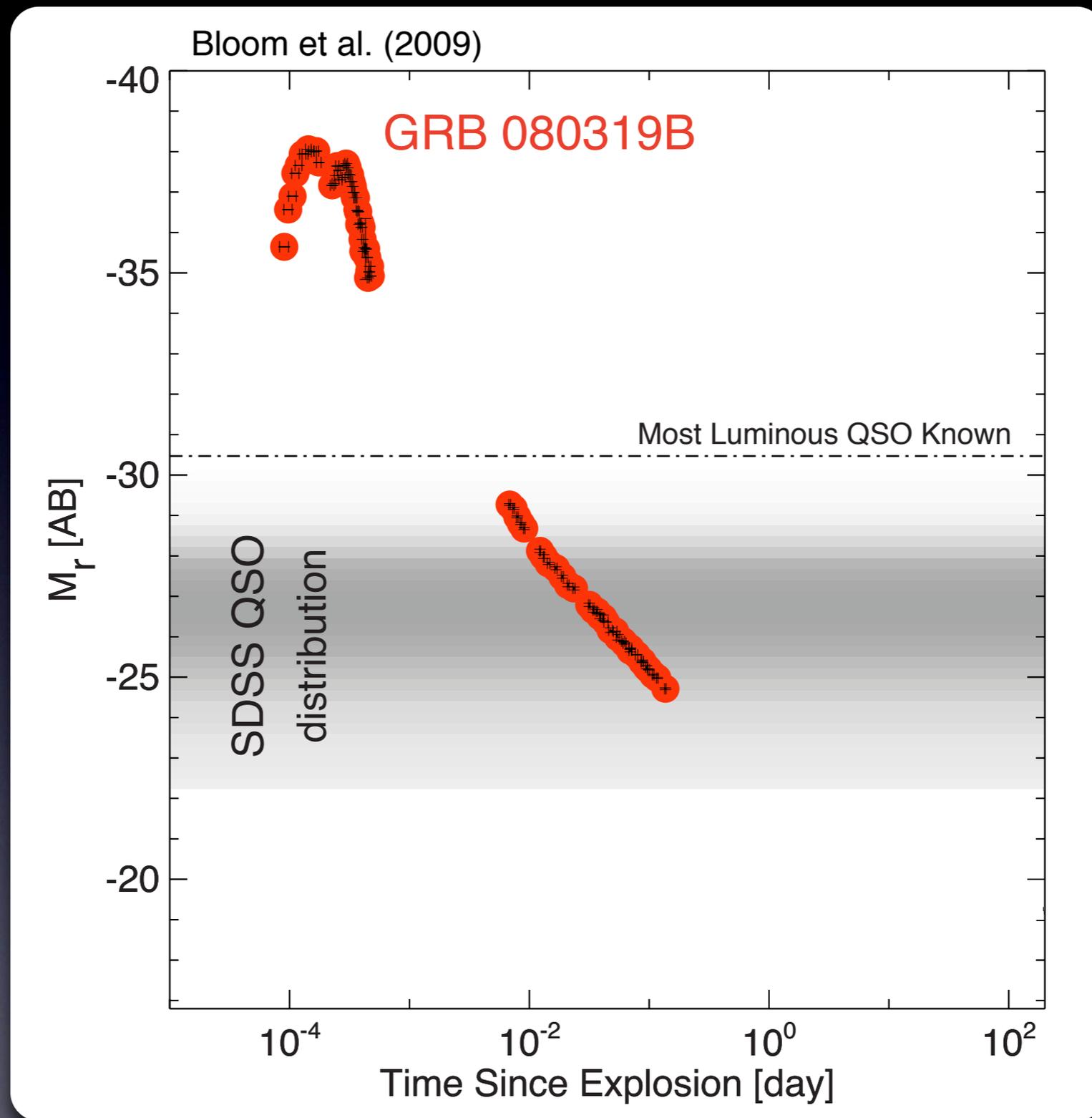
$1/10^5 \text{ yr}^{-1}$: rate in a galaxy

Several a day: full sky rate detectable from Earth

GRBs as cosmological probes

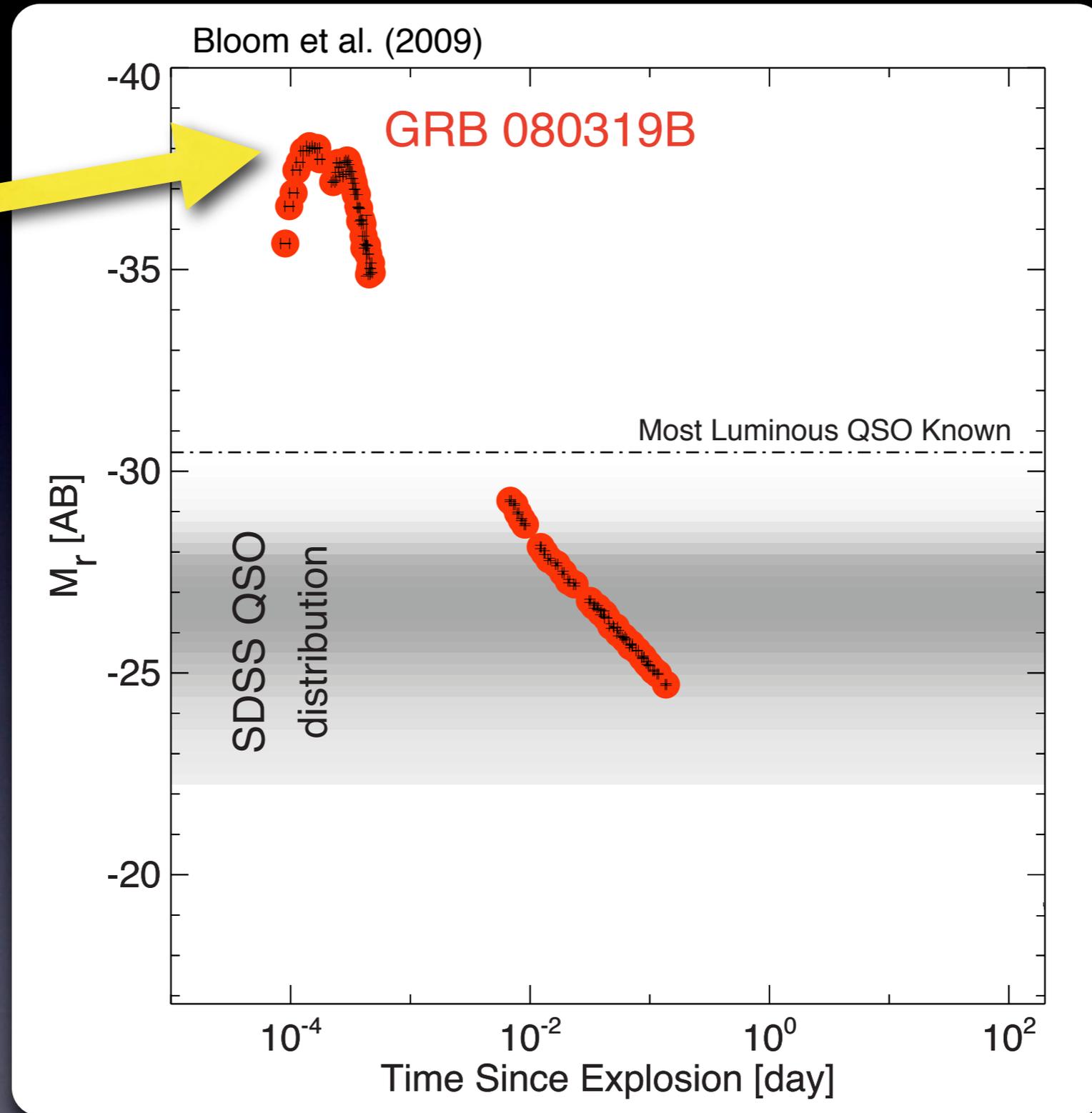


Golden Age of Gamma-Ray Bursts



Golden Age of Gamma-Ray Bursts

Visual
magnitude
 $m = 5.6$

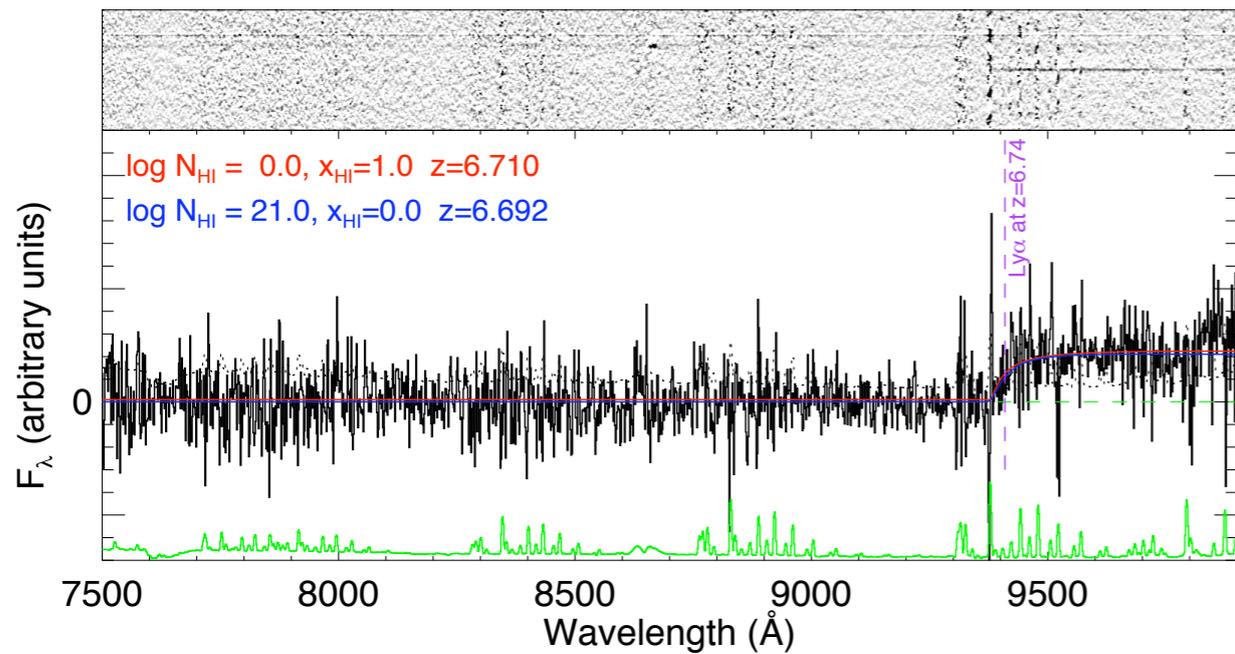


GRB 080319B

Brightest source recorded by humanity ($z=0.937$)

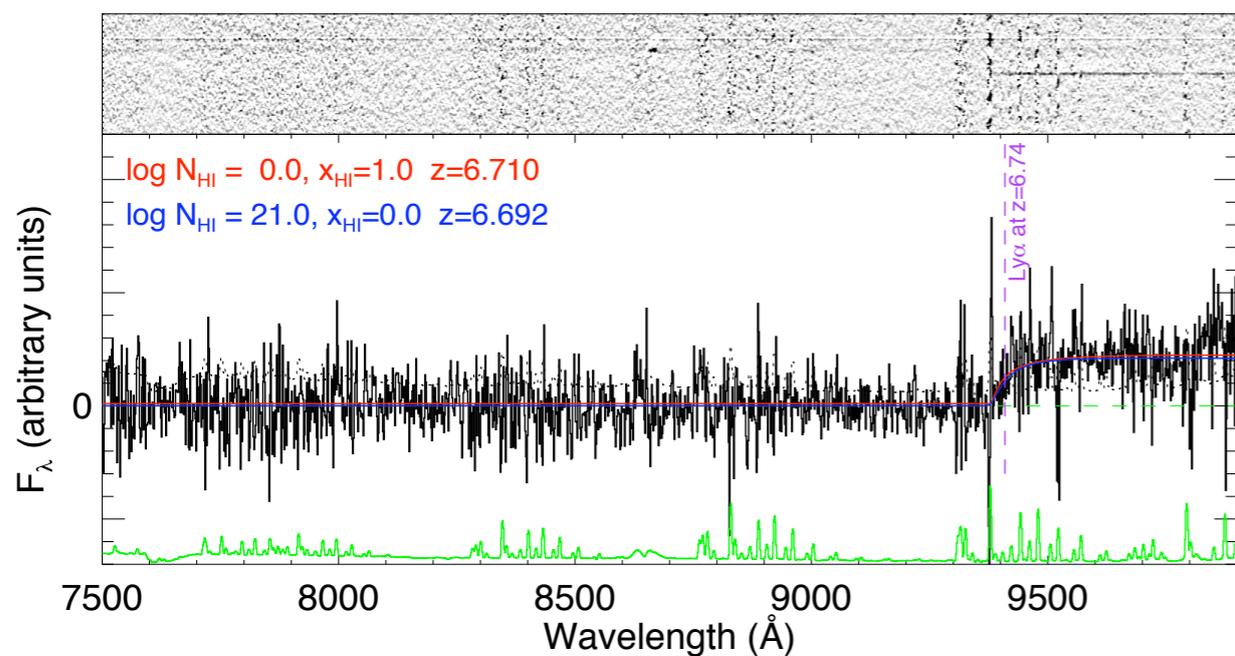
Golden Age of Gamma-Ray Bursts

Greiner, Krühler, Fynbo, et al. (2010)



Golden Age of Gamma-Ray Bursts

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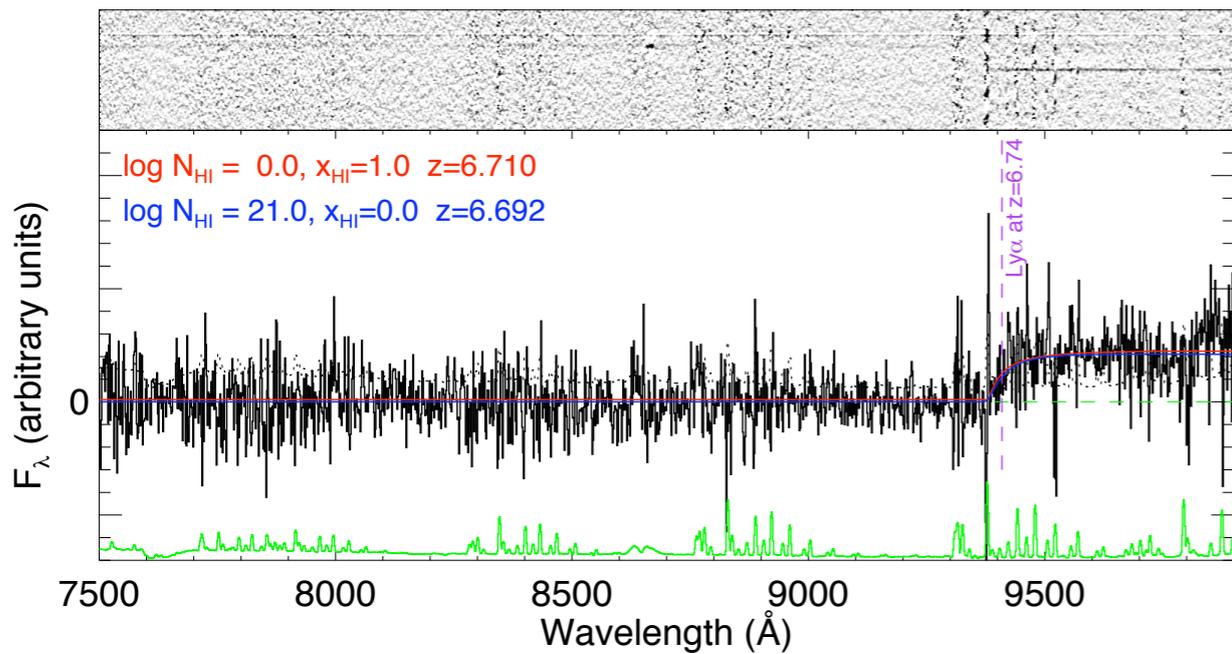


GRB 080913

Second most distant object known ($z=6.7$)

Golden Age of Gamma-Ray Bursts

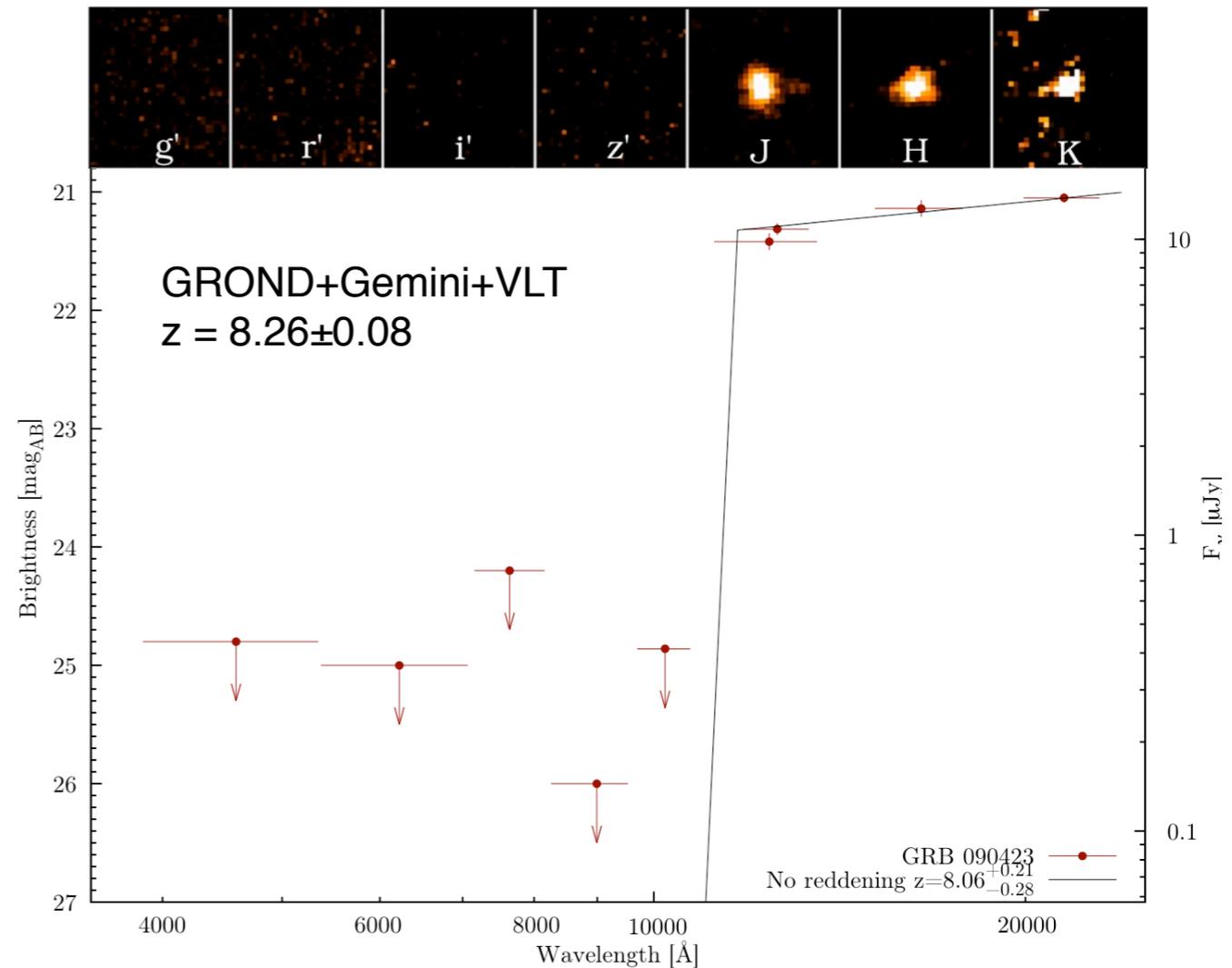
Greiner, Krühler, Fynbo, et al. (2010)



GRB 080913

Second most distant object known ($z=6.7$)

Tanvir et al. (2009), Salvaterra et al. (2009)

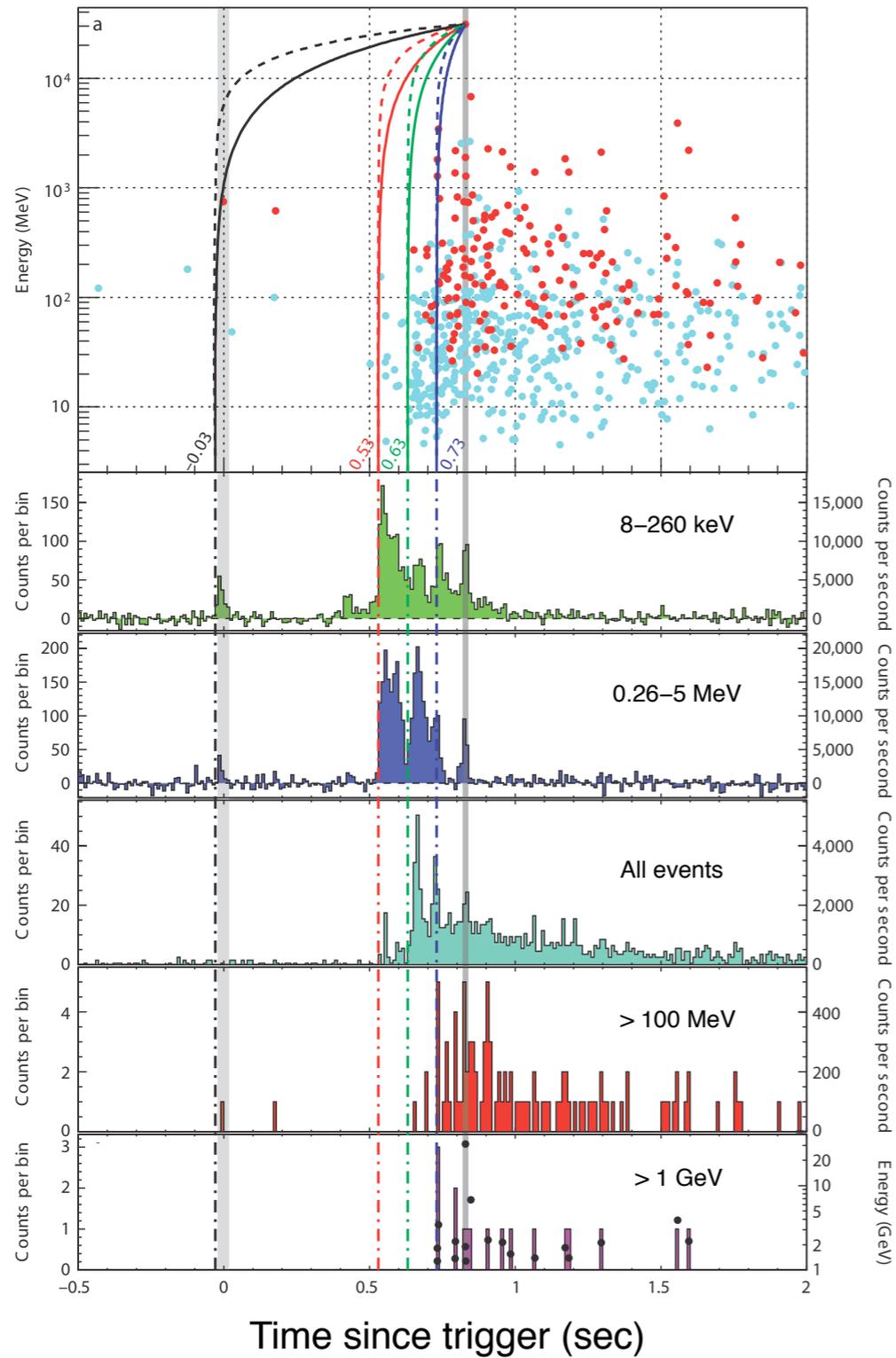


GRB 090423

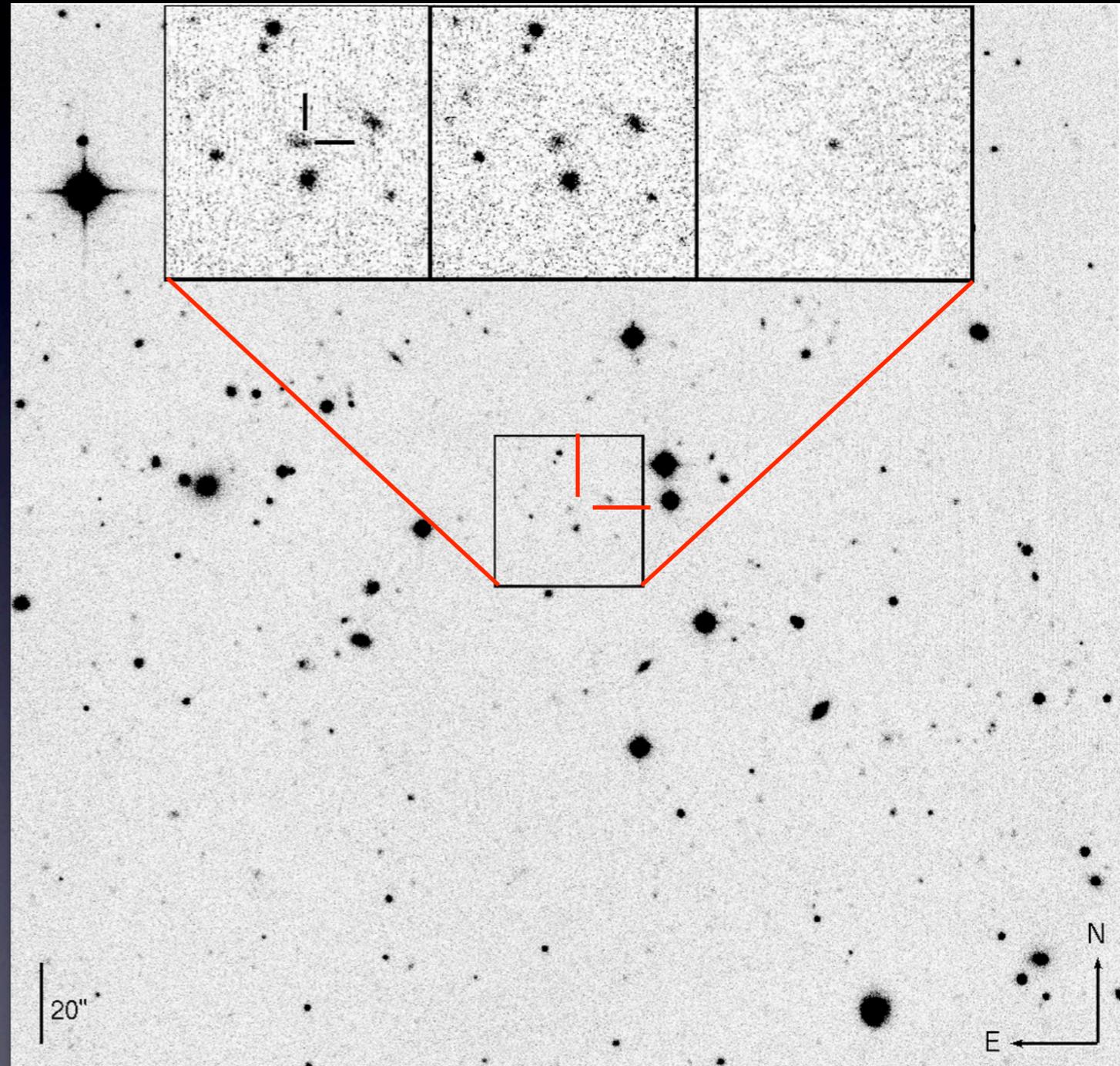
The most distant object known ($z=8.26$)

Golden Age of Gamma-Ray Bursts

Abdo et al. (2009a)

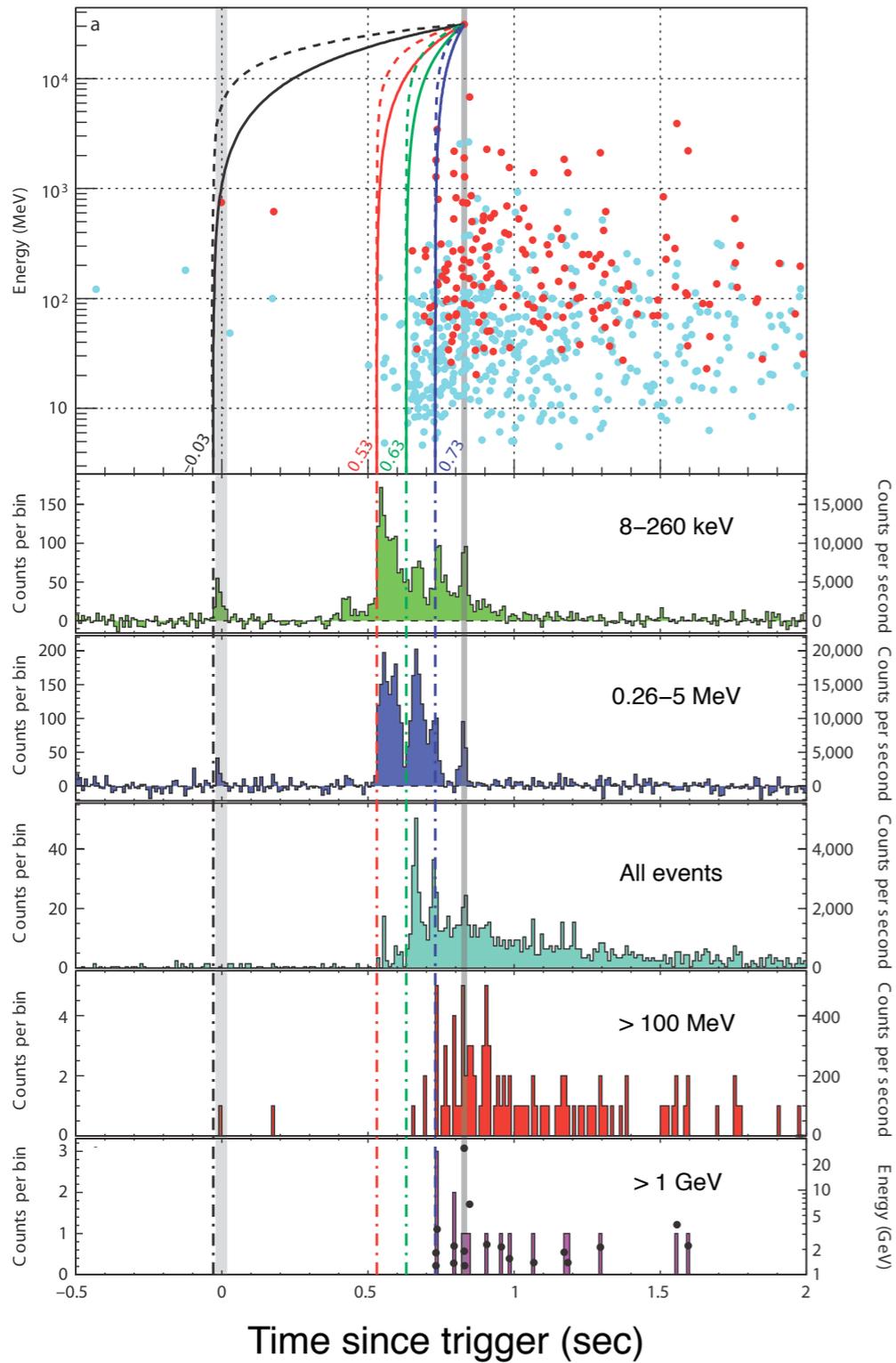


McBreen, Krühler, Rau, Greiner et al. (2010)
Rau, McBreen, Krühler, Greiner (2009)

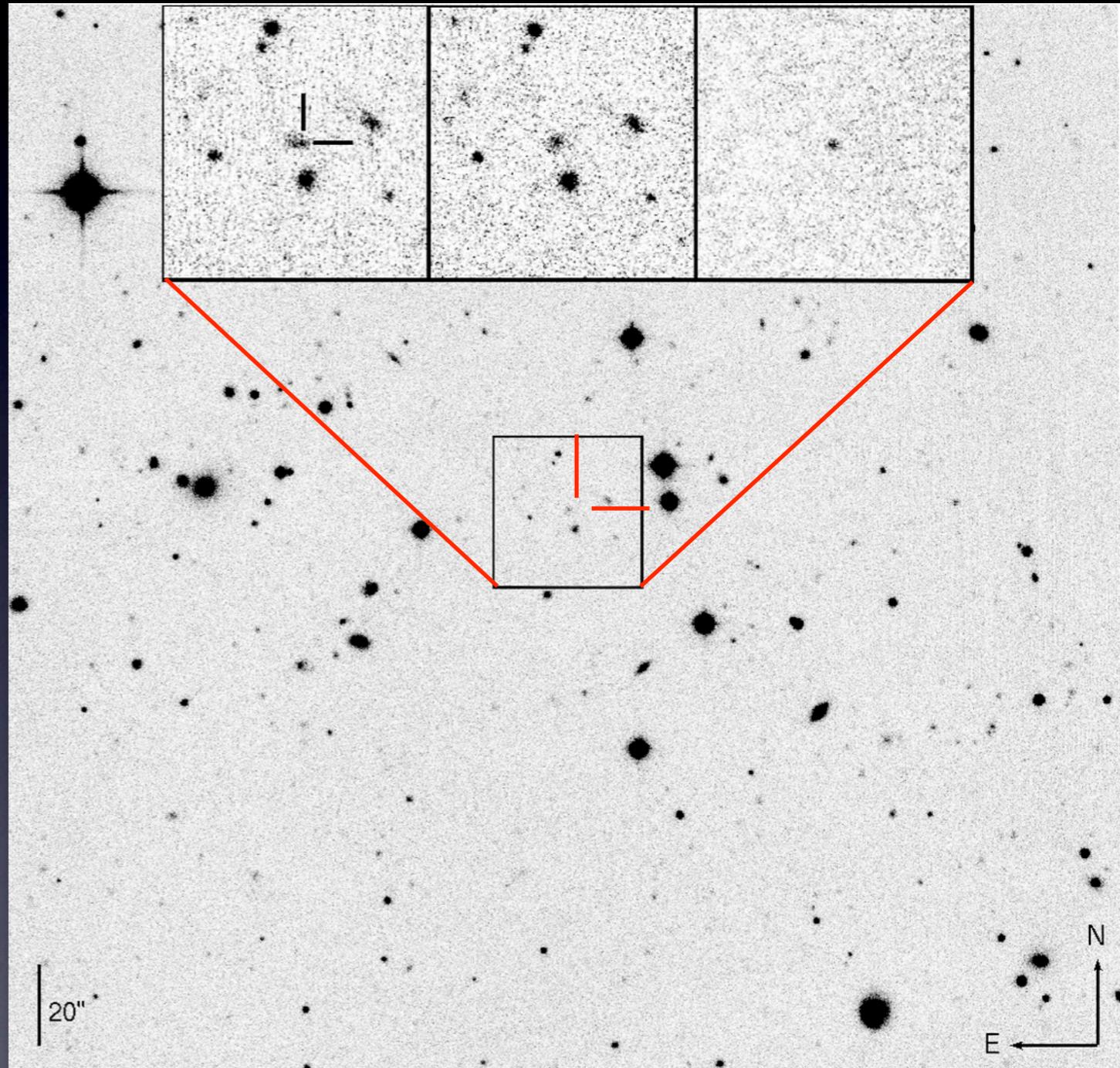


Golden Age of Gamma-Ray Bursts

Abdo et al. (2009a)



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Rau, McBreen, Krühler, Greiner (2009)

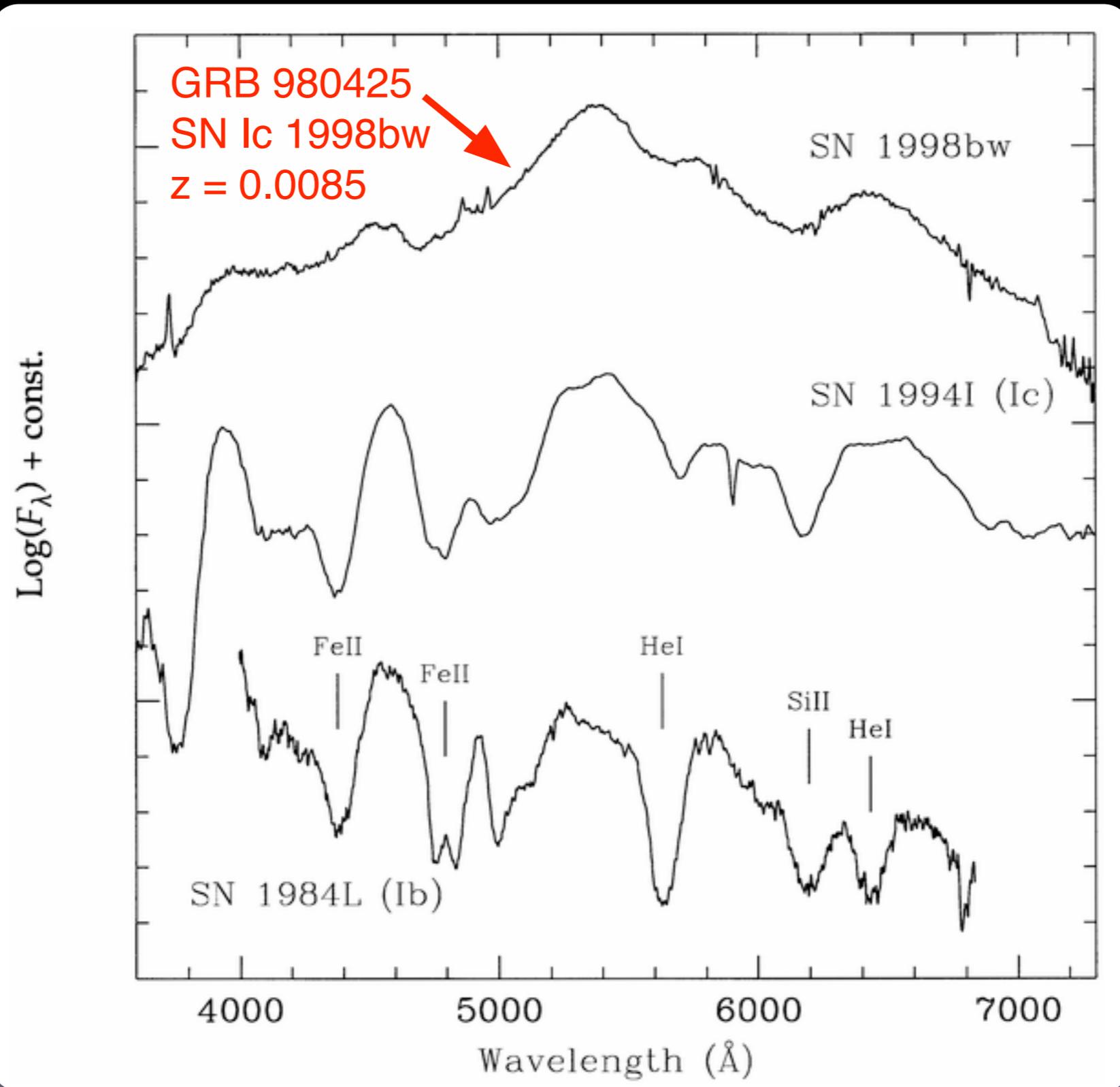


GRB 090510

Lorentz Invariance tested to highest precision ($z=0.903$)

Gamma-ray burst - SN connection

Gamma-ray burst - SN connection



Gamma-ray burst - SN connection

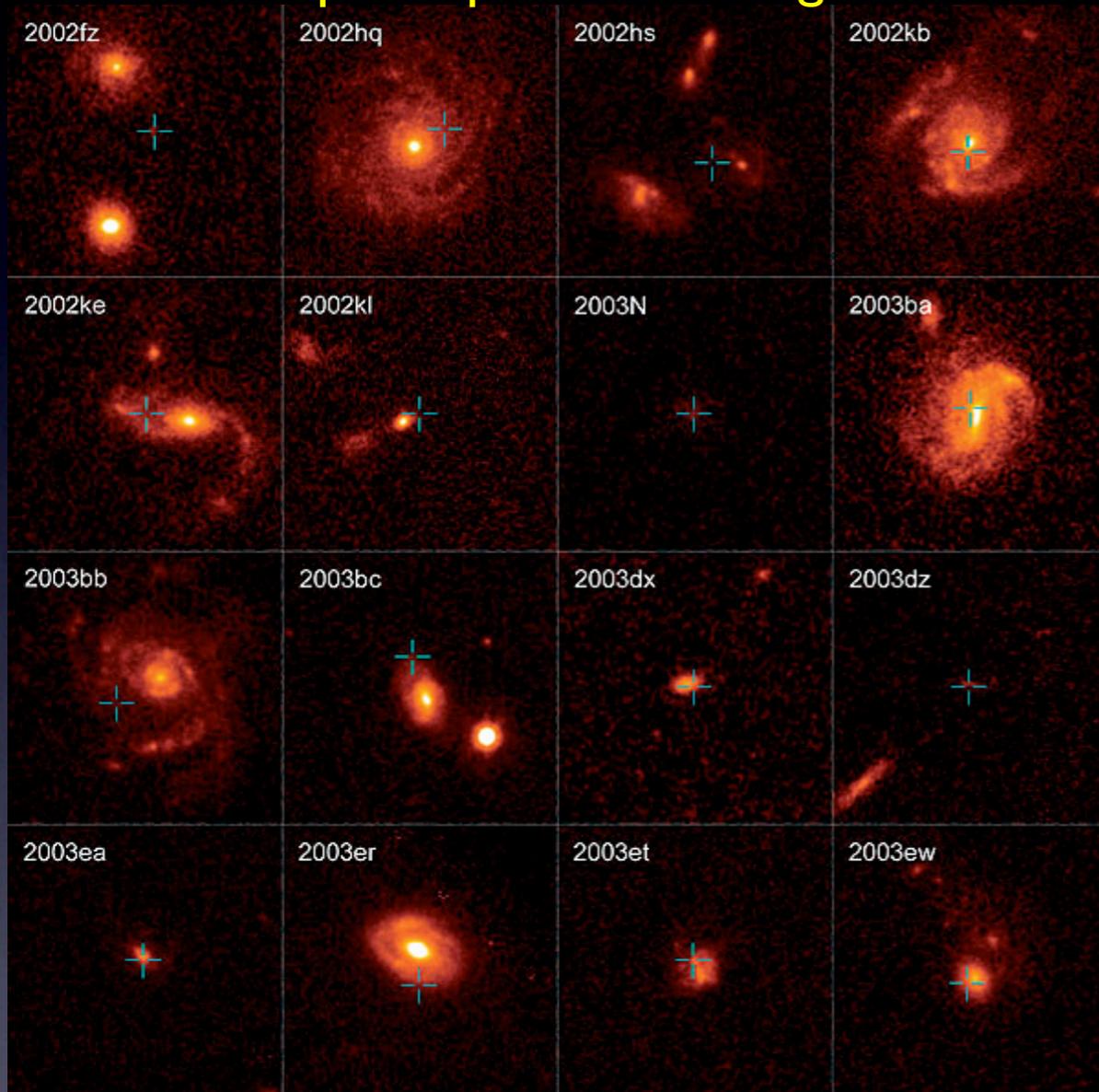
Gamma-ray burst - SN connection

GRB with spectroscopically confirmed SN

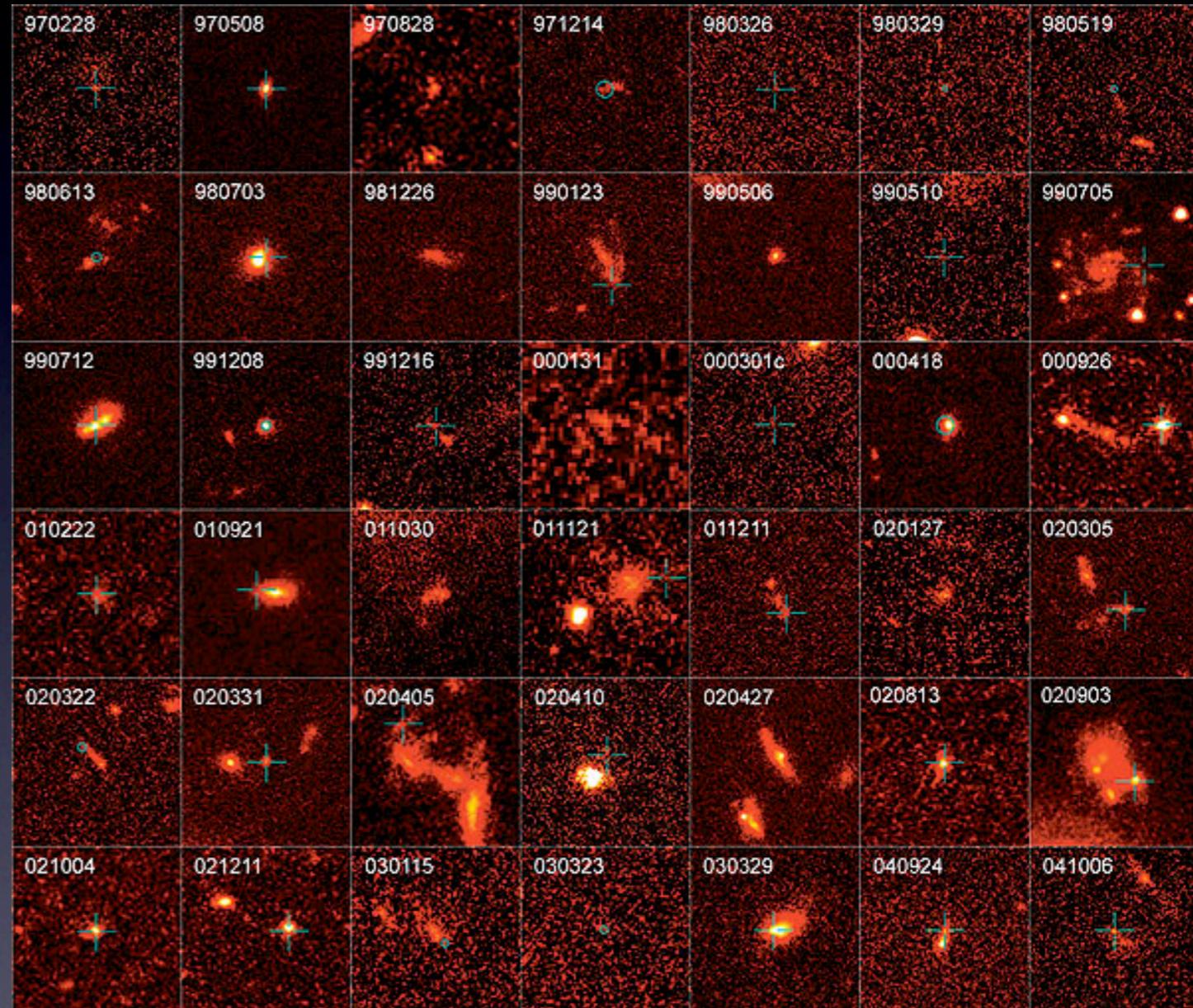
GRB	z	12 + log(O/H) (T_e)	Host type	M_B	References
980425	0.0085	8.25	Dwarf spiral	-17.6	Hammer et al. (2006)
020903	0.25	7.97	Irregular	-18.8	Hammer et al. (2006)
030329	0.168	7.72	Irregular	-16.5	Levesque et al. (2010)
031203	0.105	8.02 ± 0.15	Irregular	-21.0	Prochaska et al. (2004)
060218	0.0335	7.54+0.16	Irregular	-15.9	Wiersema et al. (2007)
100316D	0.0591	8.23 ± 0.15	Spiral? Irr?	-19	Starling et al. (2010)

GRB environment

Core-collapse supernova host galaxies

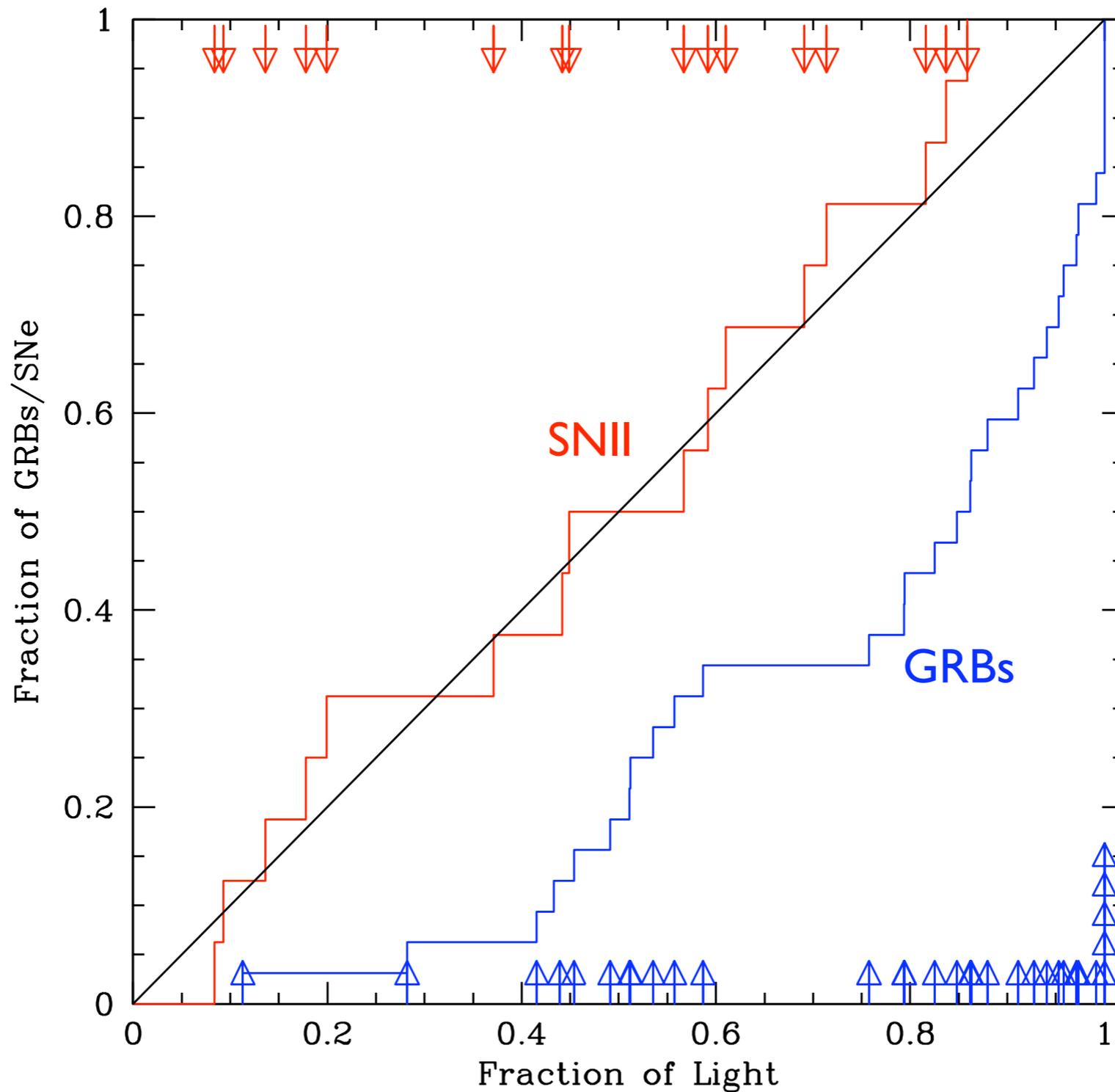


GRB host galaxies

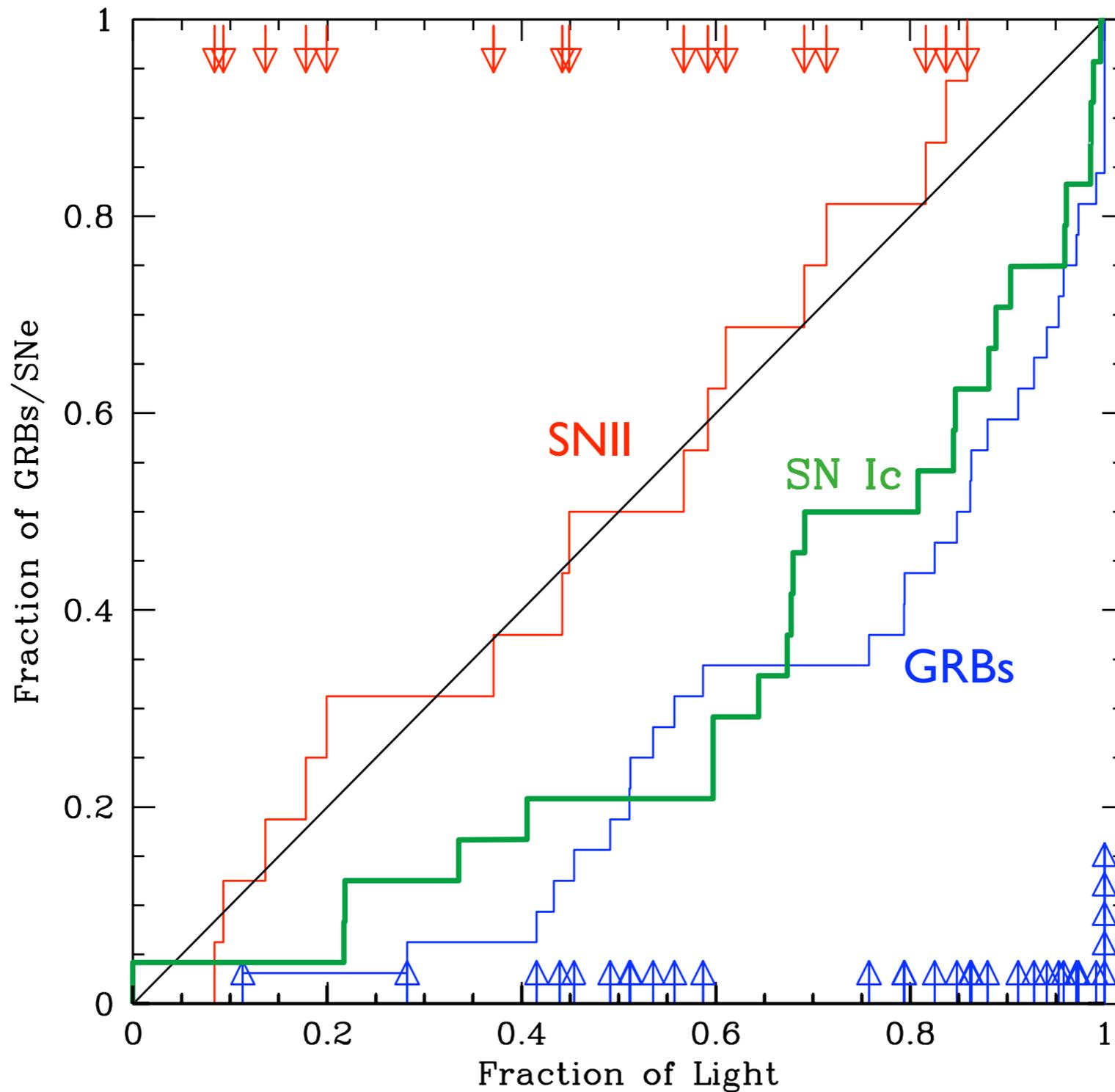


Fruchter et al. (2006)
(see also Svensson et al. 2010)

GRB environment



GRB environment

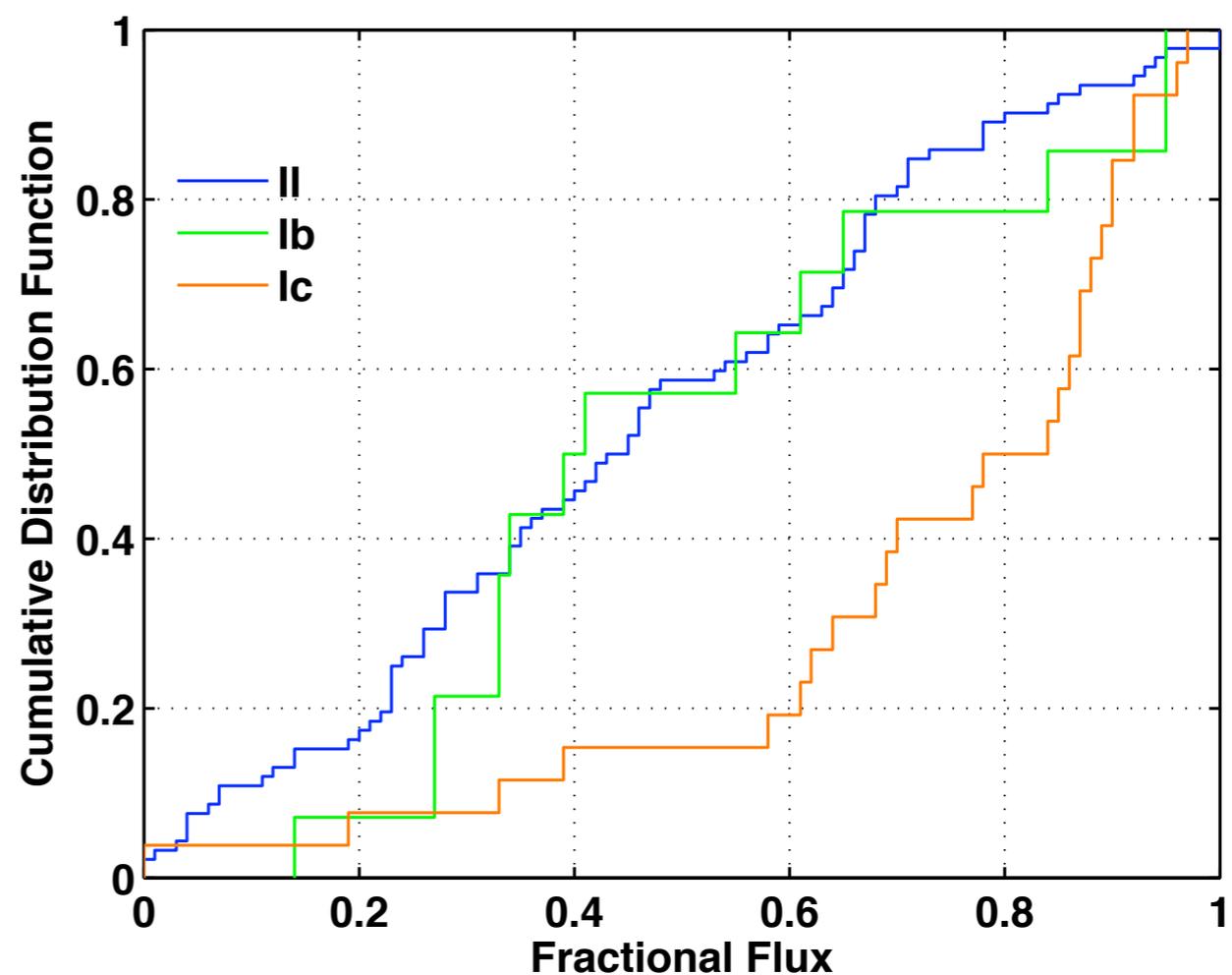
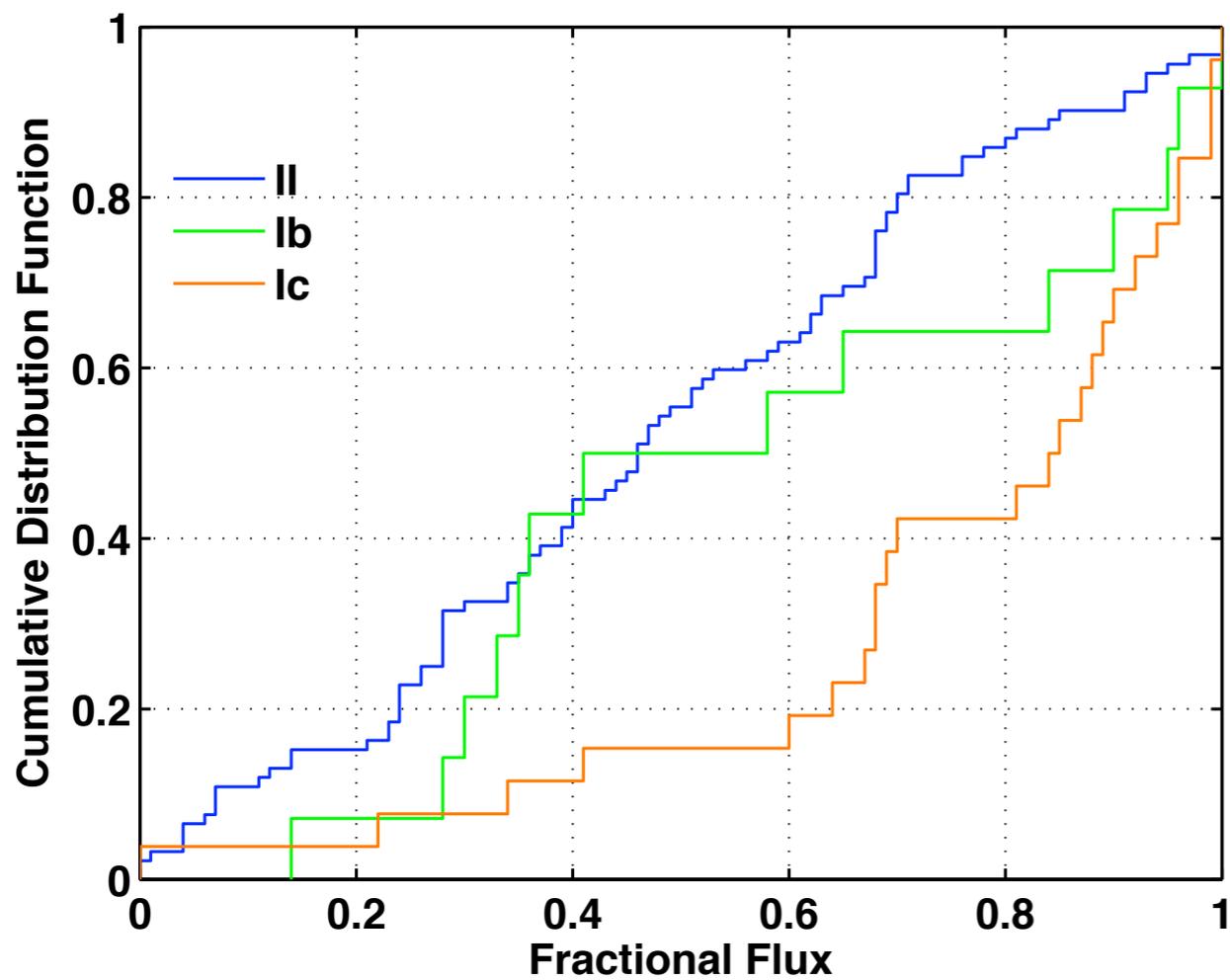


Fruchter et al. (2006)
(see also Svensson et al. 2010)

Kelly, Kirshner & Pahre (2008)

GRB environment

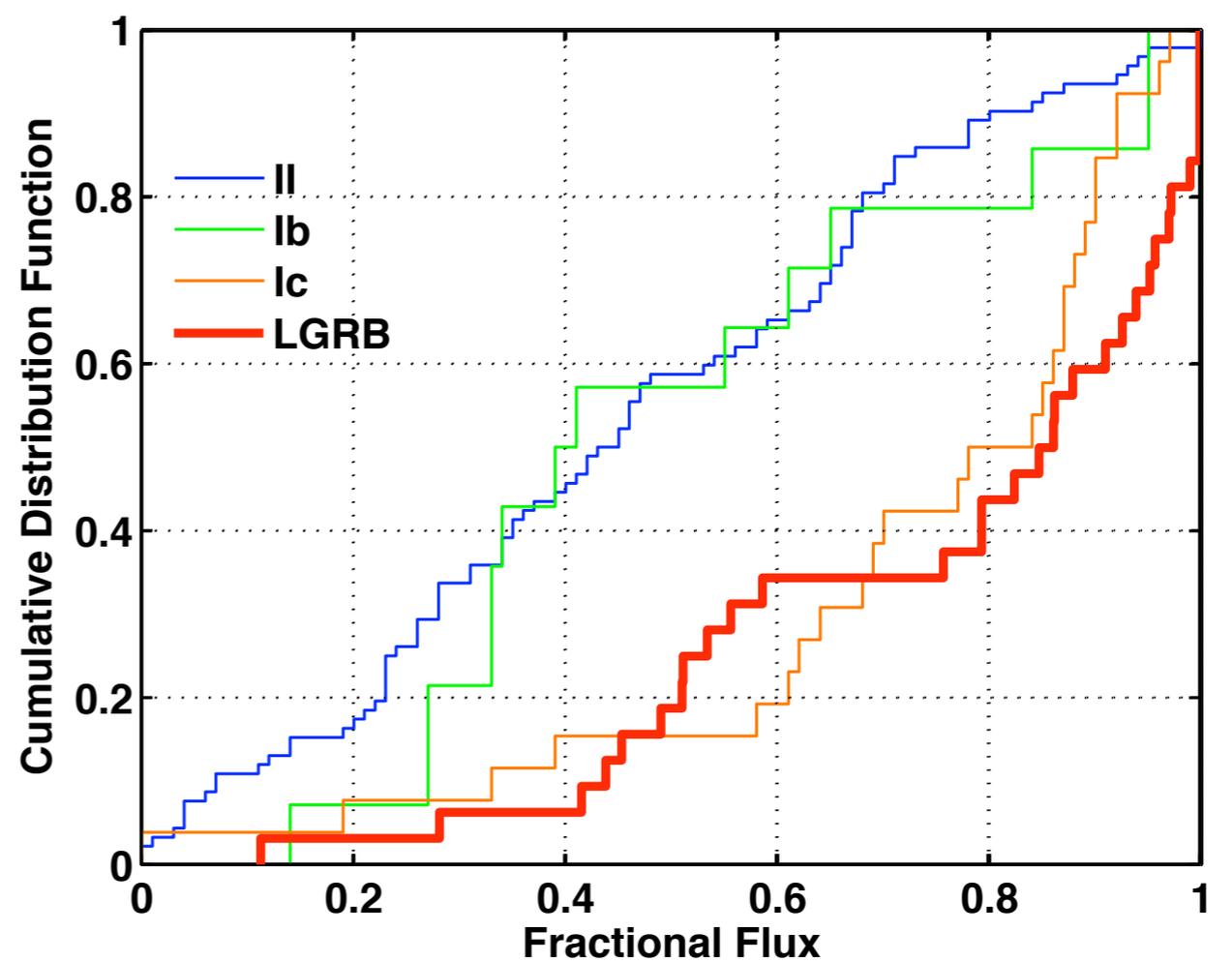
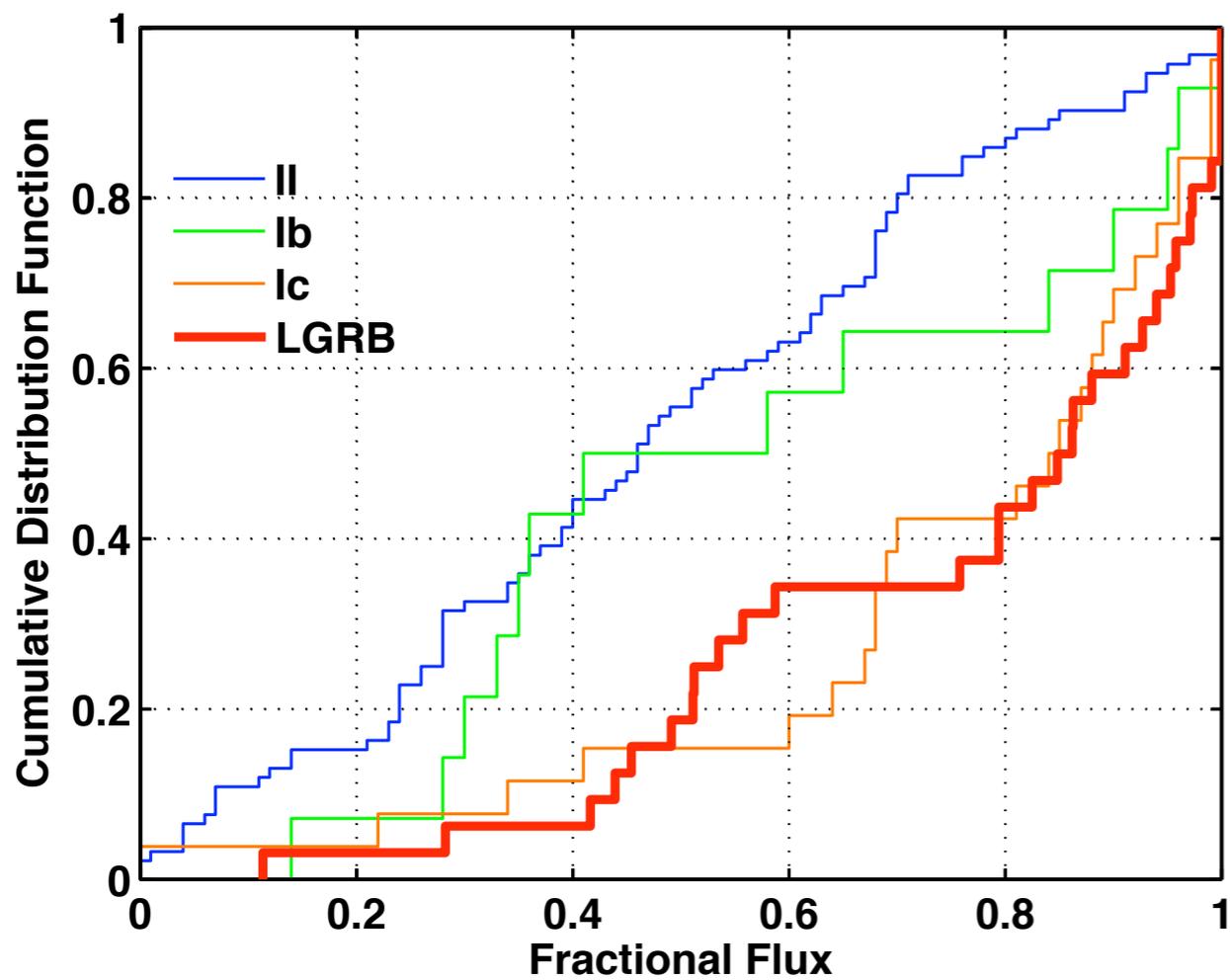
Light distribution of Wolf-Rayet stars vs. GRBs



Leloudas et al. (2010)
(see also Han et al. 2010)

GRB environment

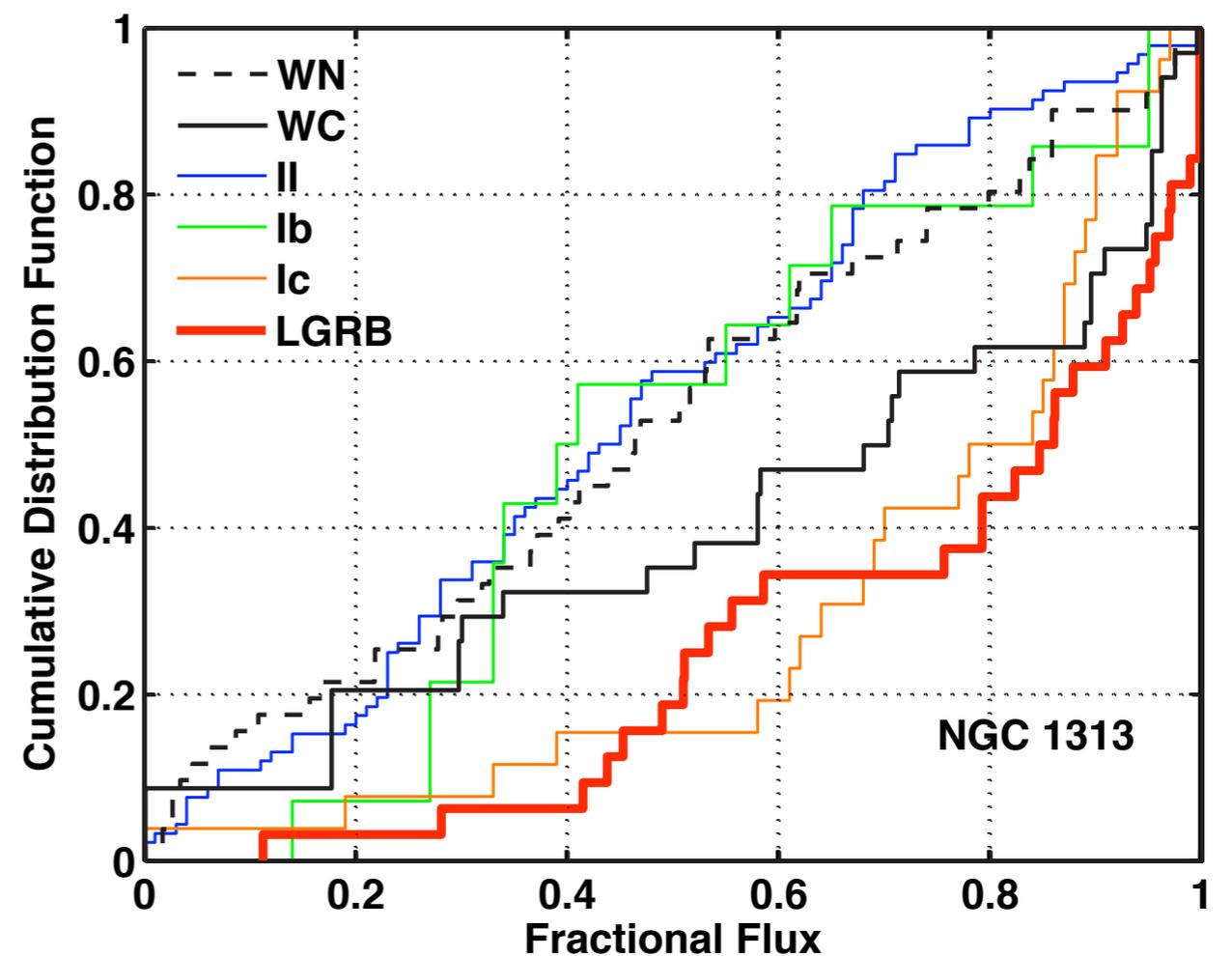
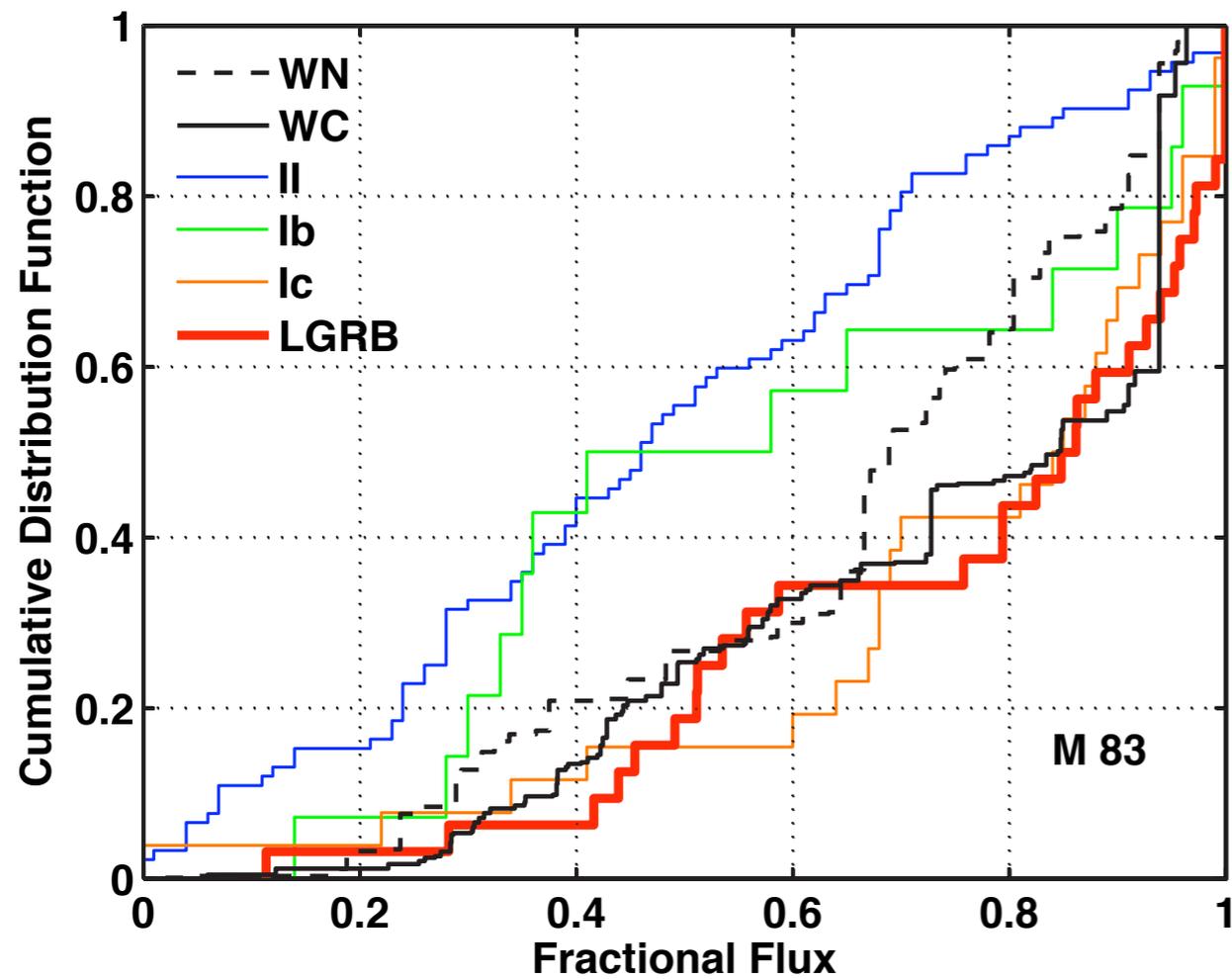
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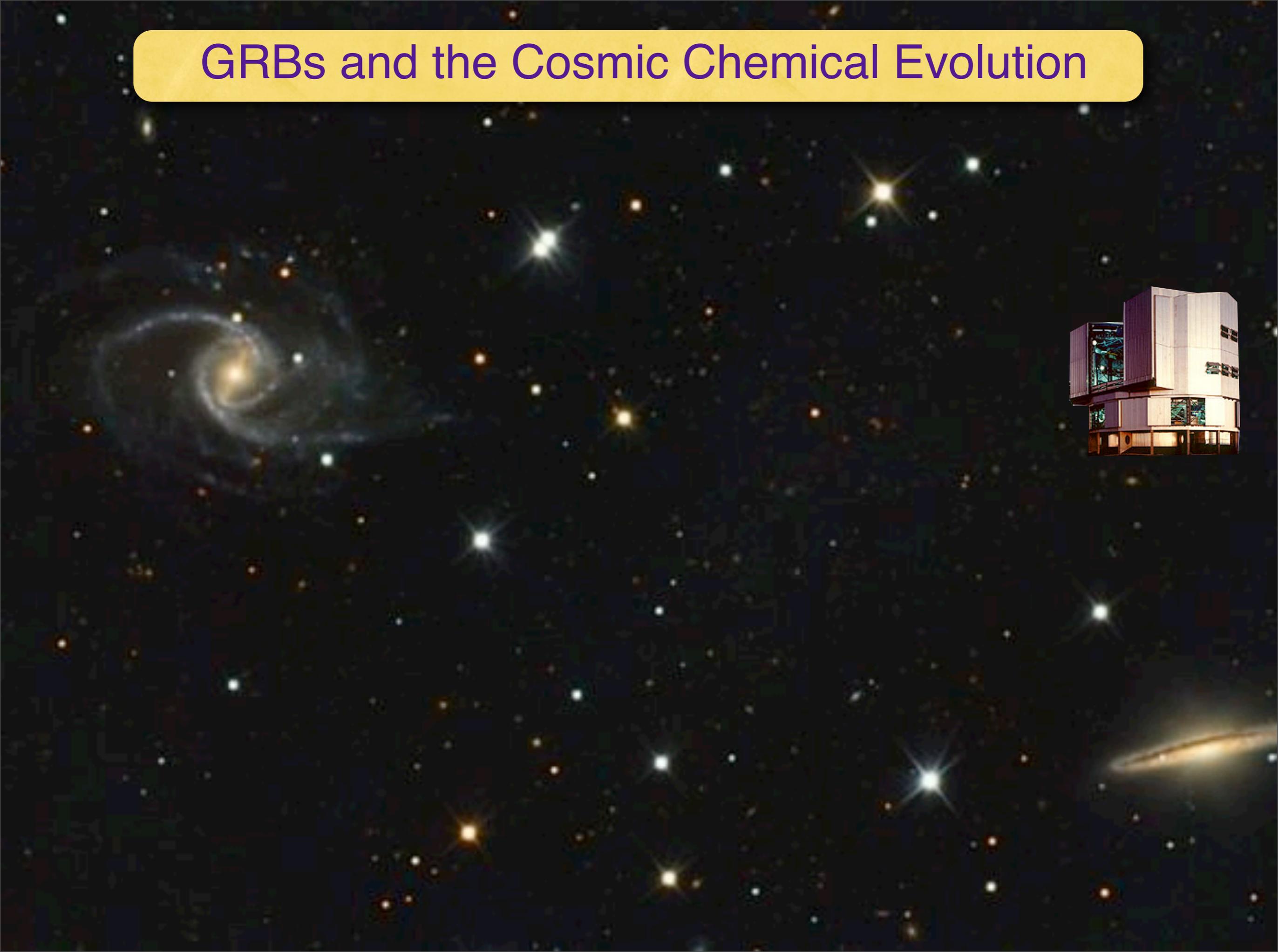
*GRBs to investigate the
Cosmic Chemical Evolution*

GRBs and the Cosmic Chemical Evolution

QSO



GRBs and the Cosmic Chemical Evolution



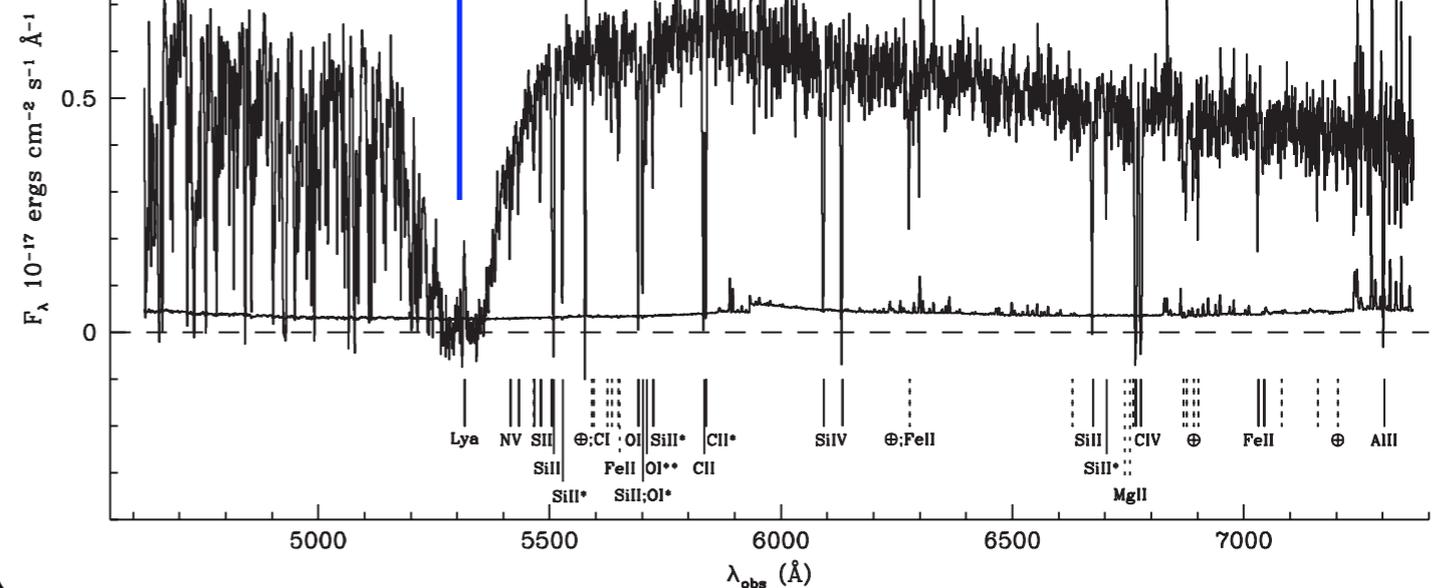
GRBs and the Cosmic Chemical Evolution

GRB



GRB 030323 $z=3.372$ (Vreeswijk et al. 2004)

GRB damped Lyman α (DLA)

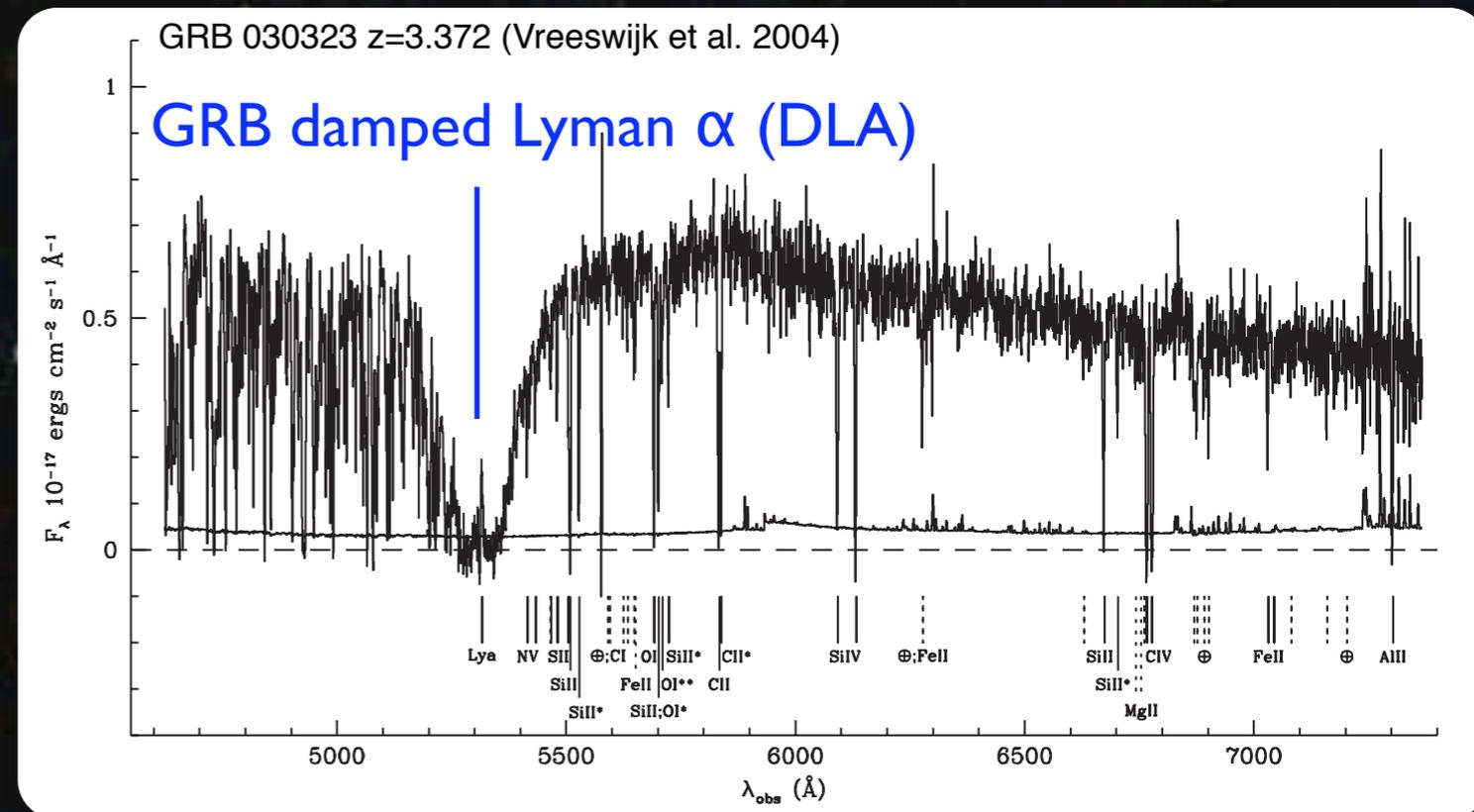


GRBs and the Cosmic Chemical Evolution

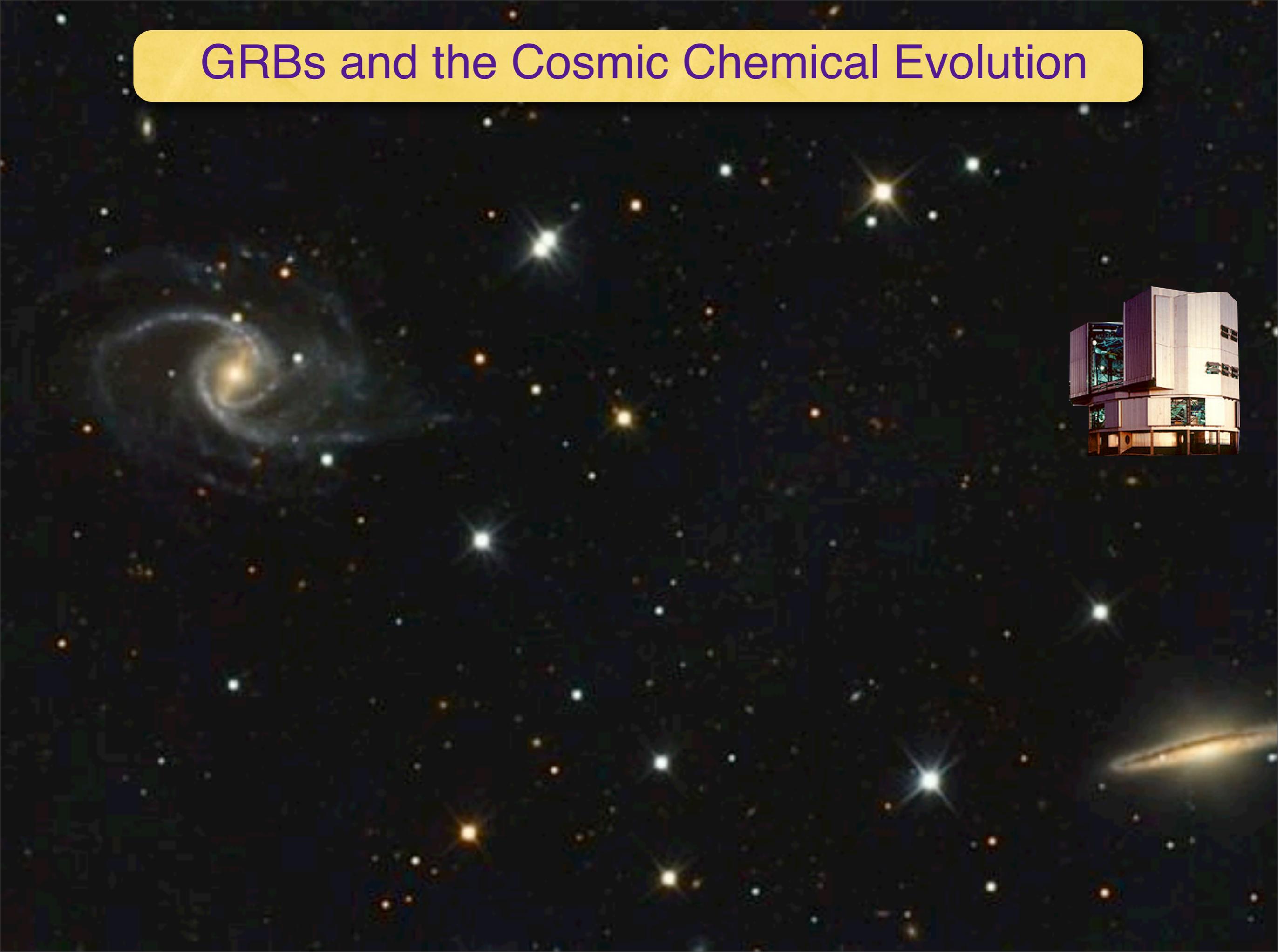
GRB



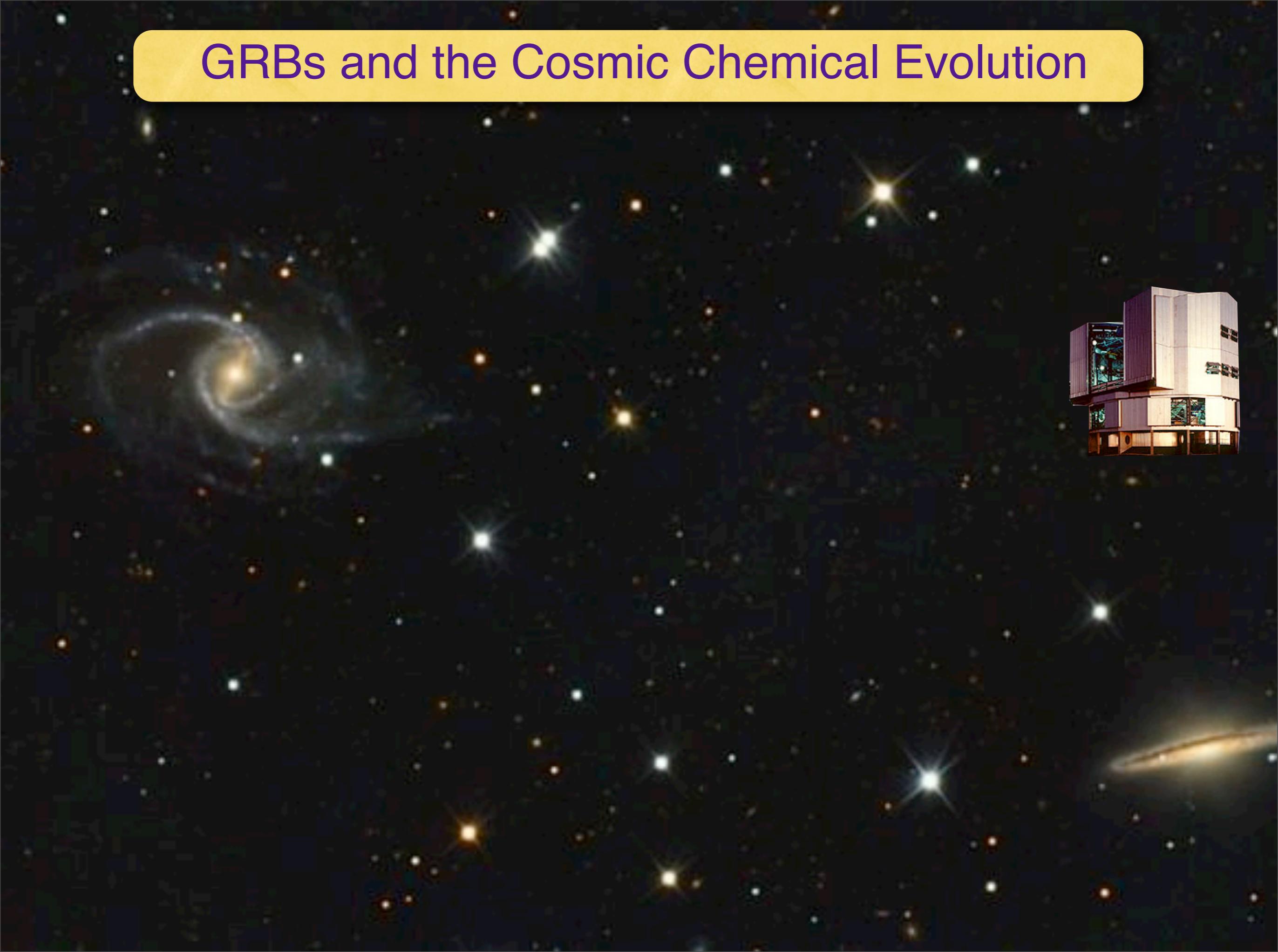
Cold interstellar medium ($T \lesssim 1000$ K)
Heavy element enrichment
Dust extinction
Dust depletion
Molecular hydrogen



GRBs and the Cosmic Chemical Evolution



GRBs and the Cosmic Chemical Evolution

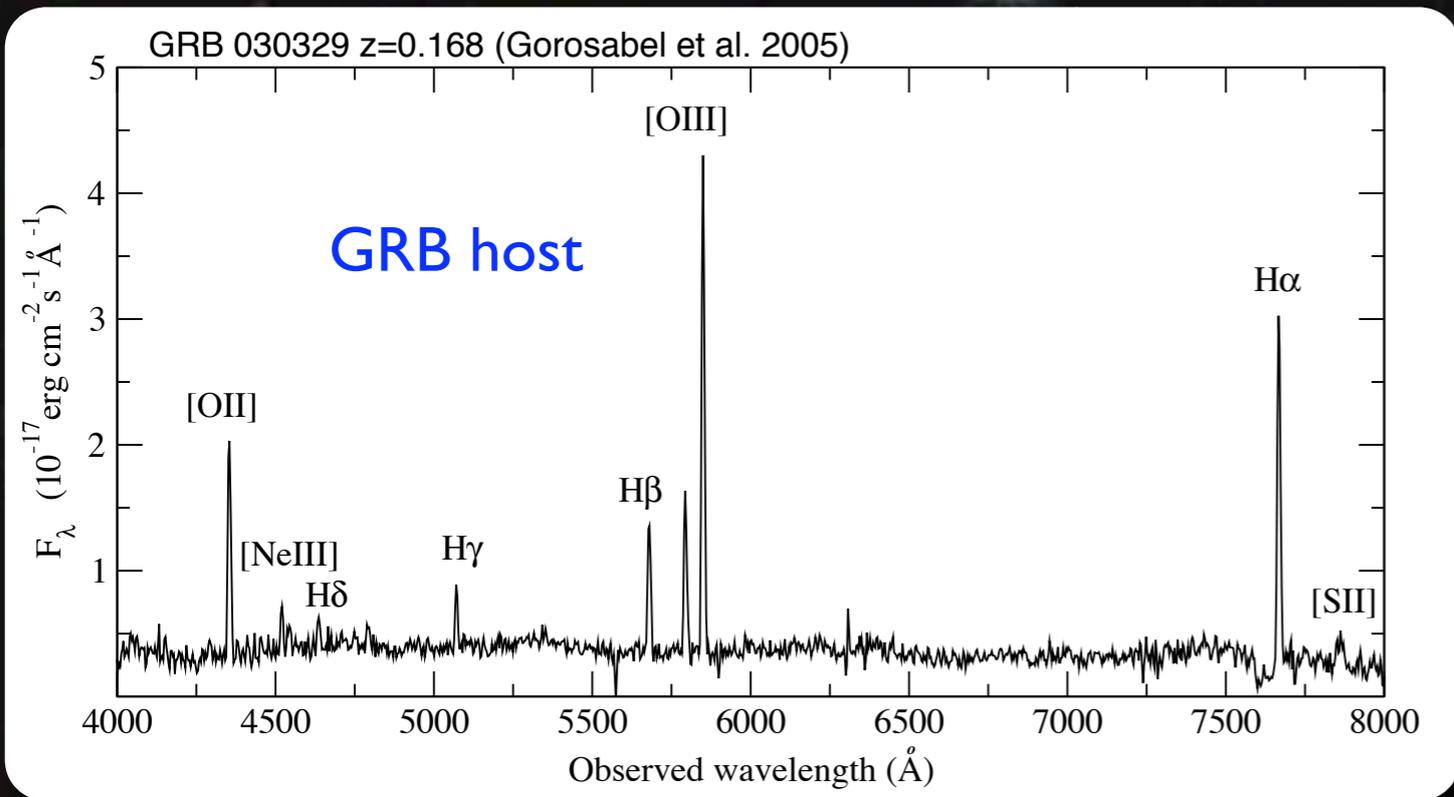
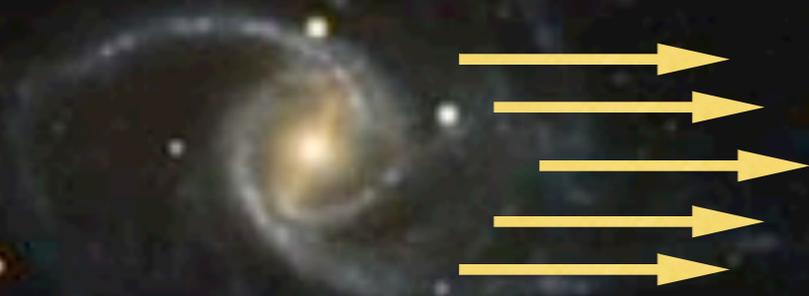


GRBs and the Cosmic Chemical Evolution

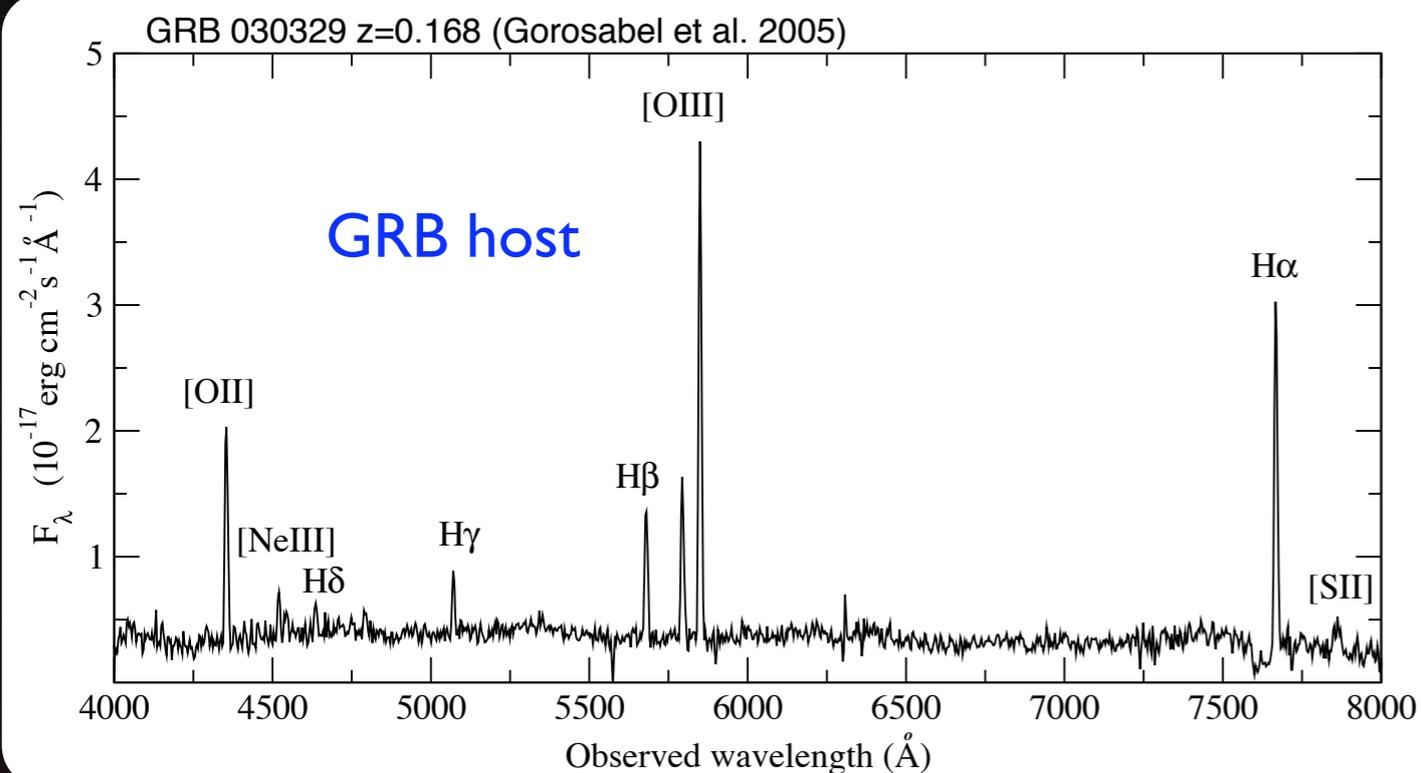
GRB



GRBs and the Cosmic Chemical Evolution

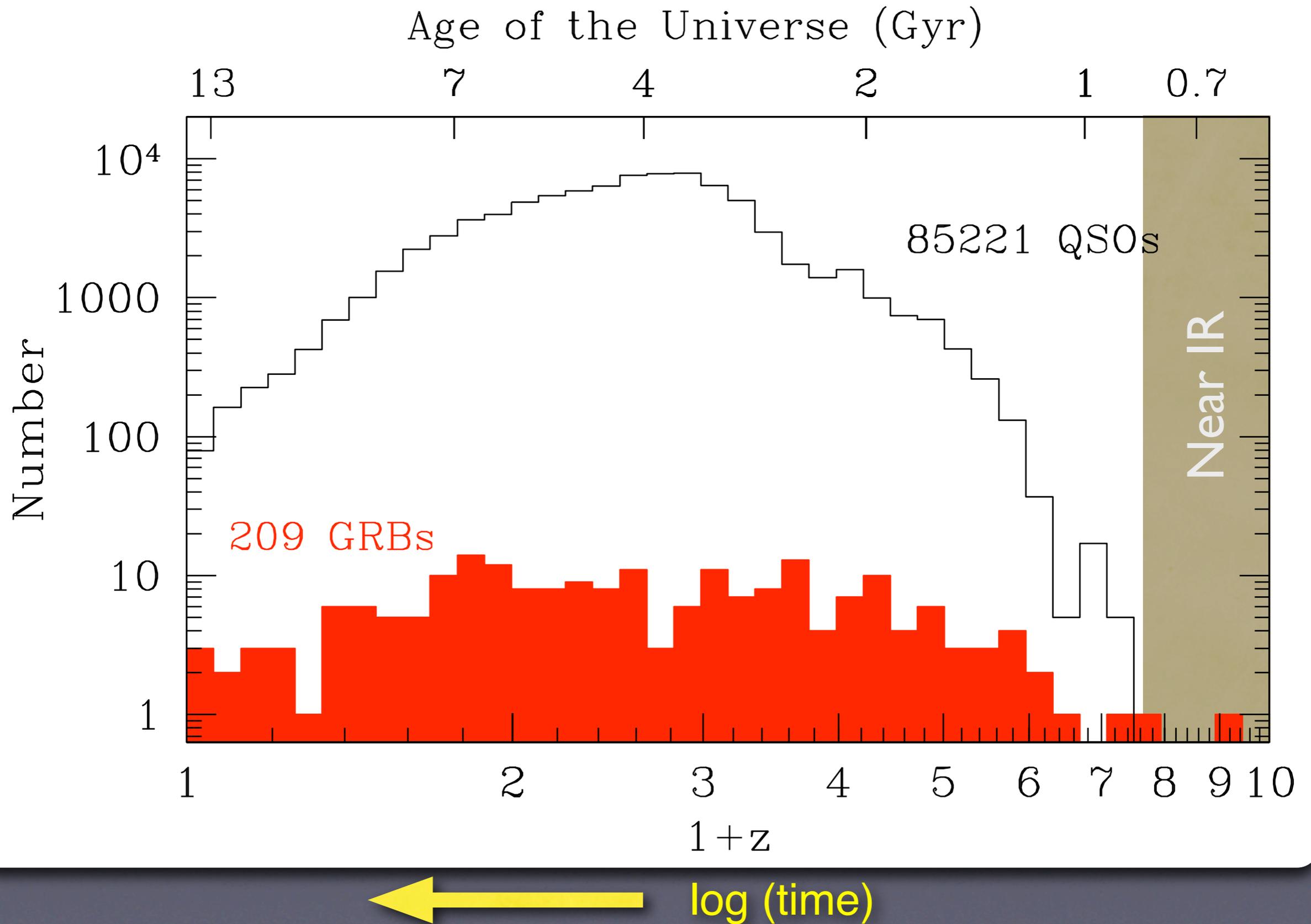


GRBs and the Cosmic Chemical Evolution

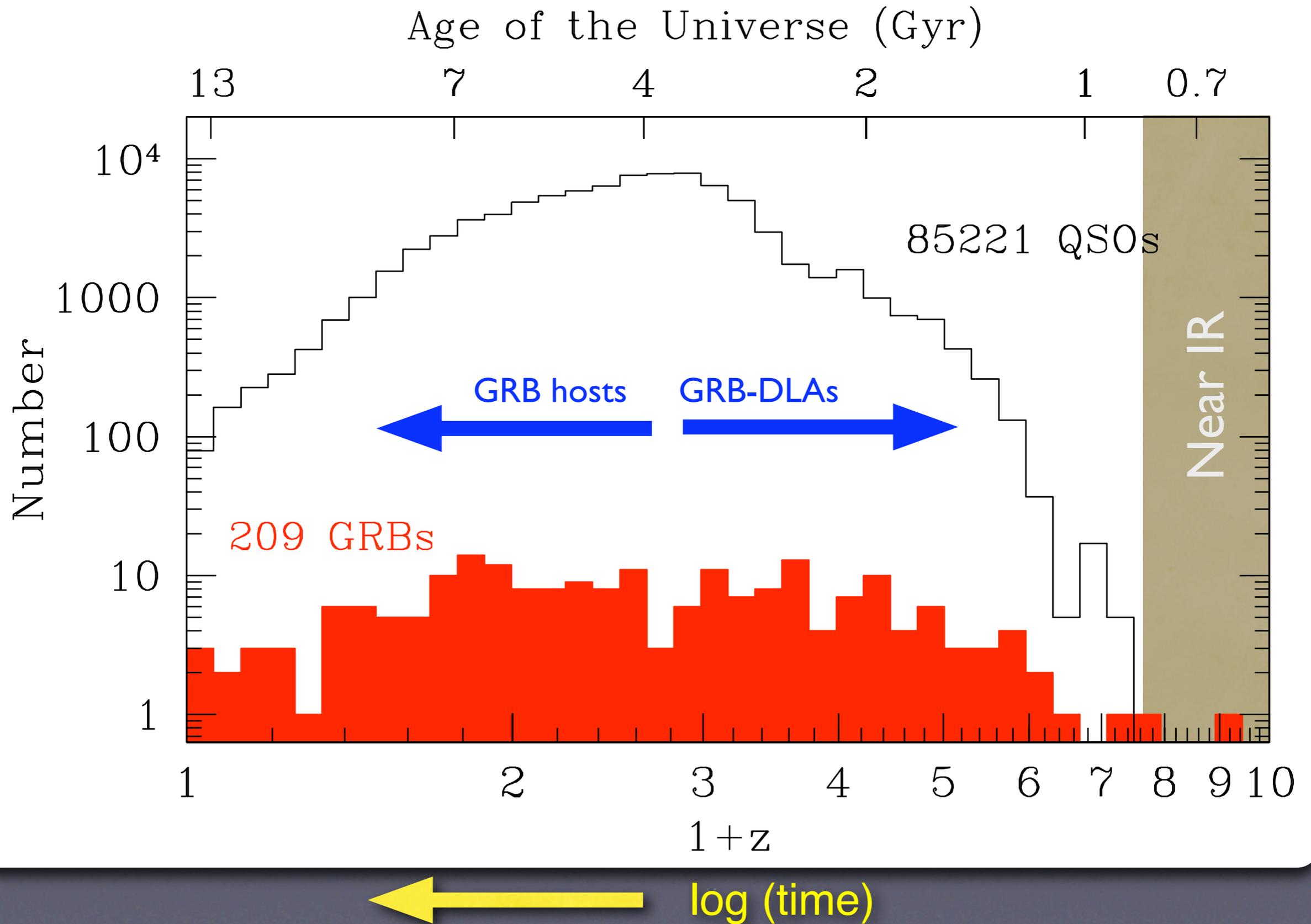


Warm interstellar medium ($T \sim 10^4$ K)
Heavy element enrichment
Dust extinction
Star formation rate
Gas electron density
Gas temperature

GRBs as cosmological probes

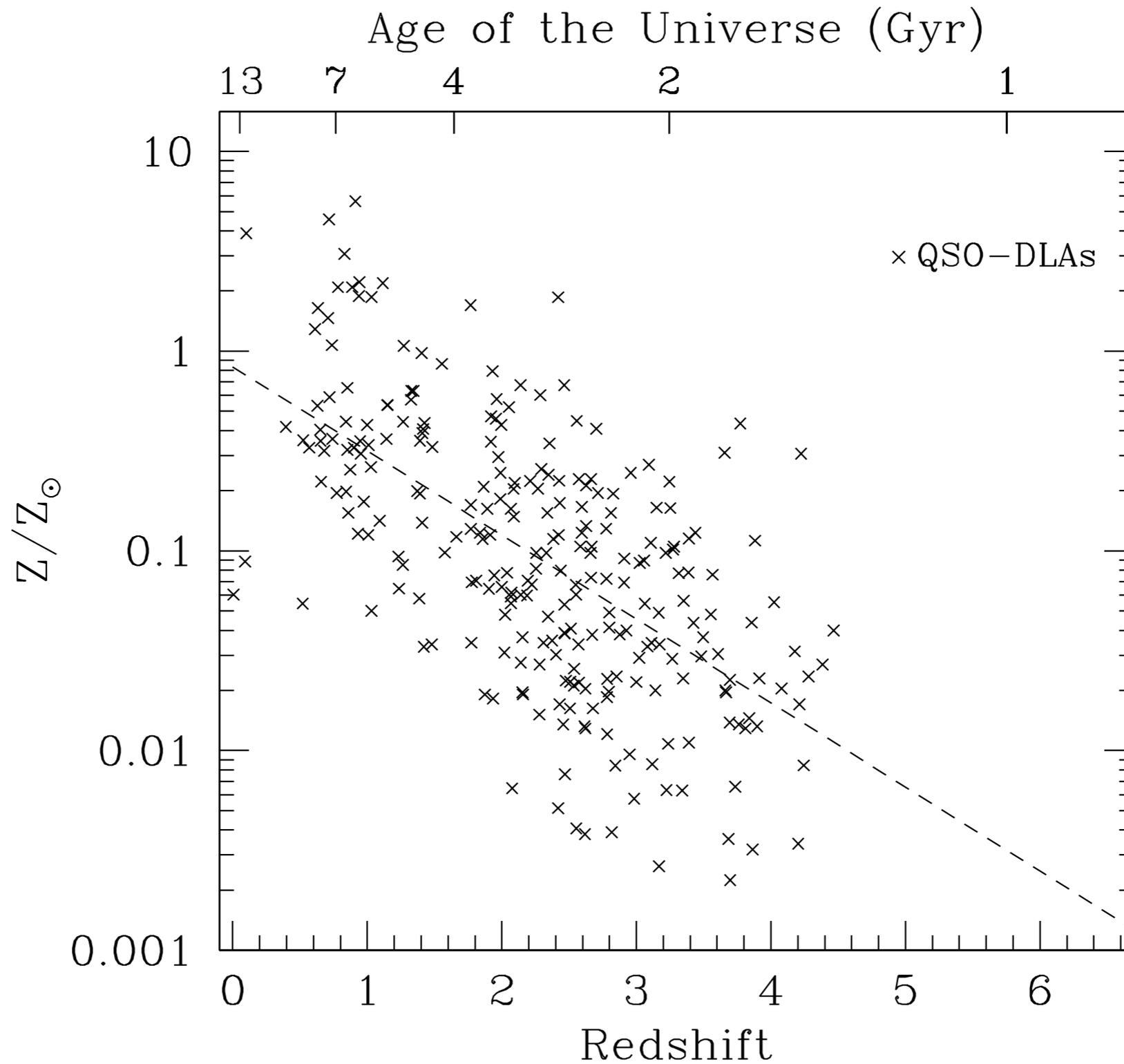


GRBs as cosmological probes



Cosmic chemical evolution with GRBs

Cosmic chemical enrichment



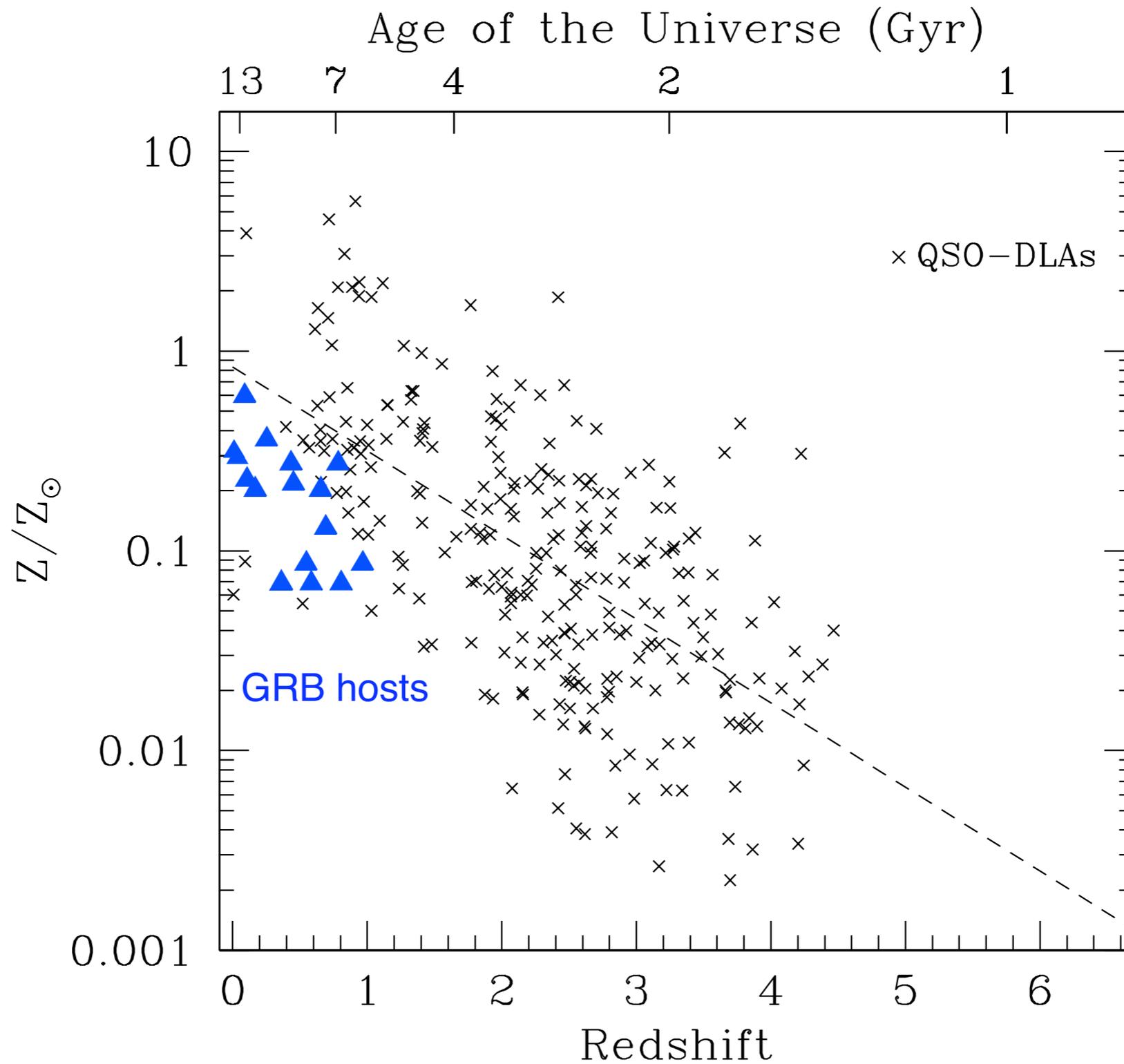
Savaglio (2006)

Rau, Savaglio, Krühler, Afonso, Greiner et al. (ApJ, sub.)

Savaglio, Rau, Greiner, Krühler et al. (Science, sub.)

Cosmic chemical evolution with GRBs

Cosmic chemical enrichment



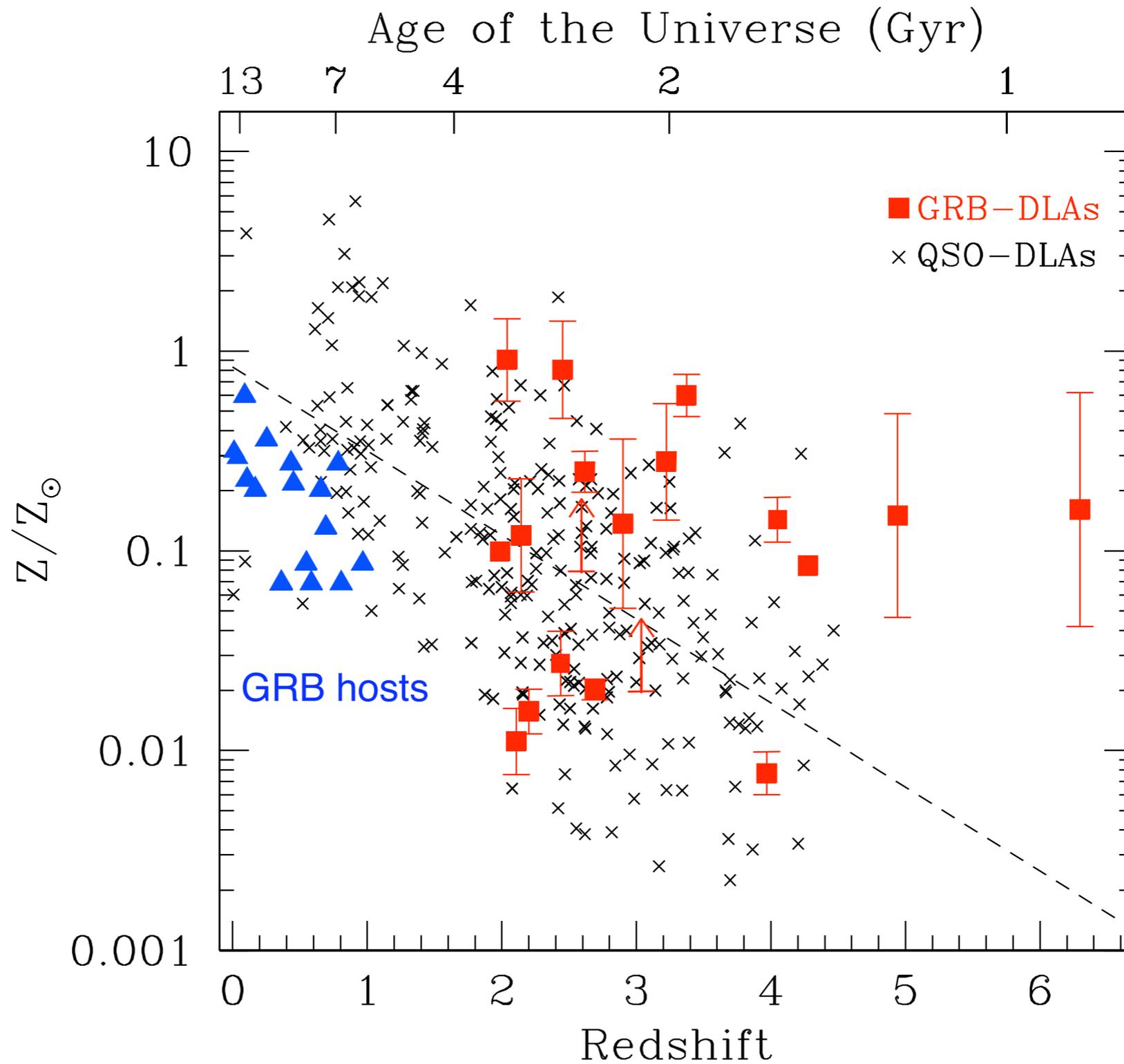
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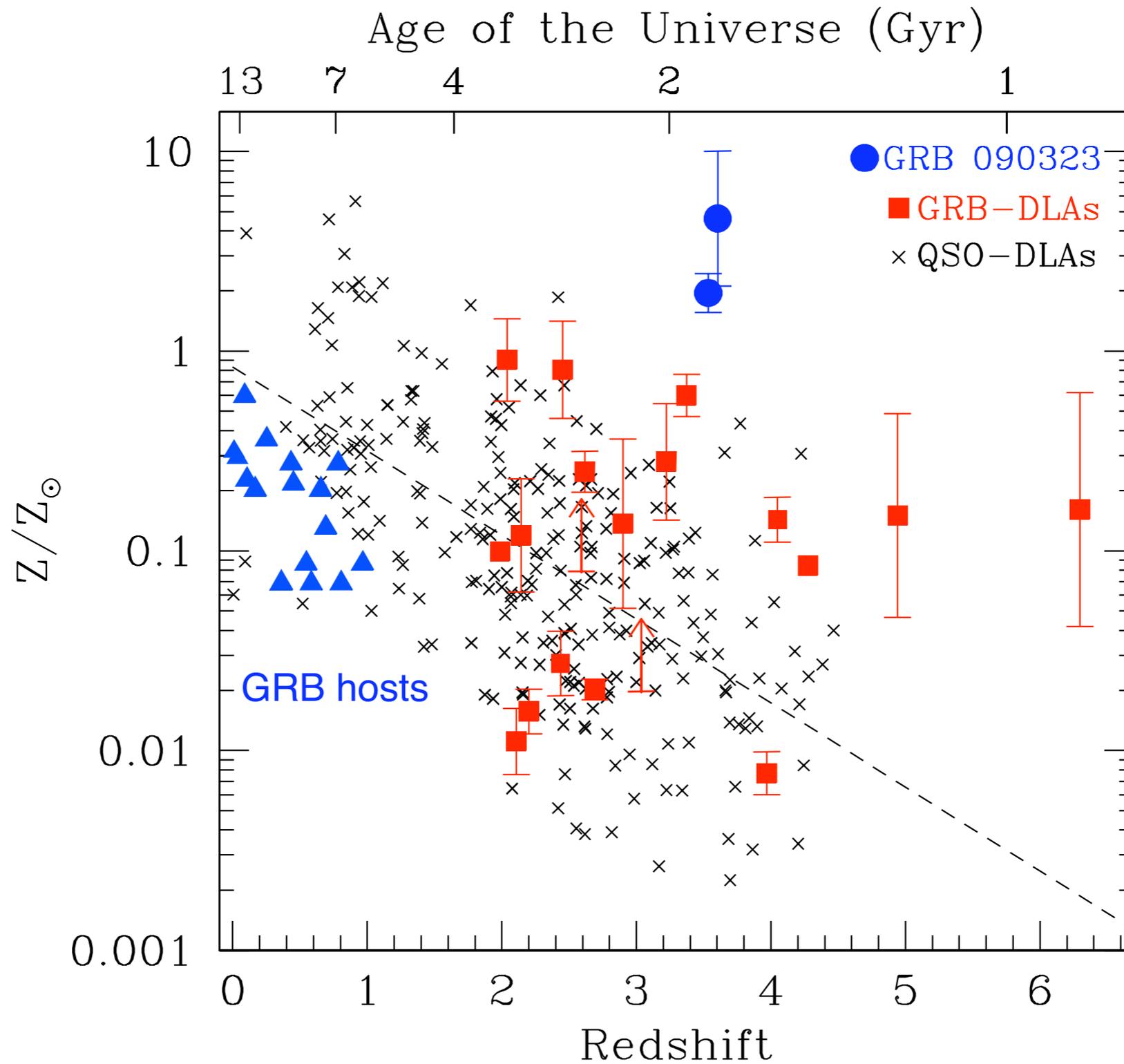
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Cosmic chemical evolution with GRBs

Cosmic chemical enrichment



Savaglio (2006)

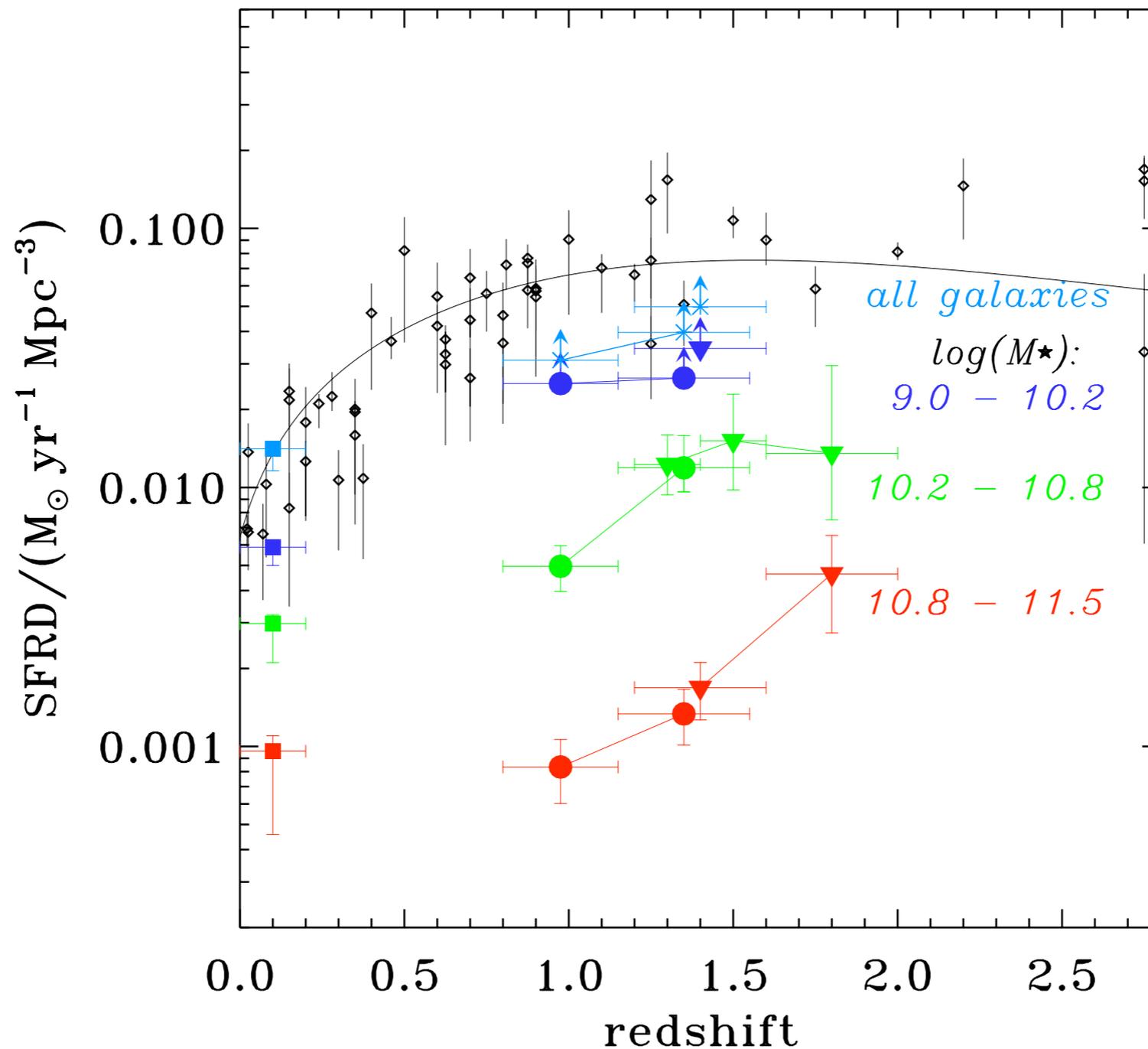
Rau, Savaglio, Krühler, Afonso, Greiner et al. (ApJ, sub.)

Savaglio, Rau, Greiner, Krühler et al. (Science, sub.)

*How can we relate the GRB-host population
to galaxy formation & evolution?*

The Galaxy Evolution context

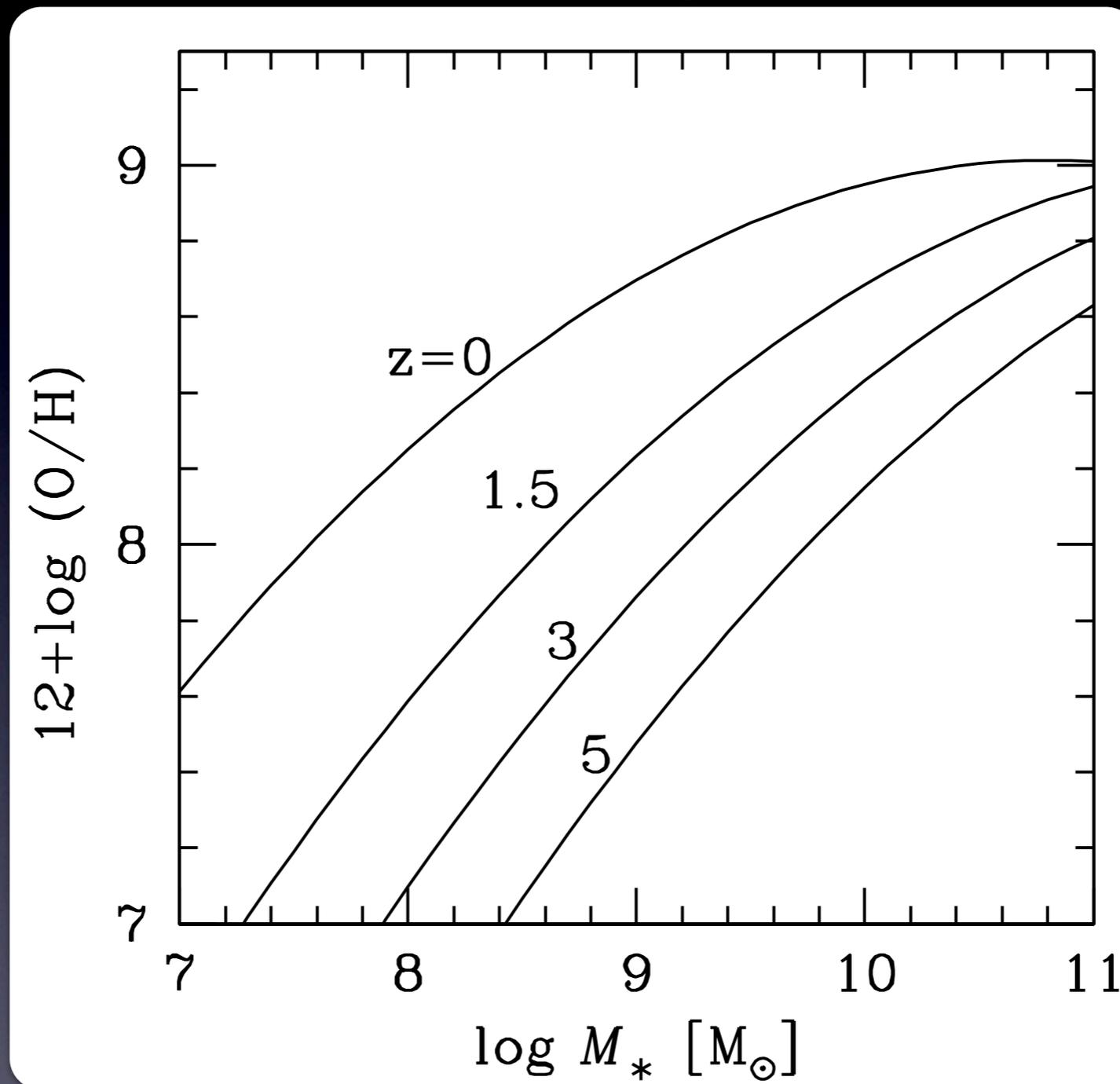
Madau plot per stellar-mass bin



The Galaxy Evolution context

Mass–metallicity relation

Cosmic chemical enrichment



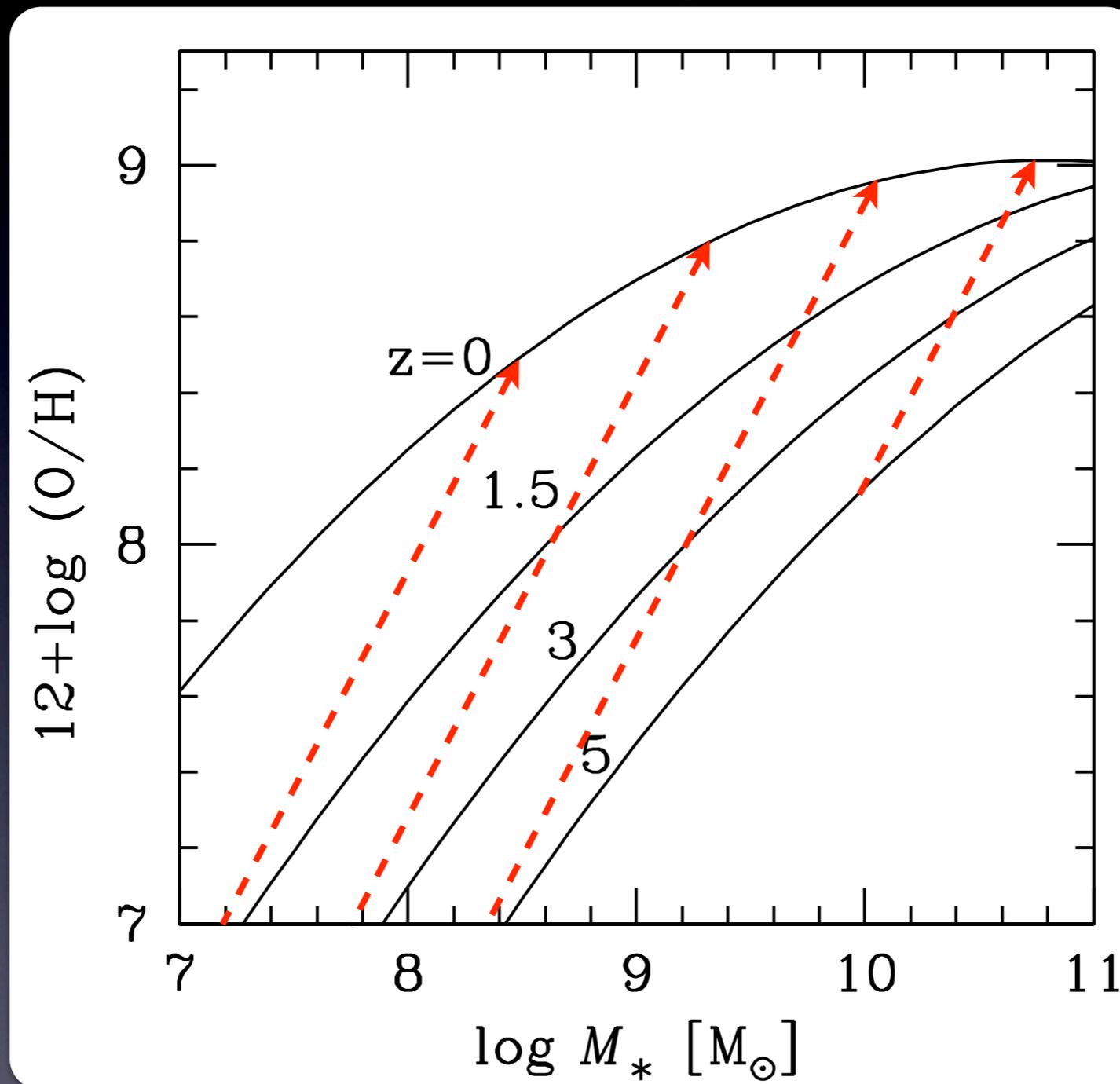
Galaxy stellar mass



The Galaxy Evolution context

Mass–metallicity relation

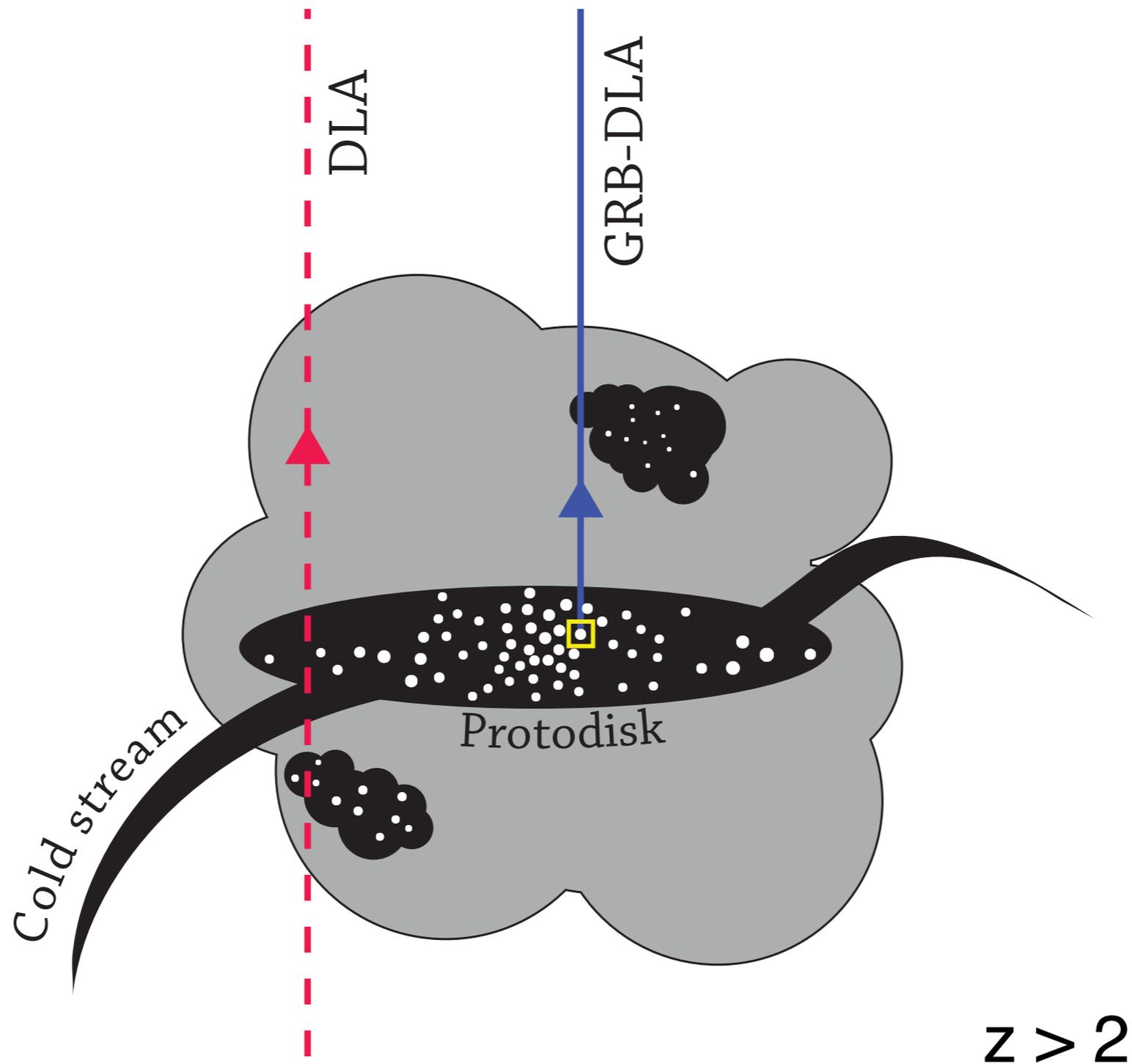
Cosmic chemical enrichment



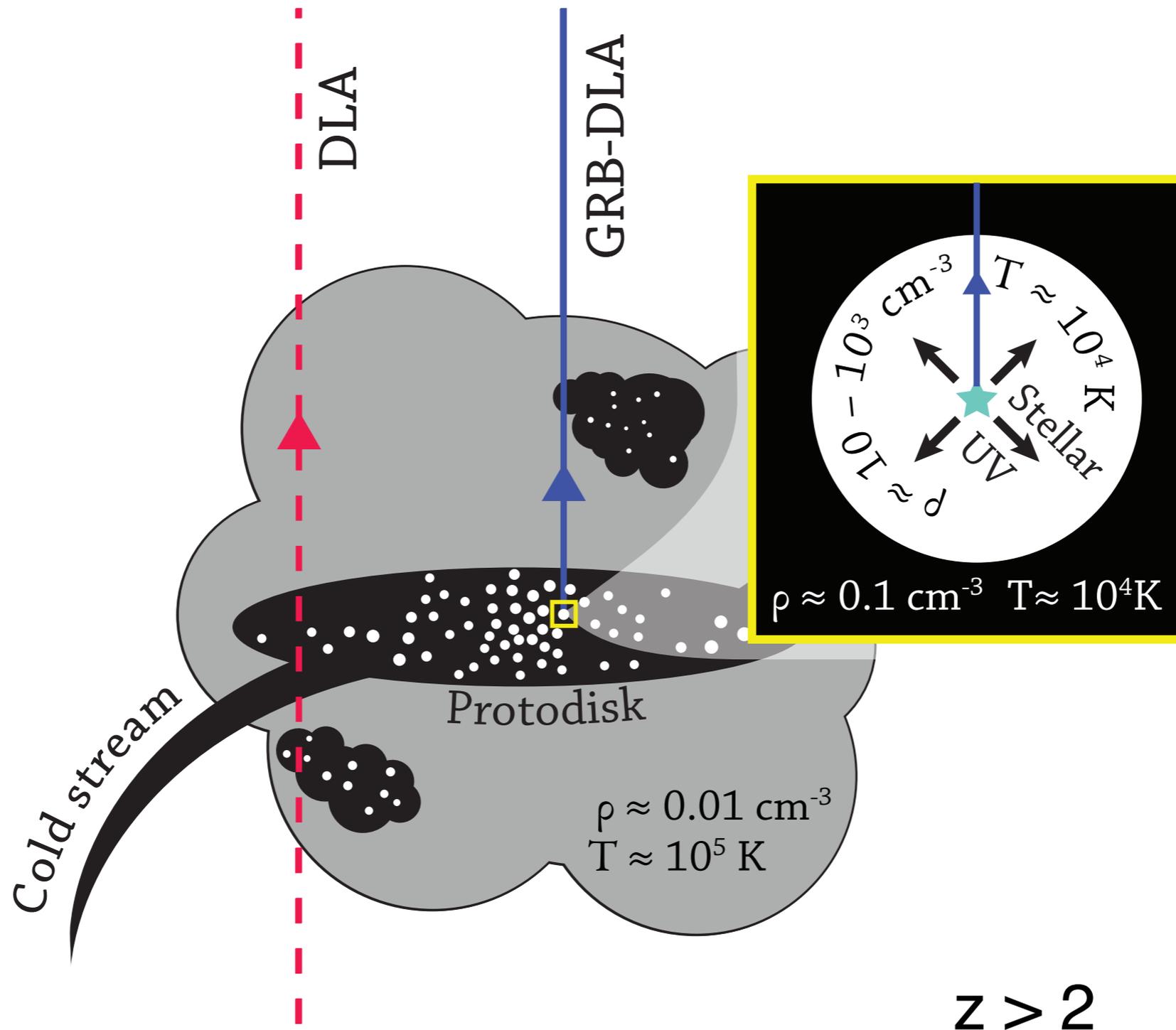
Galaxy stellar mass



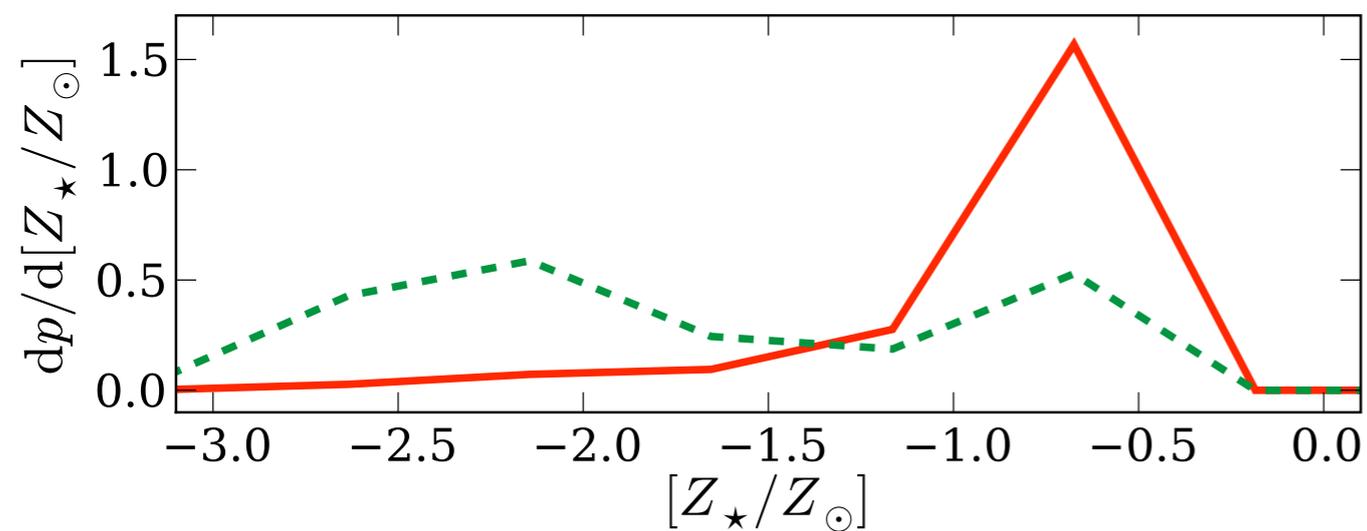
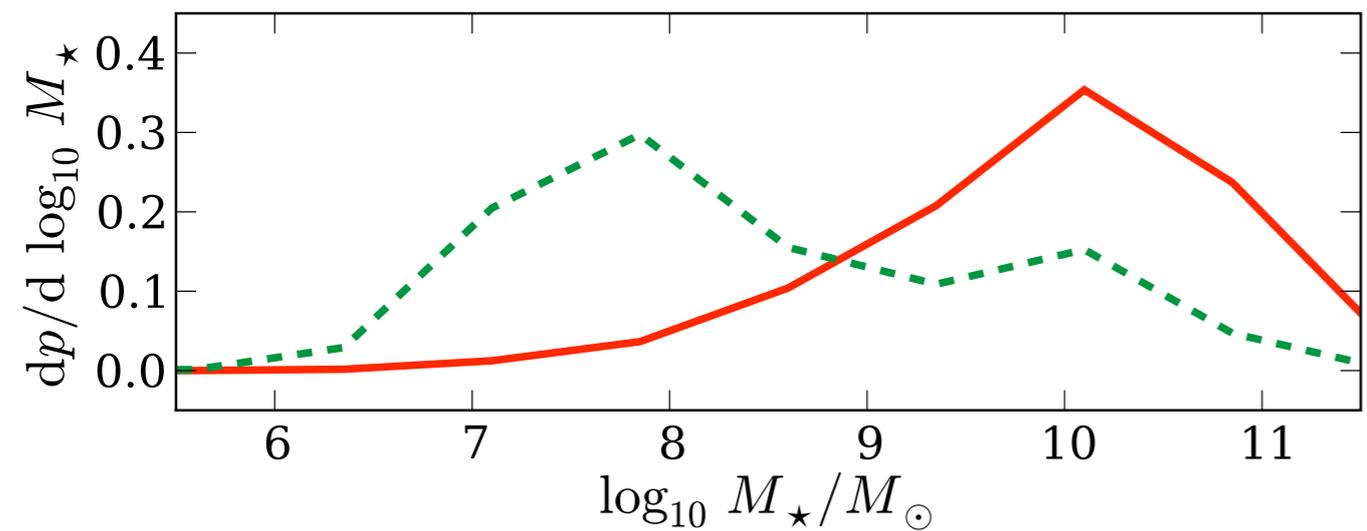
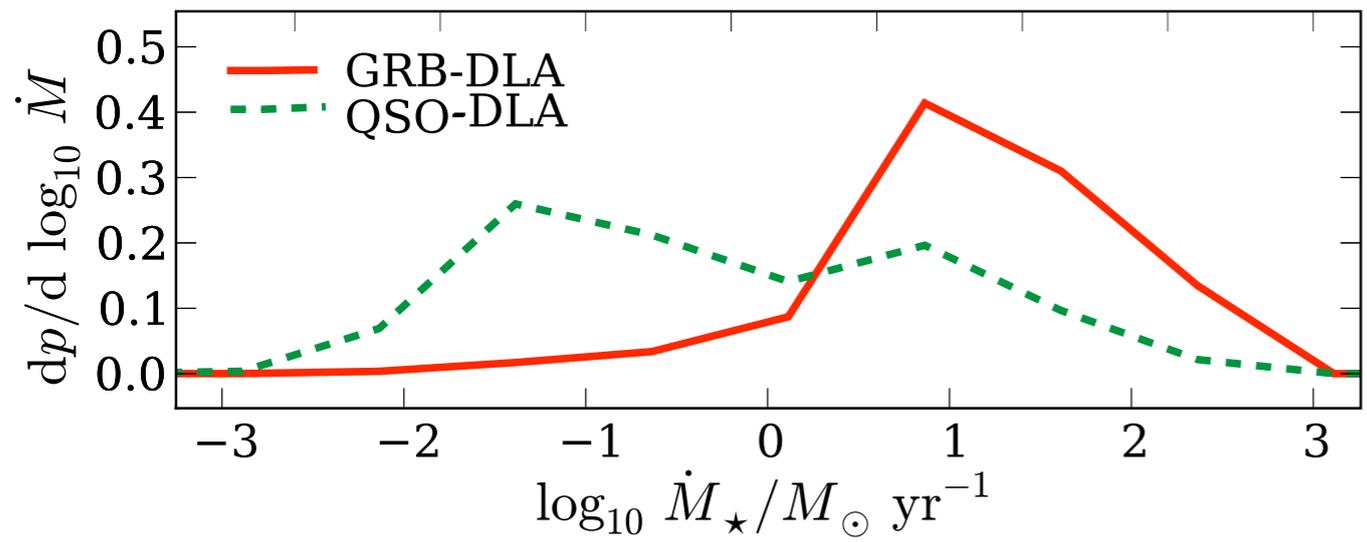
GRB-DLAs vs. QSO-DLAs



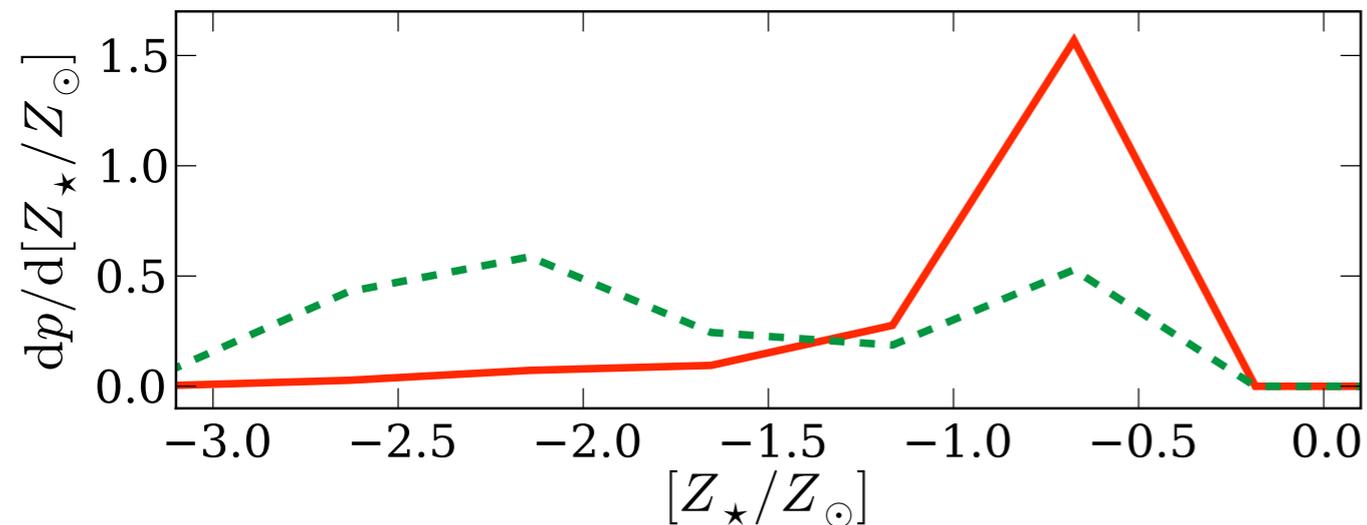
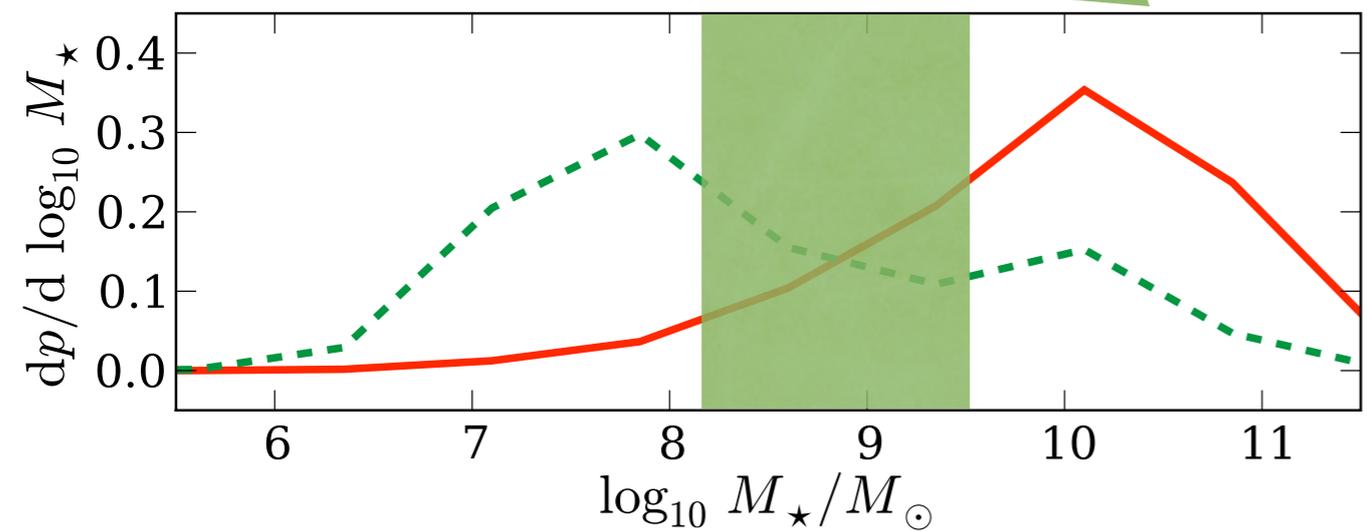
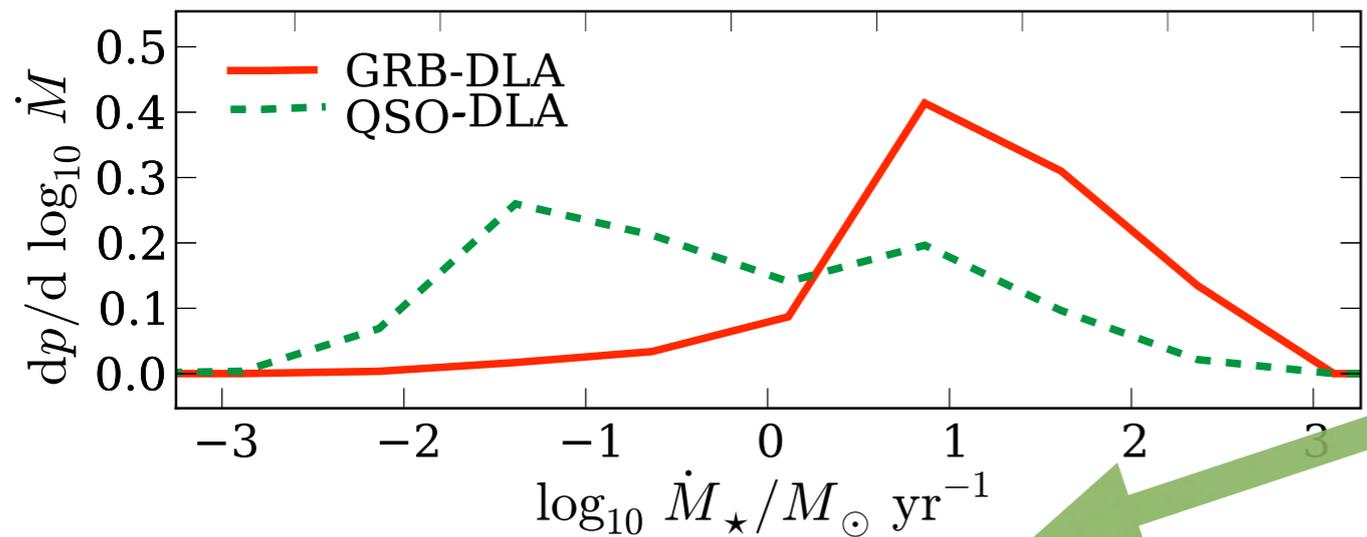
GRB-DLAs vs. QSO-DLAs



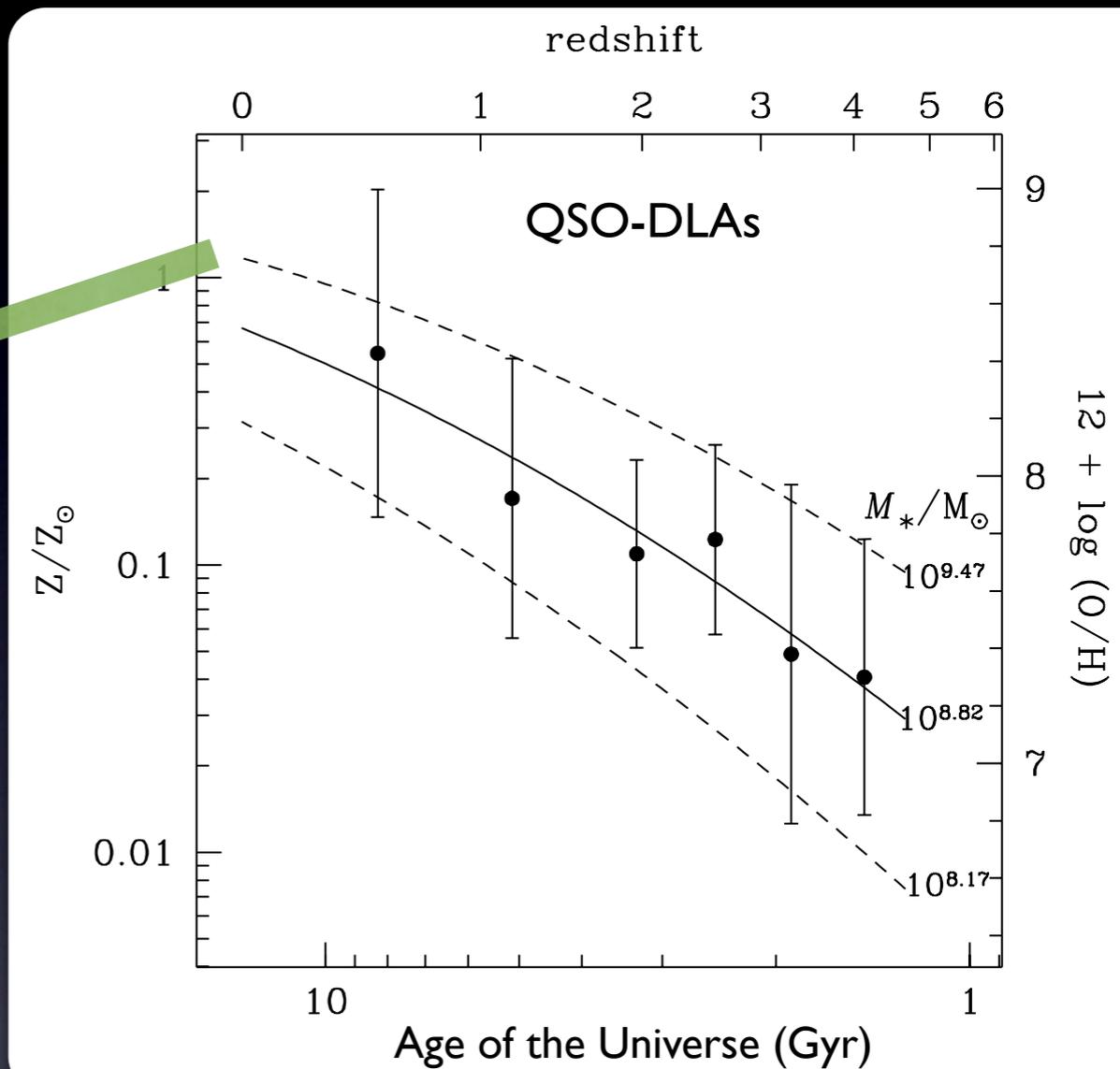
GRB-DLAs vs. QSO-DLAs



GRB-DLAs vs. QSO-DLAs

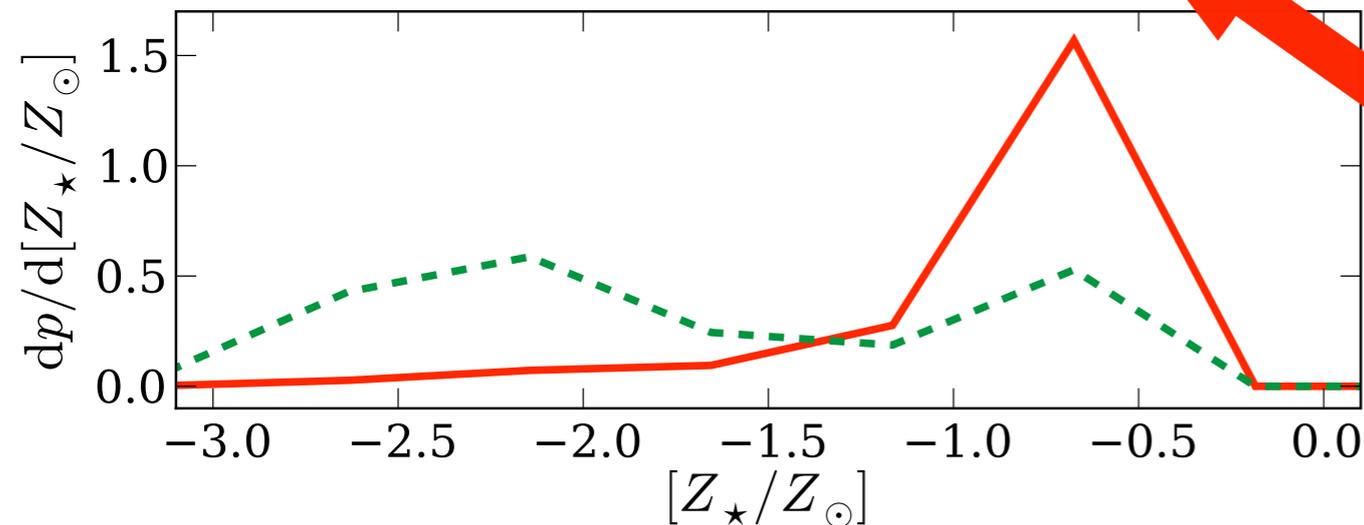
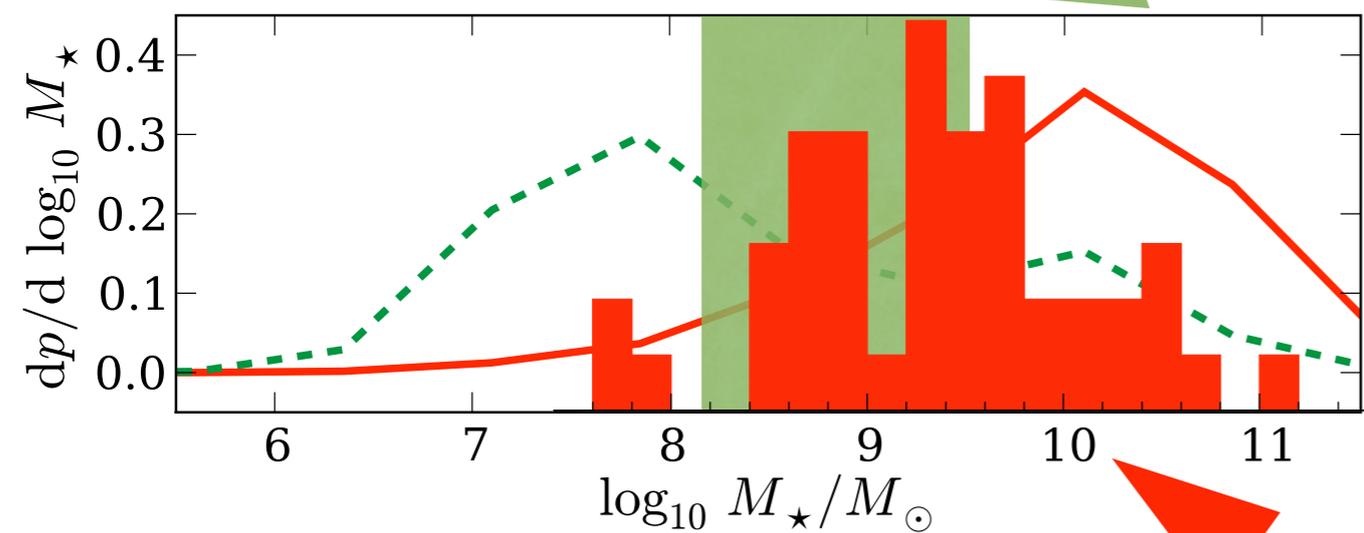
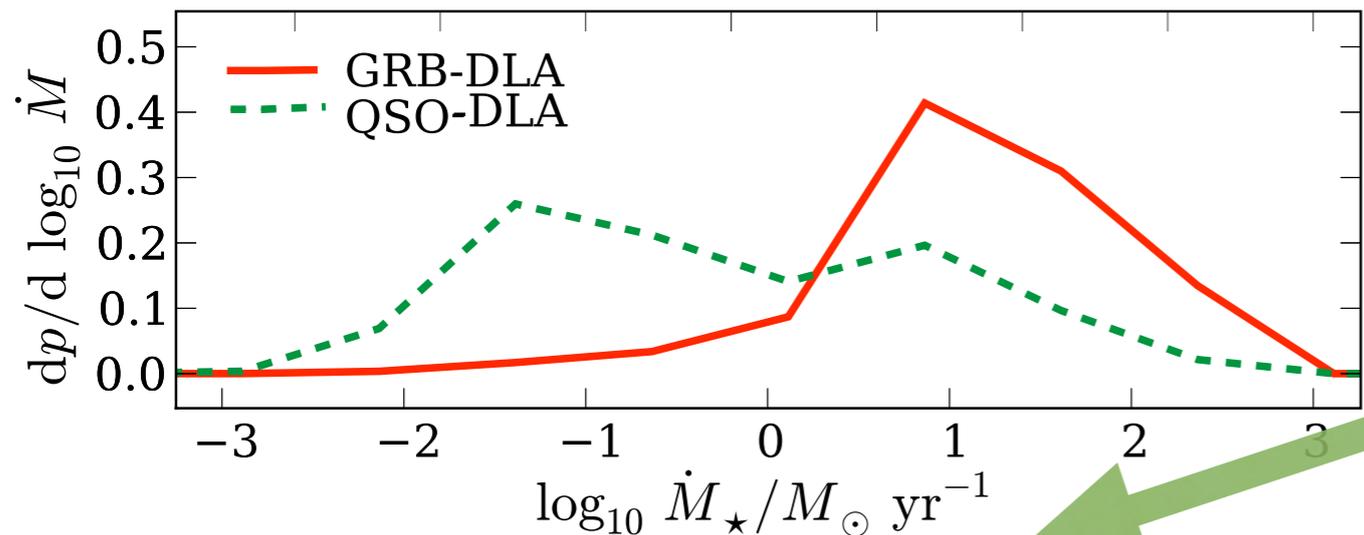


Pontzen et al. (2009)

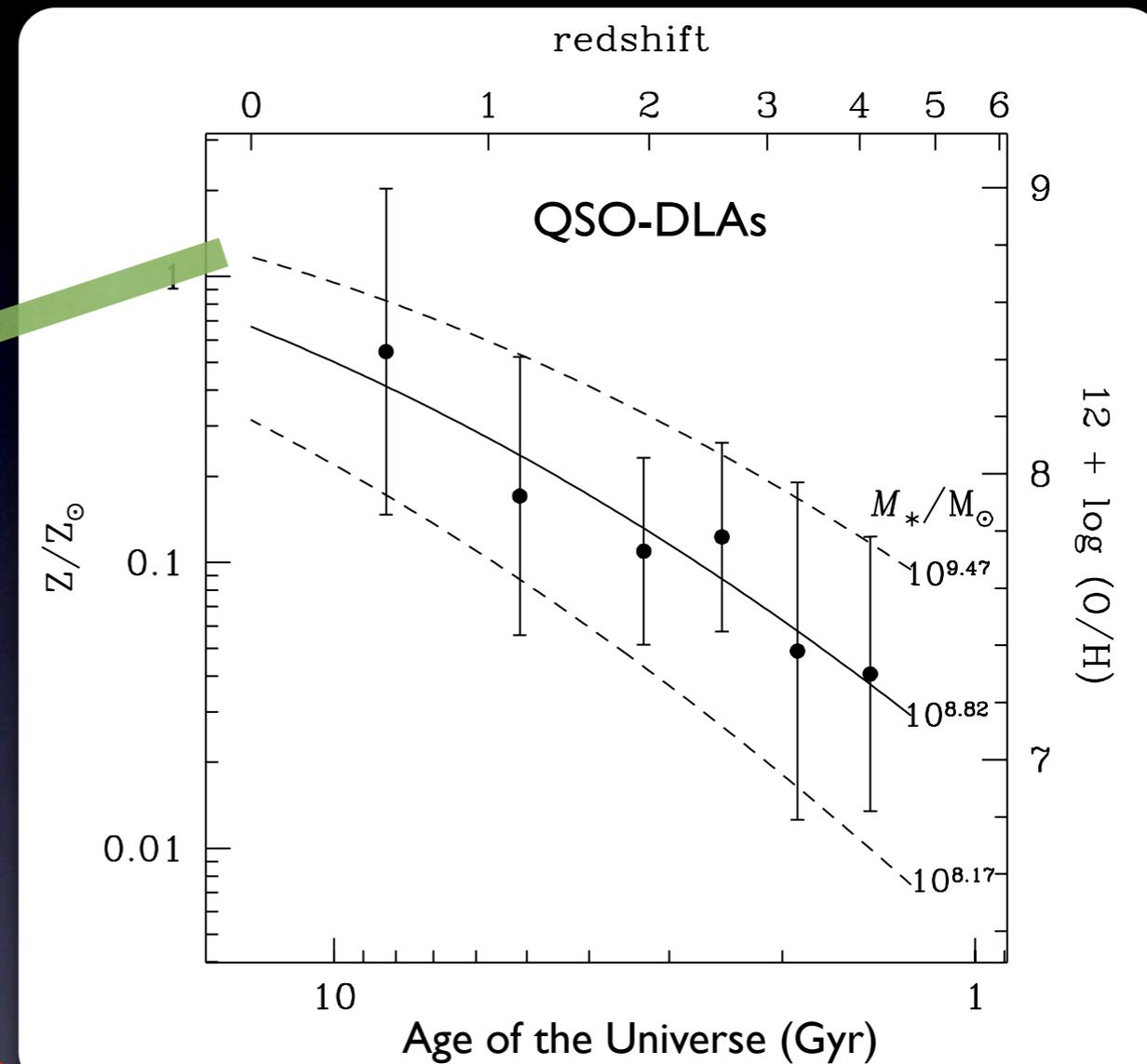


Savaglio et al. (2005)

GRB-DLAs vs. QSO-DLAs



Pontzen et al. (2009)

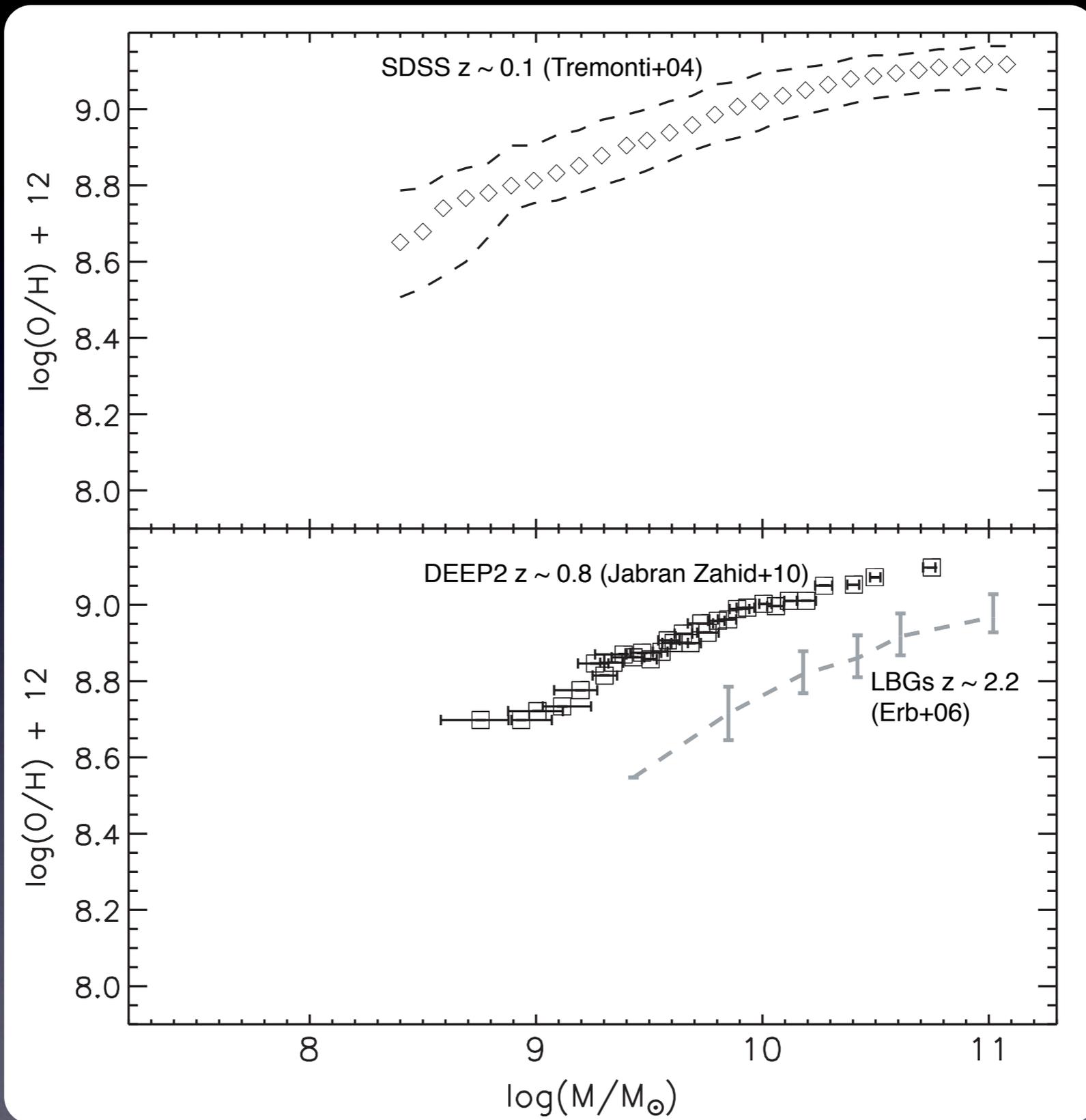


Savaglio et al. (2005)

GRB hosts $0 < z < 2.4$
(Savaglio, Glazebrook, Le Borgne 2009)

Mass-Metallicity relation

Metallicity

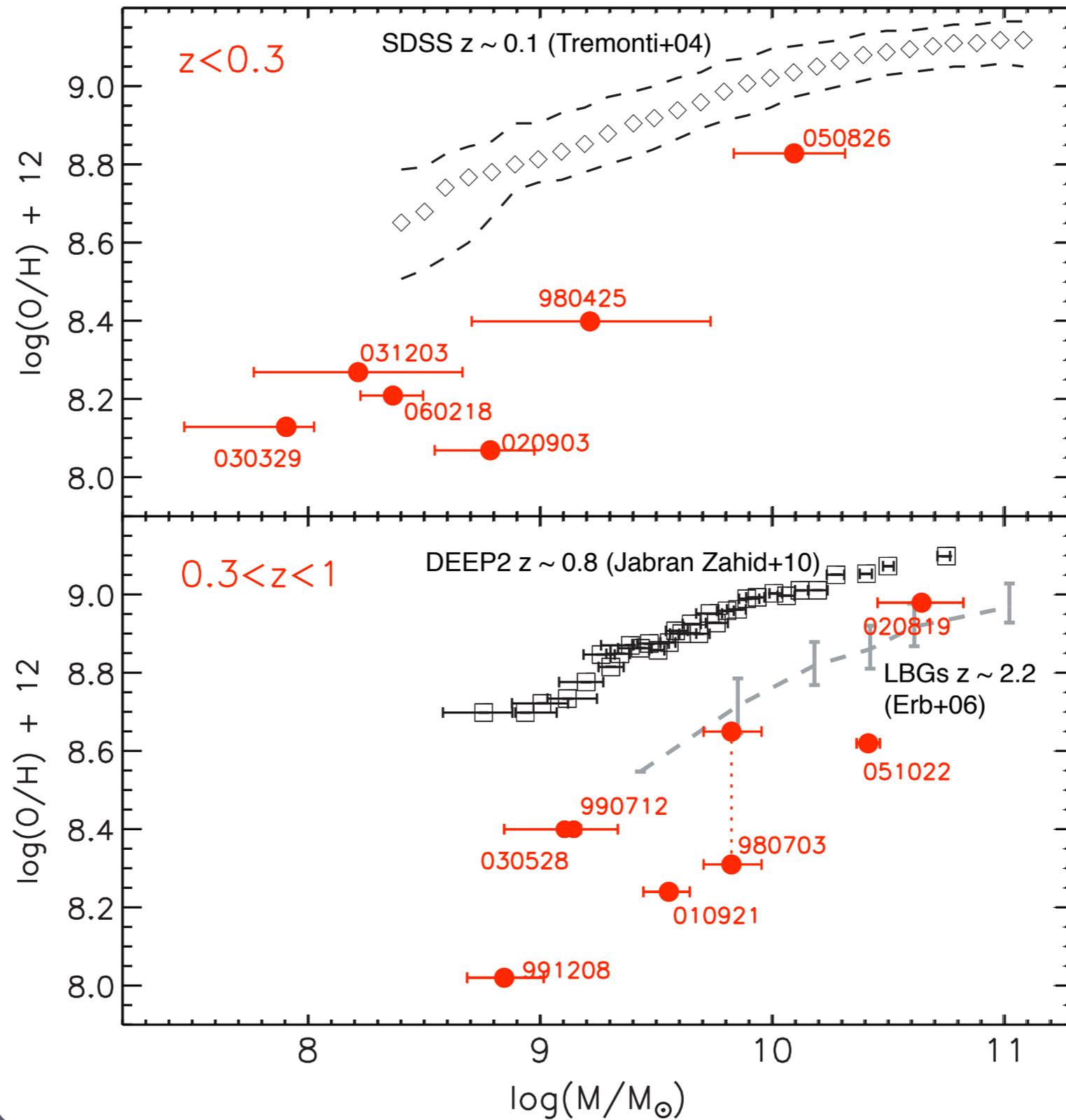


Galaxy stellar mass



Mass-Metallicity relation

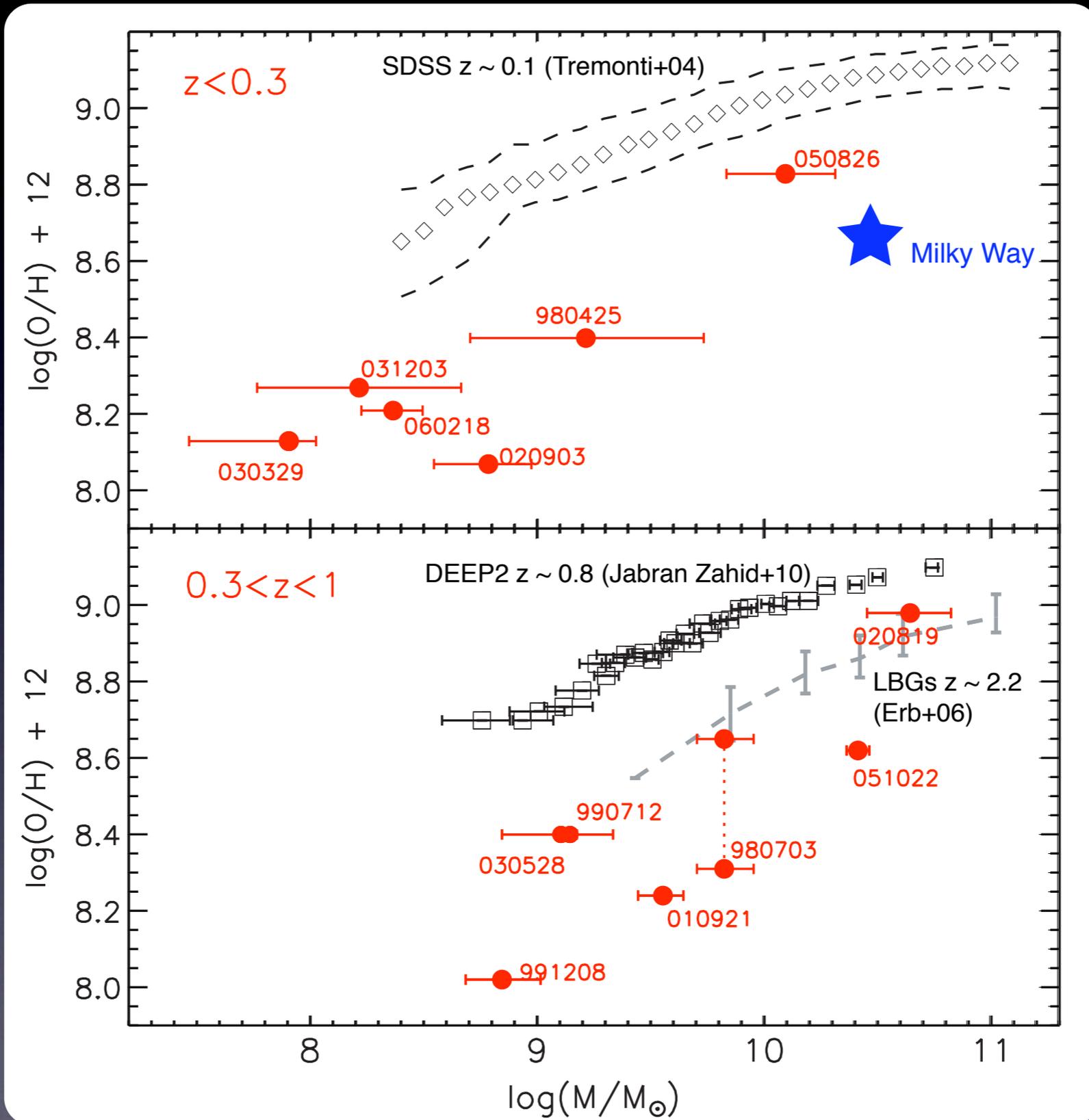
Metallicity



Galaxy stellar mass

Mass-Metallicity relation

Metallicity



Galaxy stellar mass



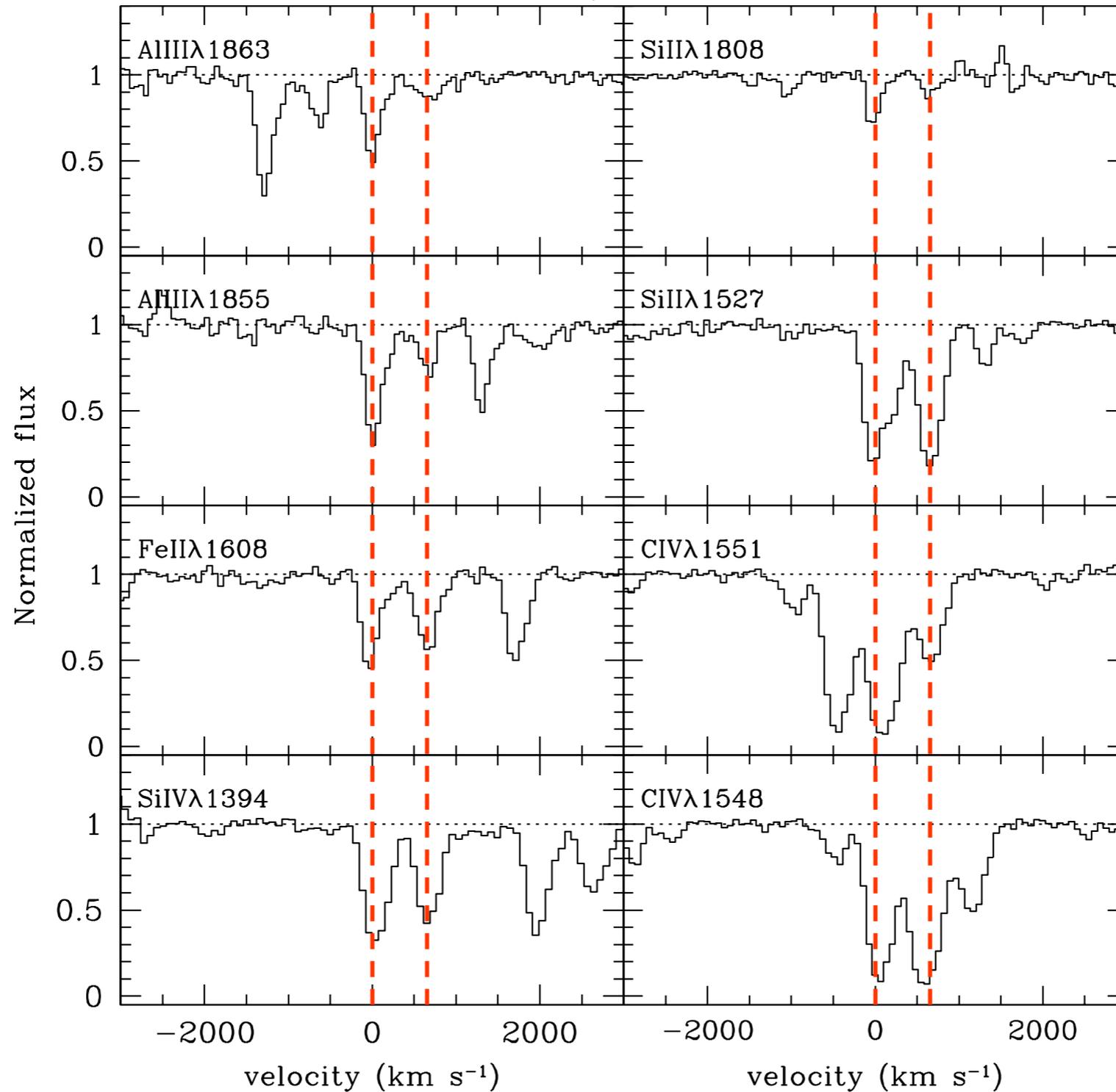
Related publications:

- Campisi, M. A., De Lucia, G., Li, L., Mao, S., & Kang, X. 2009, MNRAS, 400, 1613
- Chen, H.-W. et al. 2009, ApJ, 691, 152
- Chisari, N. E., Tissera, P. B., & Pellizza, L. J. 2010, MNRAS, submitted arXiv:1005.4036
- Fynbo, J. P. U.; Prochaska J. X.; Sommer-Larsen J.; Dessauges-Zavadsky, M., Møller P. 2008, ApJ, 683, 321
- Han, X. H., Hammer, F., Liang, Y. C., Flores, H., Rodrigues, M., Hou, J. L., Wei, J. Y. 2010, A&A, 514, 24
- Kocevski, D., West, A. A., & Modjaz, M. 2009, ApJ, 702, 377
- Kretchmer, C., & Ravindranath, S. 2007, ApJ, 654, 172
- Levesque, E. M., Kewley, L. J., Graham, J. F., & Fruchter, A. S. 2010, ApJ, 712, L26
- Modjaz, M., et al. 2008, AJ, 135, 1136
- Niino, Y., et al. 2010, submitted, arXiv:1006.5033

*Is there a connection between
sub-mm galaxies and GRB hosts ?*

Double absorbers in GRB afterglows

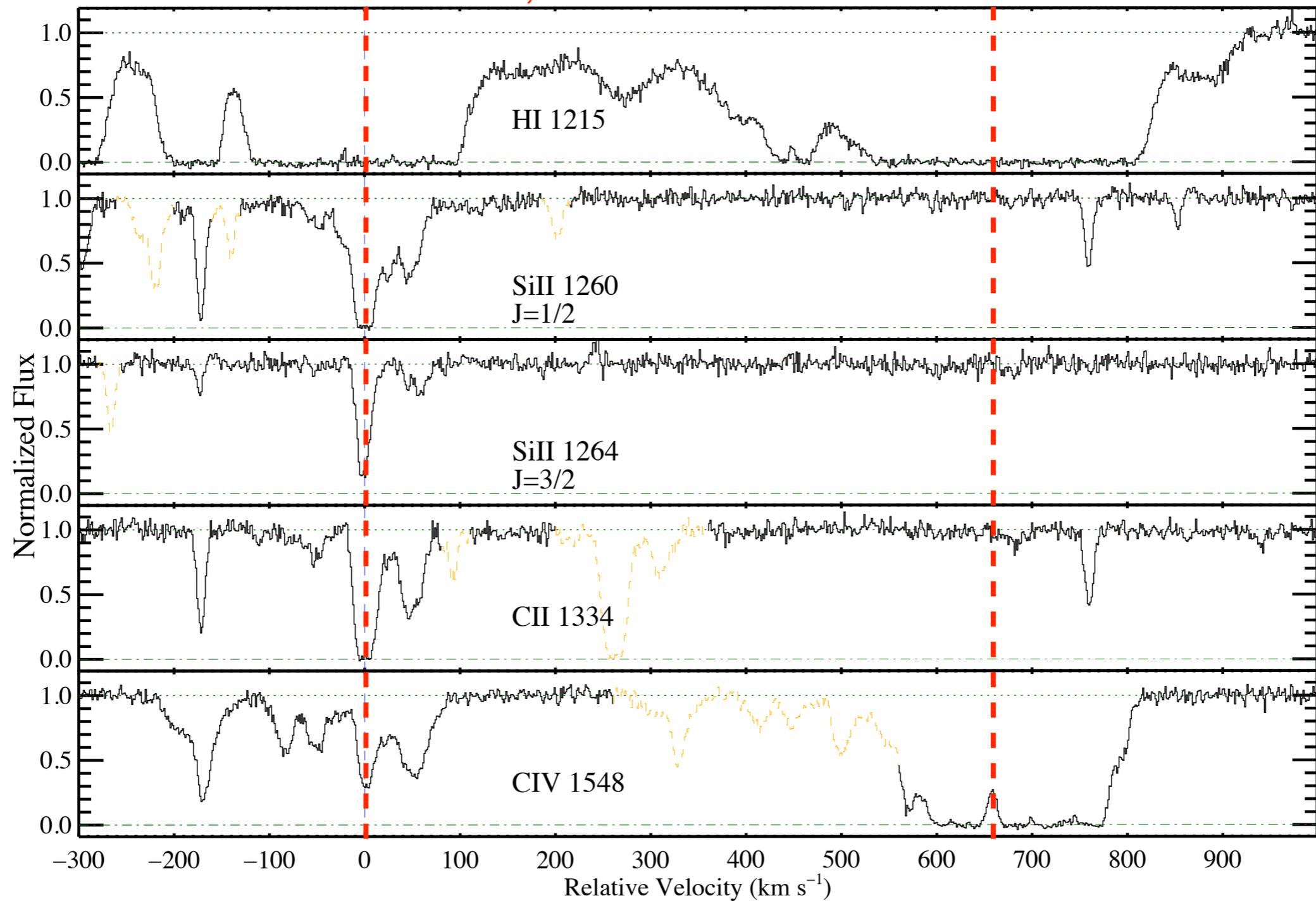
GRB 090323 $z = 3.567, 3.577$



Double absorbers in GRB afterglows

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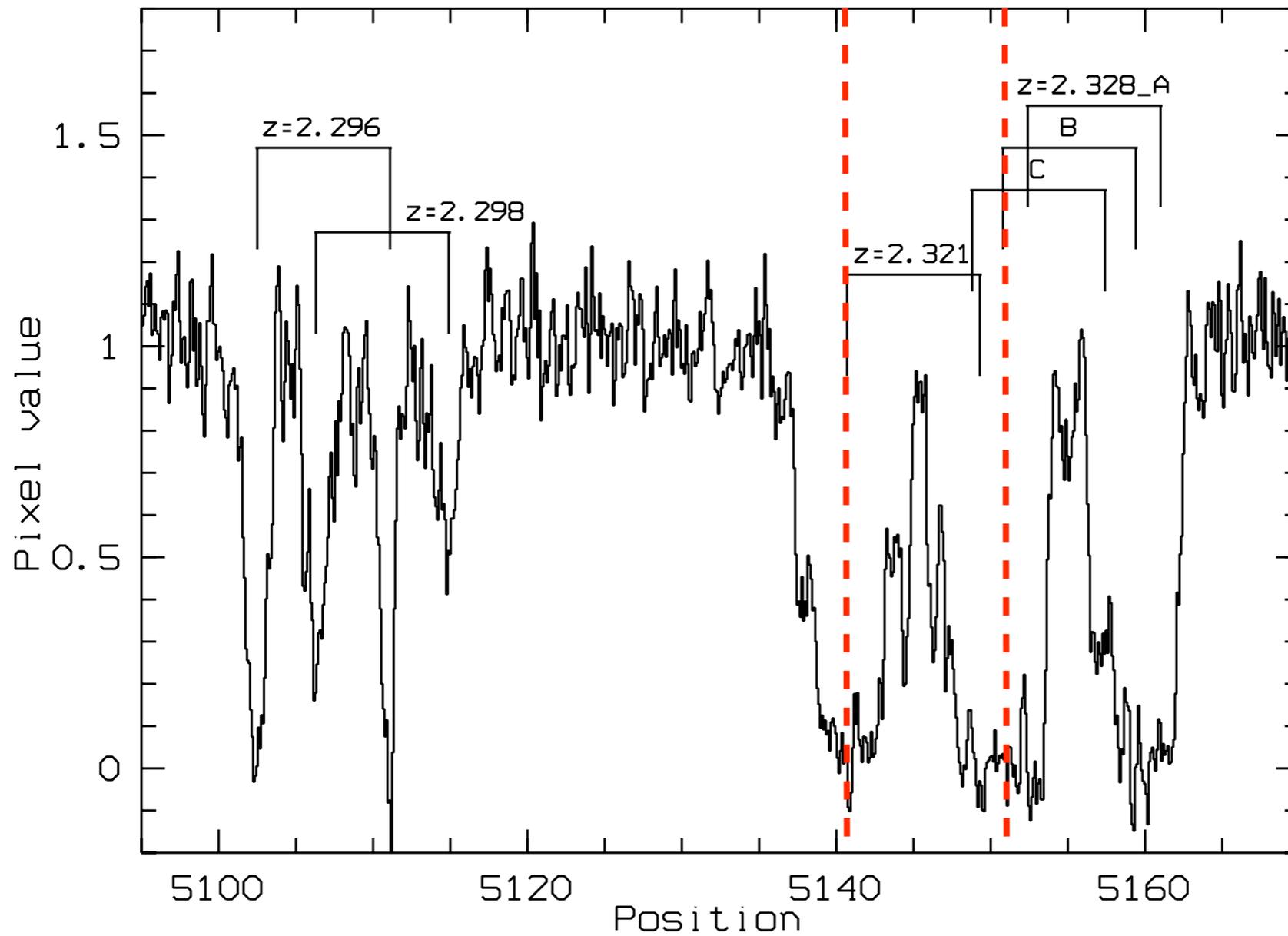
GRB 080810 $z = 3.355, 3.365$



Double absorbers in GRB afterglows

Double absorbers in GRB afterglows

GRB 021004 $z = 2.321, 2.328$



Sub-millimeter Galaxies



CHANDRA X-RAY

Alexander et al. (2005)

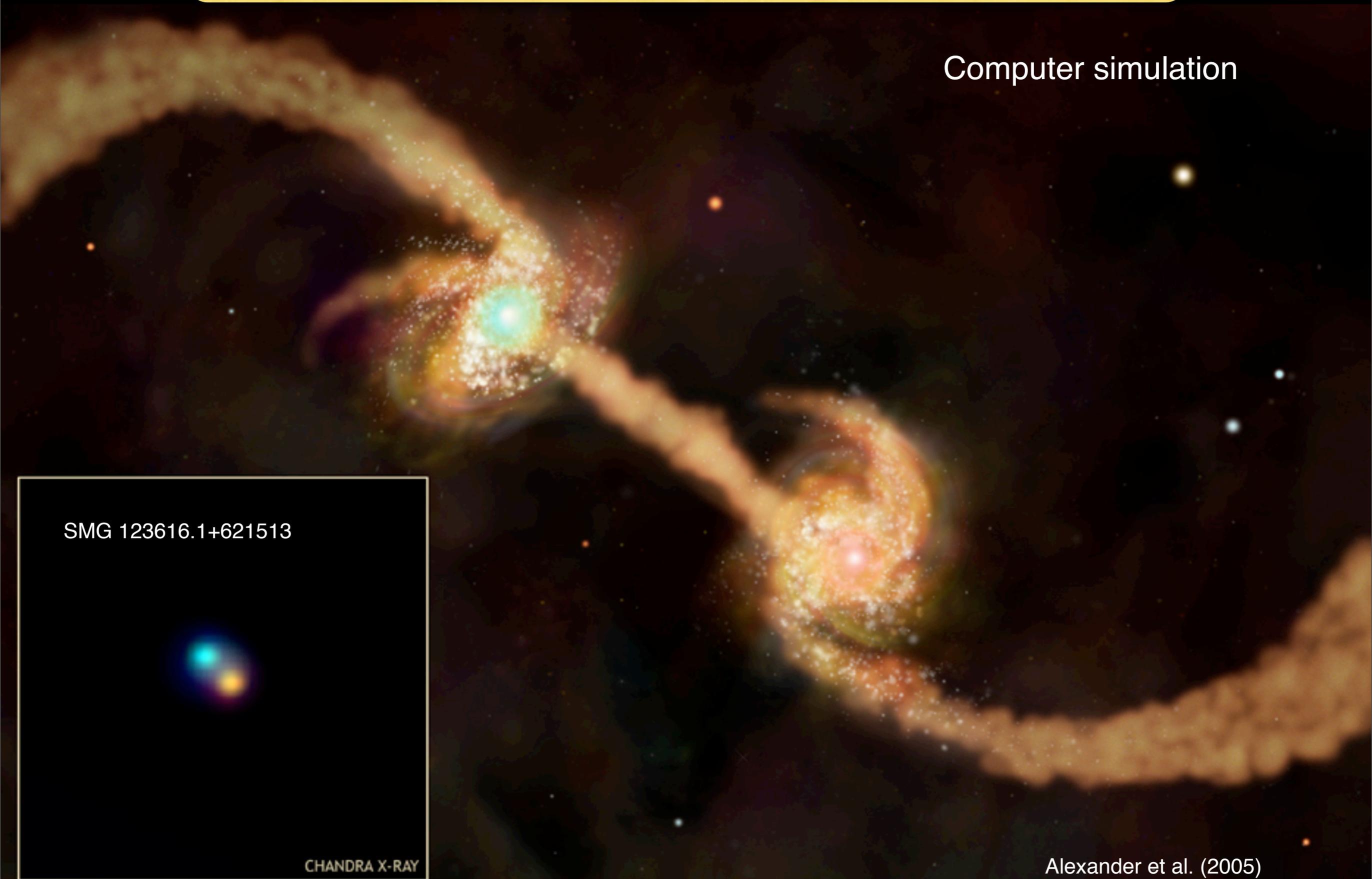
Sub-millimeter Galaxies

Computer simulation

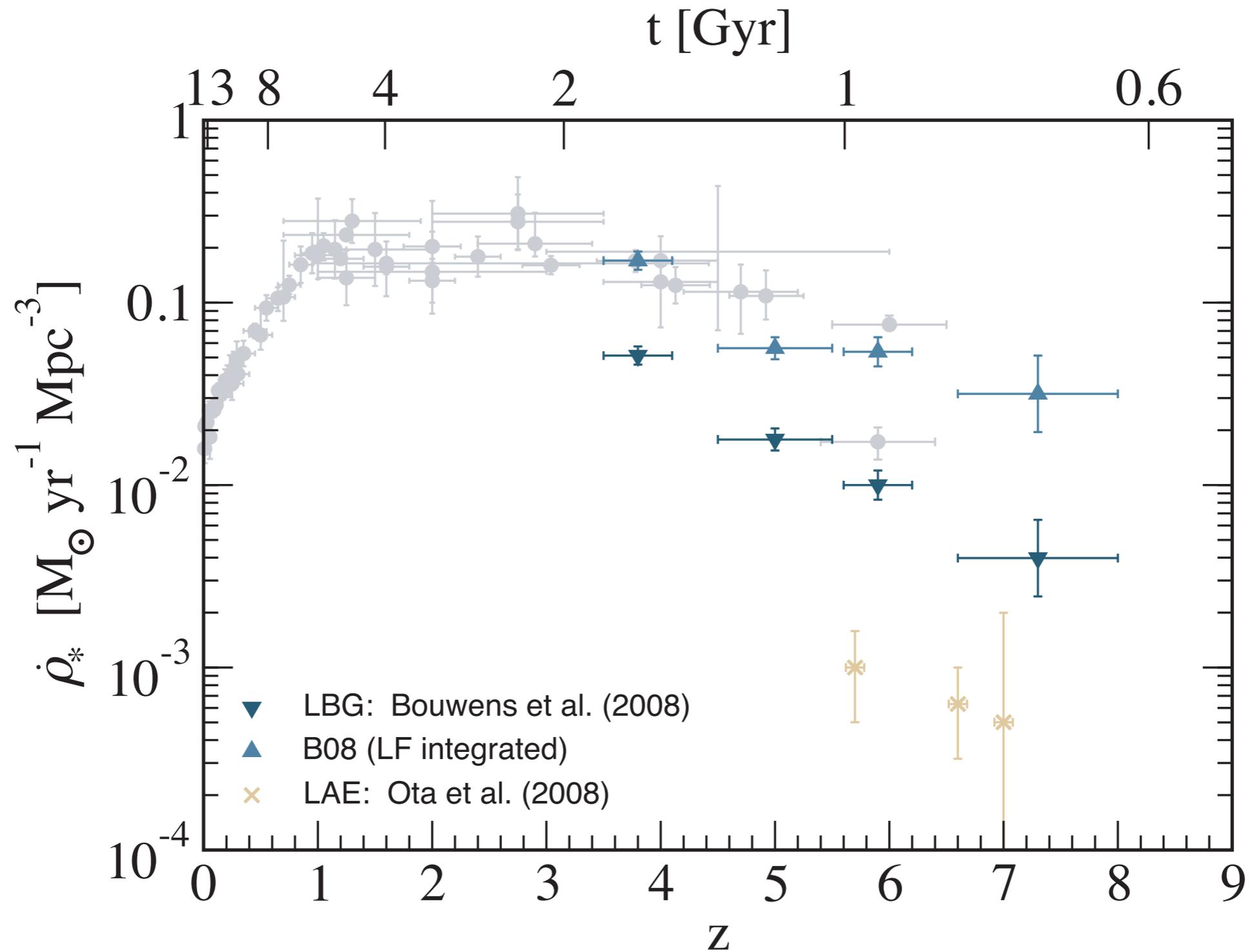
SMG 123616.1+621513

CHANDRA X-RAY

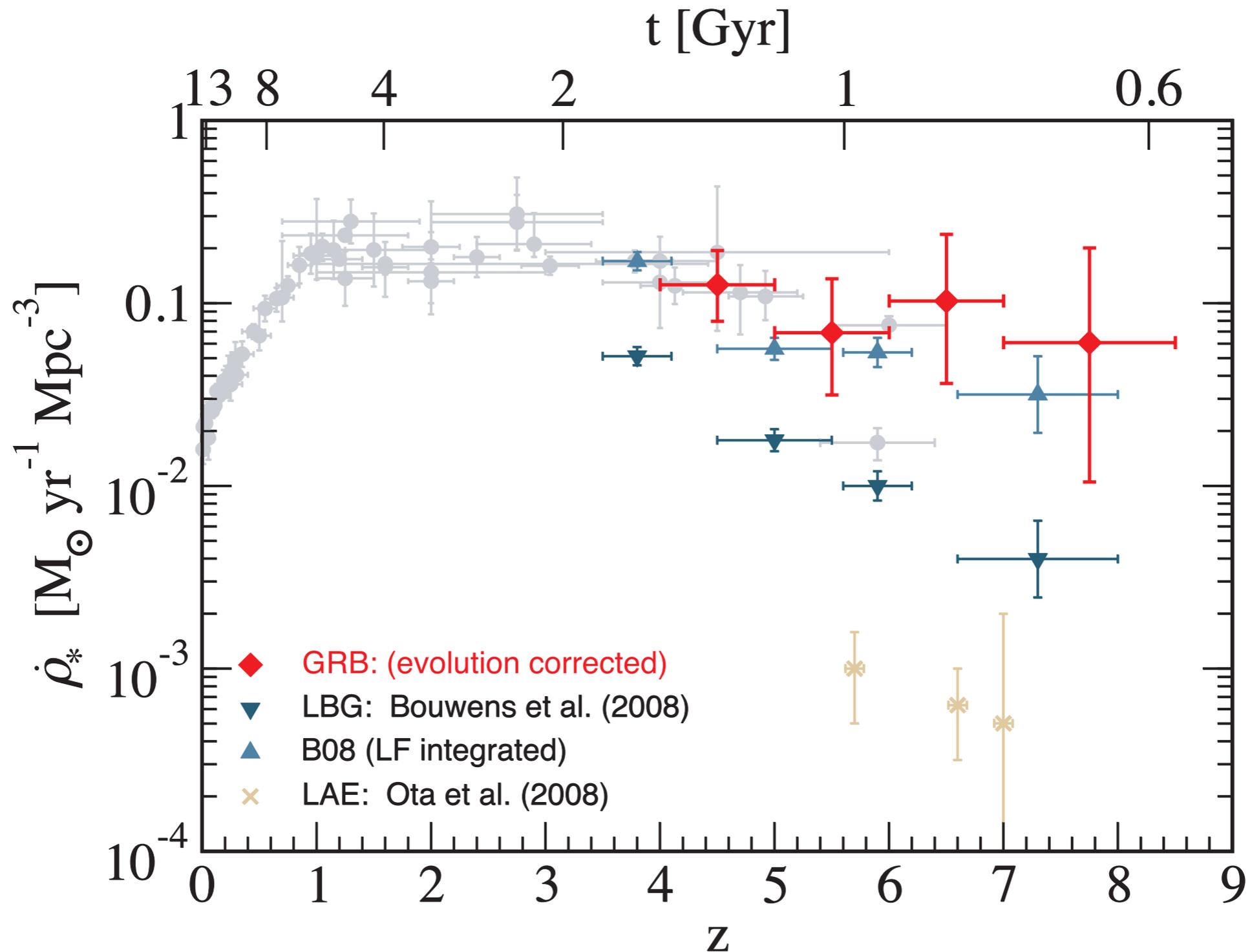
Alexander et al. (2005)



Star Formation Rate Density of the Universe



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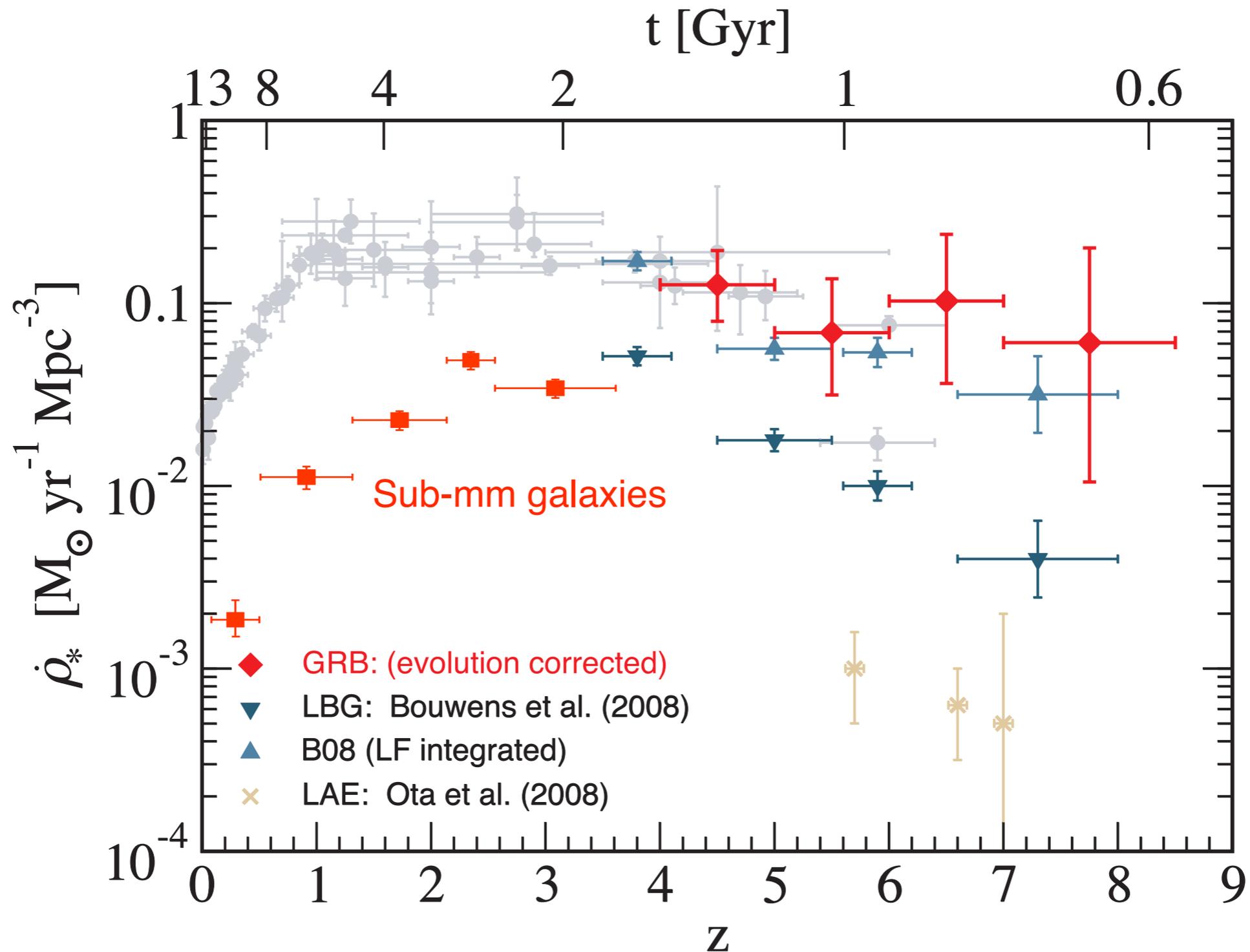


Chary, Berger, & Cowie (2007)

Yüksel, Kistler, Beacom, & Hopkins (2008)

Kistler et al. (2009)

Star Formation Rate Density of the Universe



Chary, Berger, & Cowie (2007)
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 Kistler et al. (2009)

Michałowski, Hjorth & Watson (2009)

Conclusions

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