

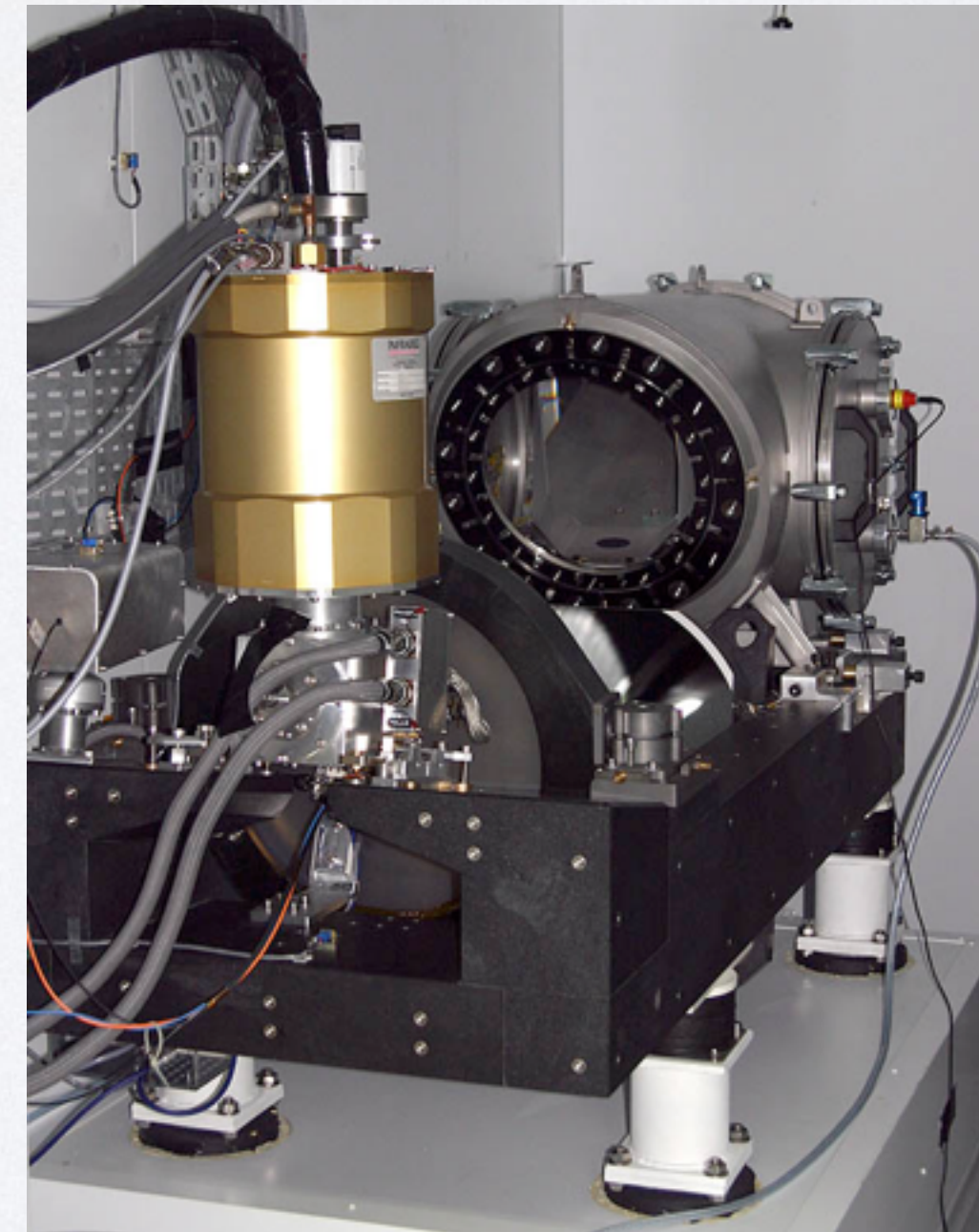
THE HIGH PRECISION SEARCH FOR NORTHERN NEPTUNES AND SUPER- EARTHS WITH *SOPHIE*

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THE SOPHIE SPECTROGRAPH

- Environmentally stabilized Echelle spectrograph (387-694 nm)
- Simultaneous Thorium-Argon reference, $R = 75\,000$
- Installed on the 1.93 m telescope at the Observatoire de Haute Provence (OHP)
- Active since 2006 (replacement of ELODIE), octogonal fibers since 2011 : **2 m/s precision**

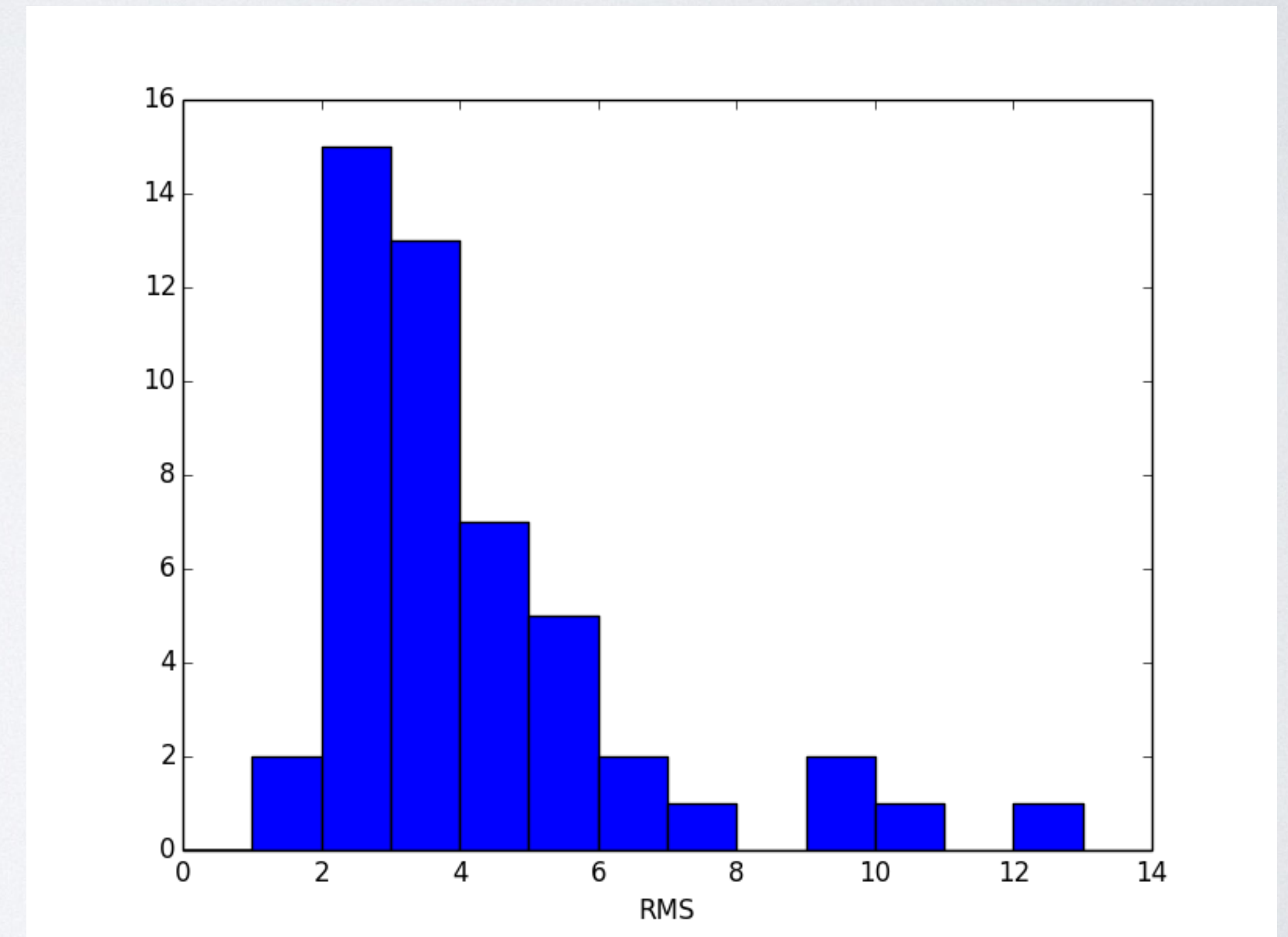


THE HIGH PRECISION SURVEY

PI : F. Bouchy

- Started in june 2011, ~50 nights/year
- 190 bright stars ($V < 10$) :
 - G & K spectral type and $0.6 < B-V < 1.4$ and volume limited to 35 pc
 - Slow rotators, non active stars
 - Single stars, without any already known companion
- Not part of HARPS-N GTO survey

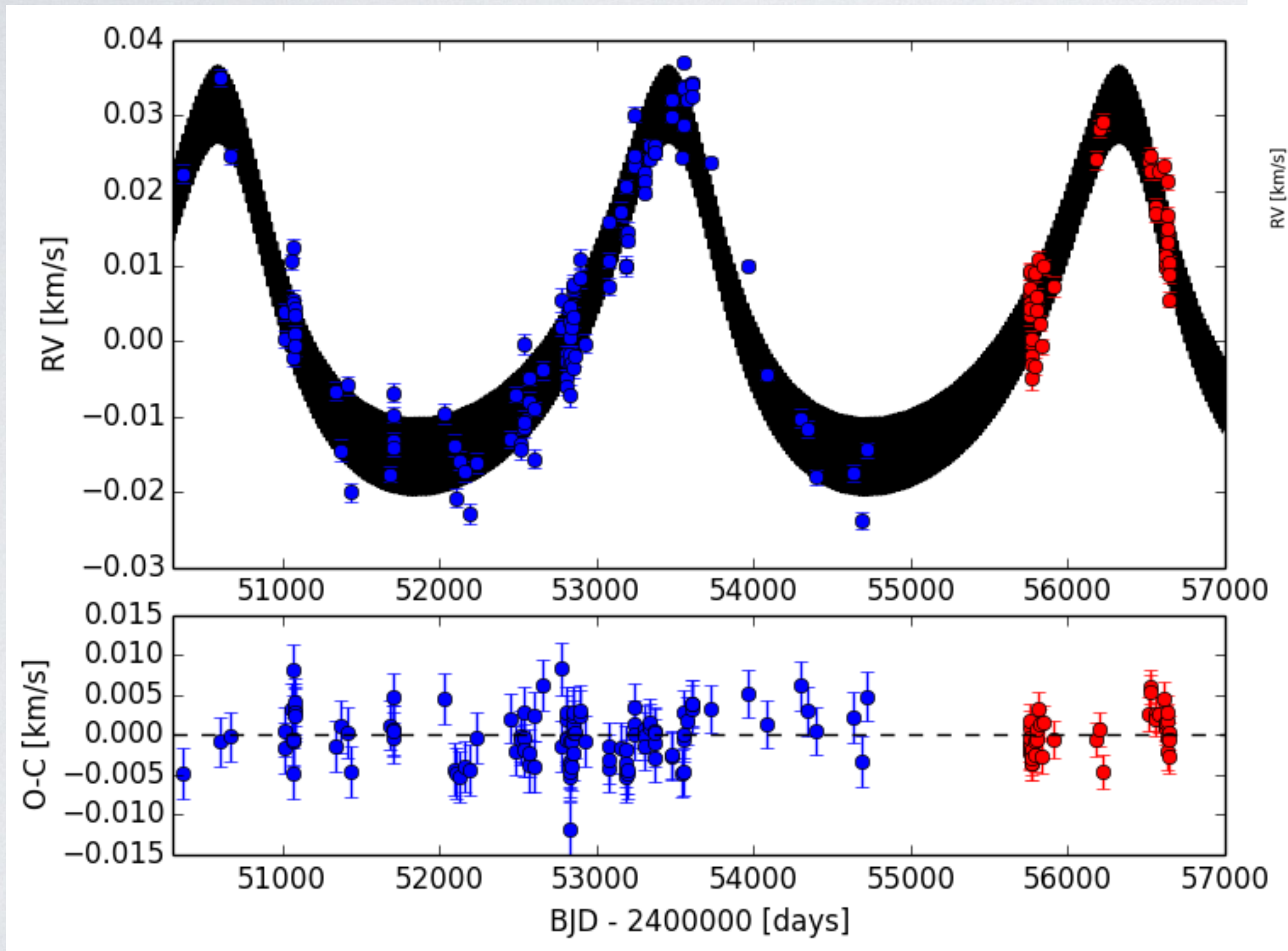
RMS distribution of targets with at least 20 measurements



VALIDATING THE UPGRADED SOPHIE PRECISION

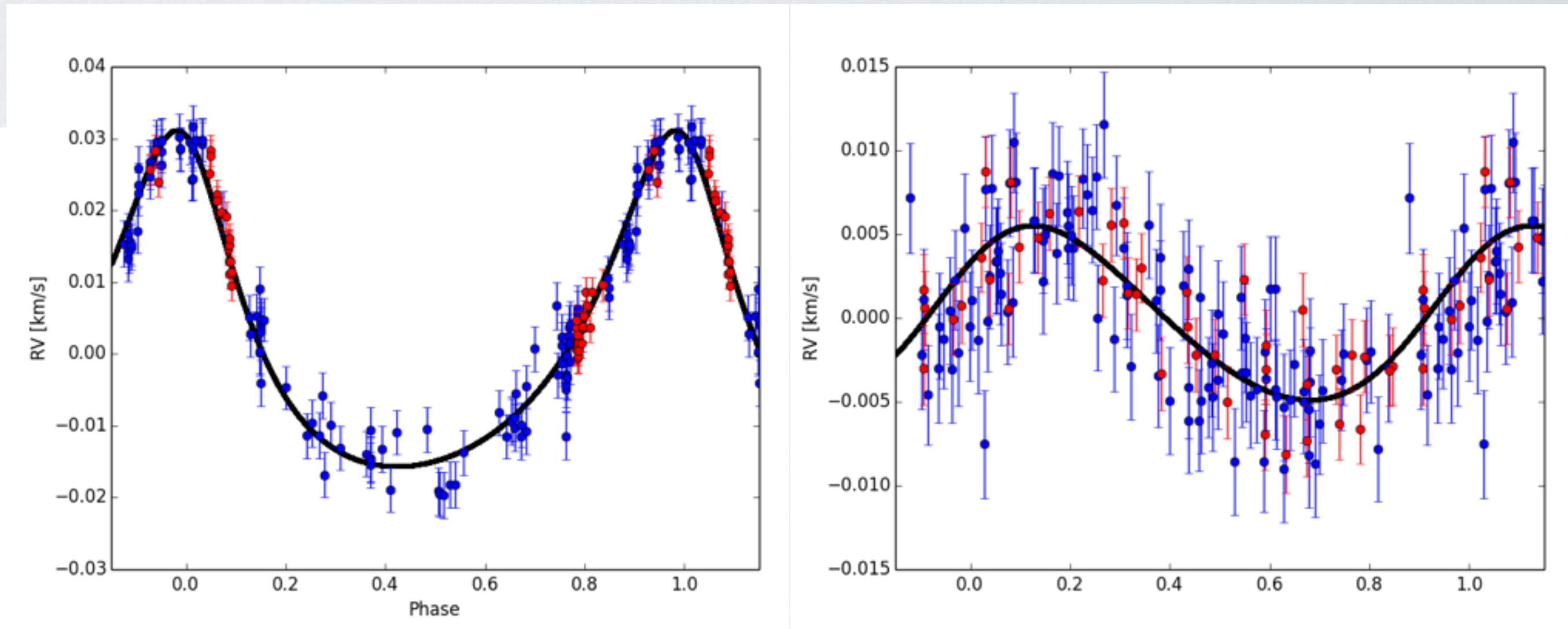
Courcol, B., Bouchy, F., et al. 2015, A&A, in press (arXiv 1506.07144)

HD 190360 : Two-planets system



HIRES
 $\sigma = 3.3 \text{ m/s}$

SOPHIE
 $\sigma = 2.4 \text{ m/s}$



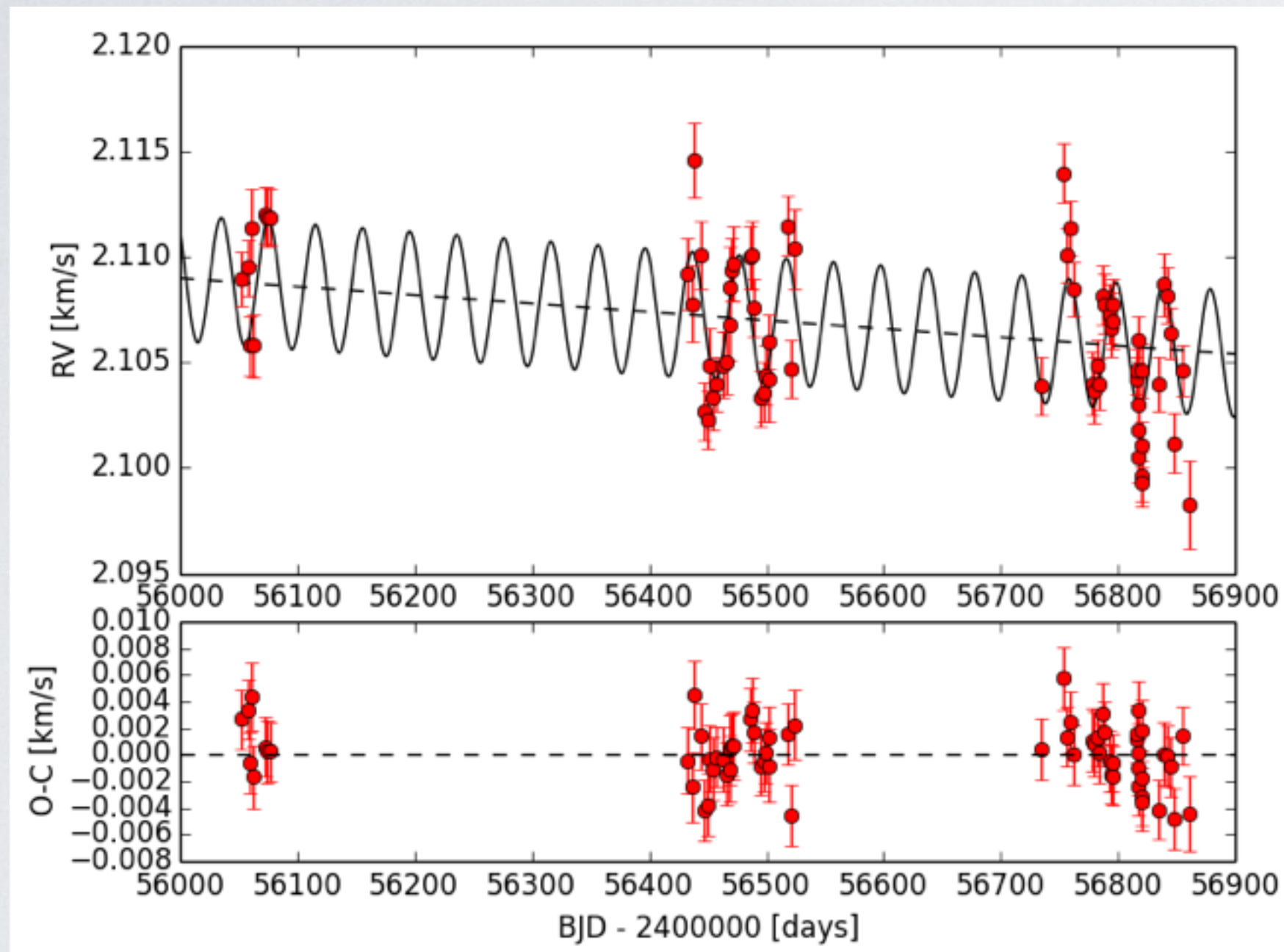
Phase folded RVs for HD 190360 b (left) and c (right)

Parameter	HD190360b	HD190360c
P [days]	2867.9 ± 7.7	17.1186 ± 0.0016
T_0 [BJD]	59271 ± 19	$55570.3^{+1.5}_{-2.9}$
e	0.343 ± 0.017	0.107 ± 0.07
ω [deg]	14.7 ± 32	305.8^{+39}_{-280}
K [m s^{-1}]	23.39 ± 0.46	5.20 ± 0.37
$m \sin i [M_{\oplus}]$	475.16 ± 49.0	20.28 ± 3.16

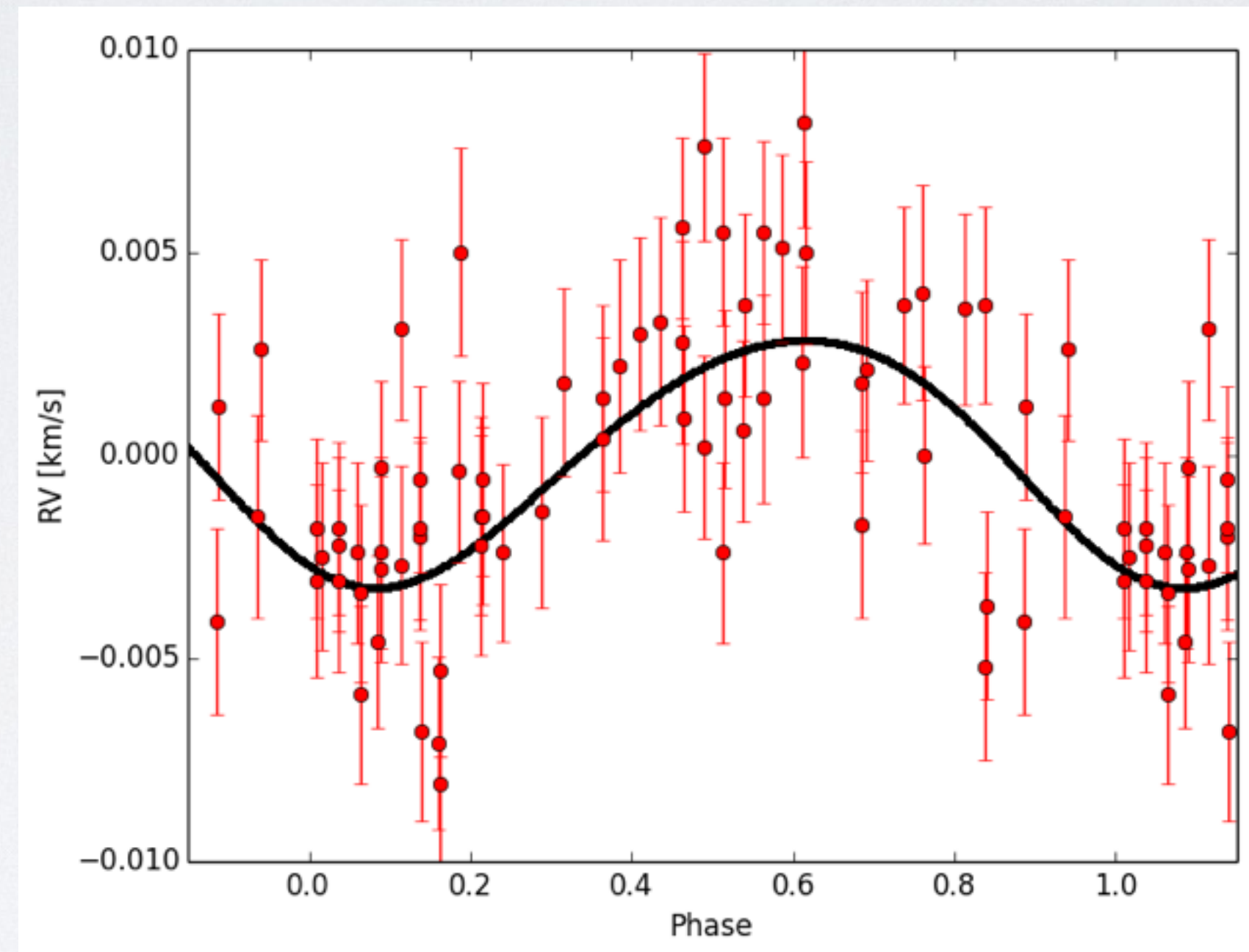
Updated parameters for the HD 190360 system

A NEW NEPTUNE AROUND THE SOLAR ANALOG HD 164595

Courcol, B., Bouchy, F., et al. 2015, A&A, in press (arXiv 1506.07144)



RVs of HD 164595



Phase folded RVs of HD 164595

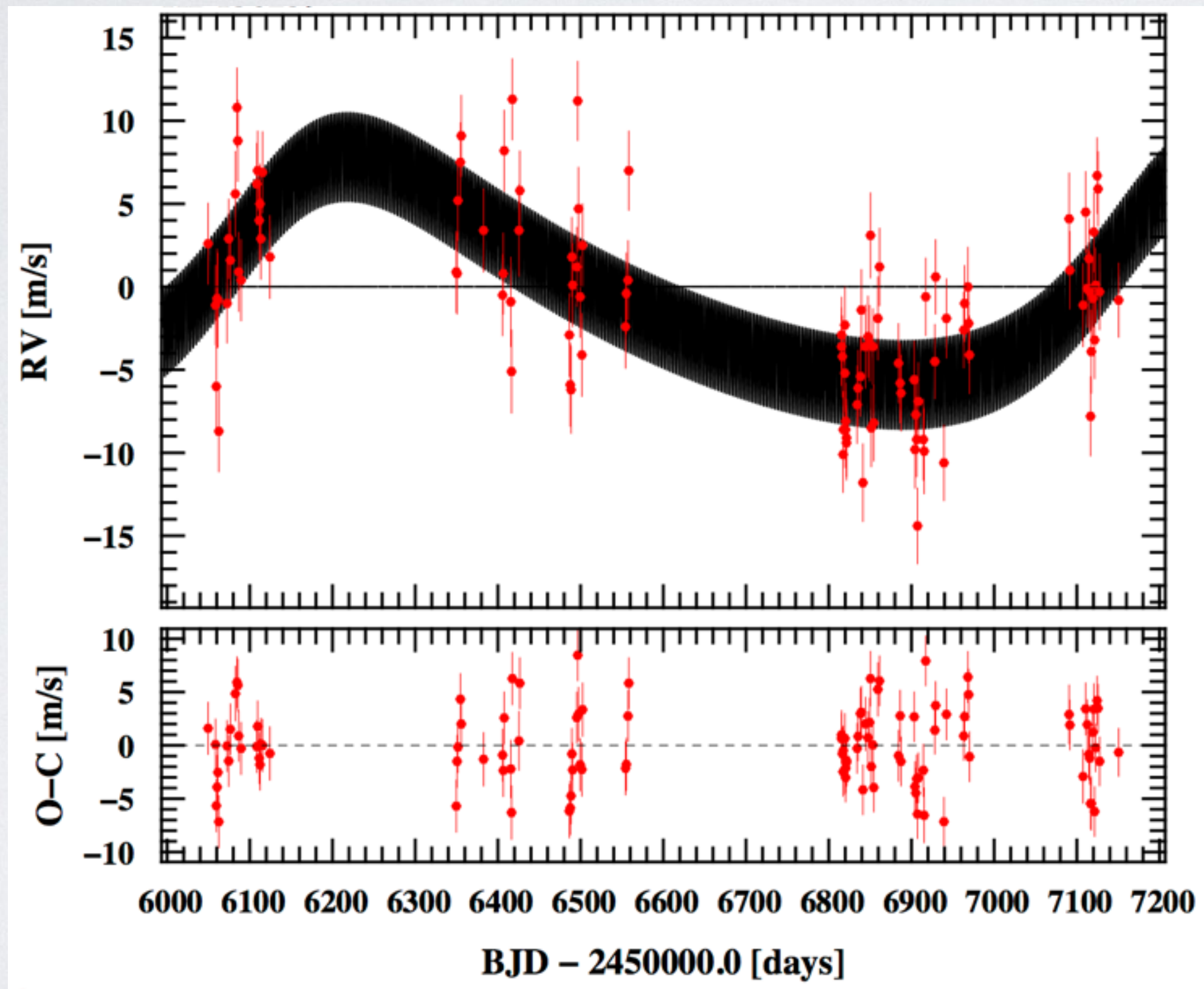
Parameter	HD164595b
P [days]	40.00 ± 0.24
T_0 [BJD]	56280 ± 12
e	$0.088^{+0.12}_{-0.066}$
ω [deg]	145^{+160}_{-110}
K [m s^{-1}]	3.05 ± 0.41
$m \sin i [M_{\oplus}]$	16.14 ± 2.72
a [AU]	0.23
drift [m/s/yr]	-2.34 ± 0.44
N_{meas}	66
Data span [days]	809
σ (O - C) [m s^{-1}]	2.3
Jitter [m s^{-1}]	1.8

Orbital and physical parameters of HD 164595b

Smallest planetary signal ever detected with SOPHIE !

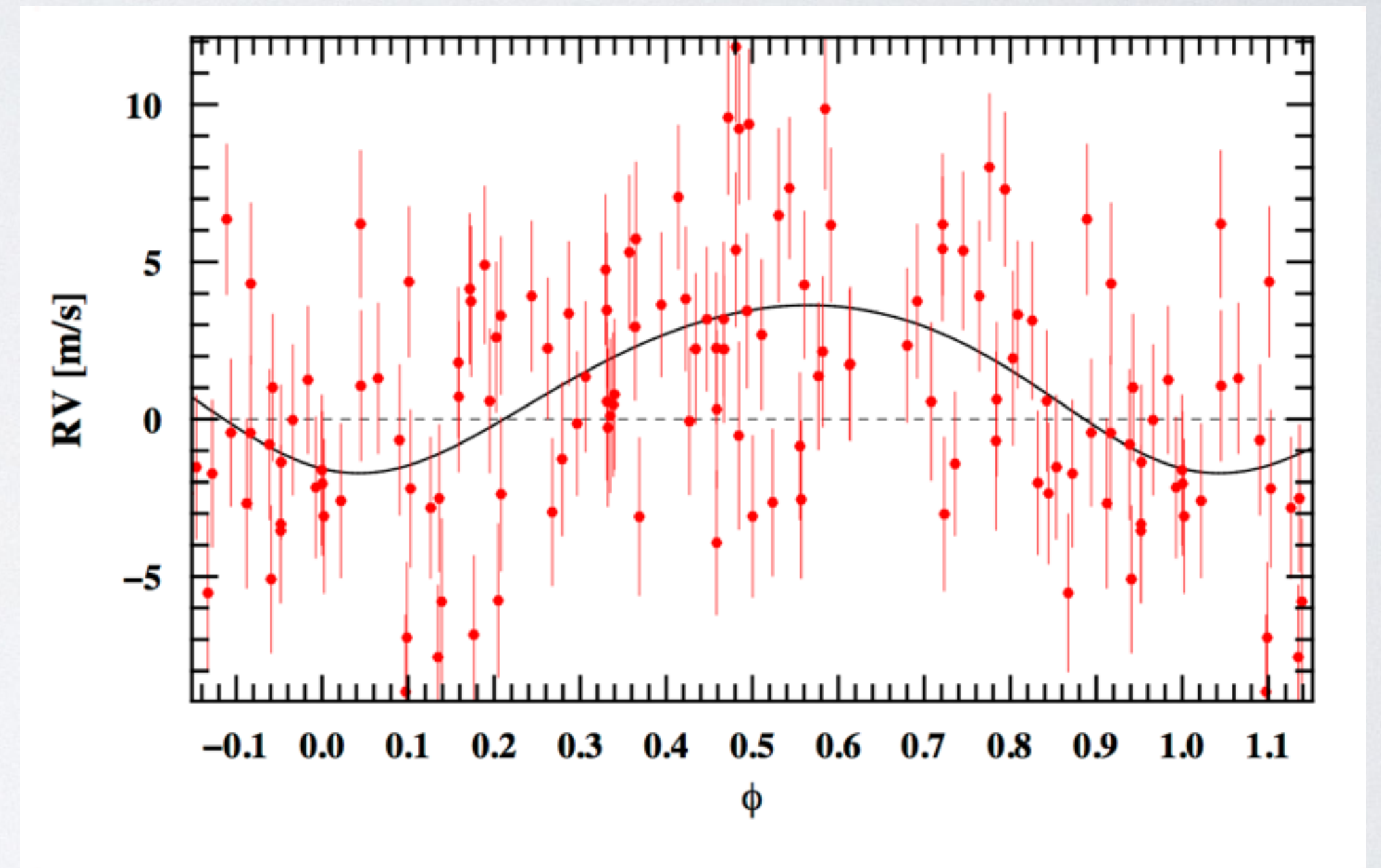
A TWO-PLANETS SYSTEM CANDIDATE

RV time series (~110 points)



Long period
Saturn-like
planet

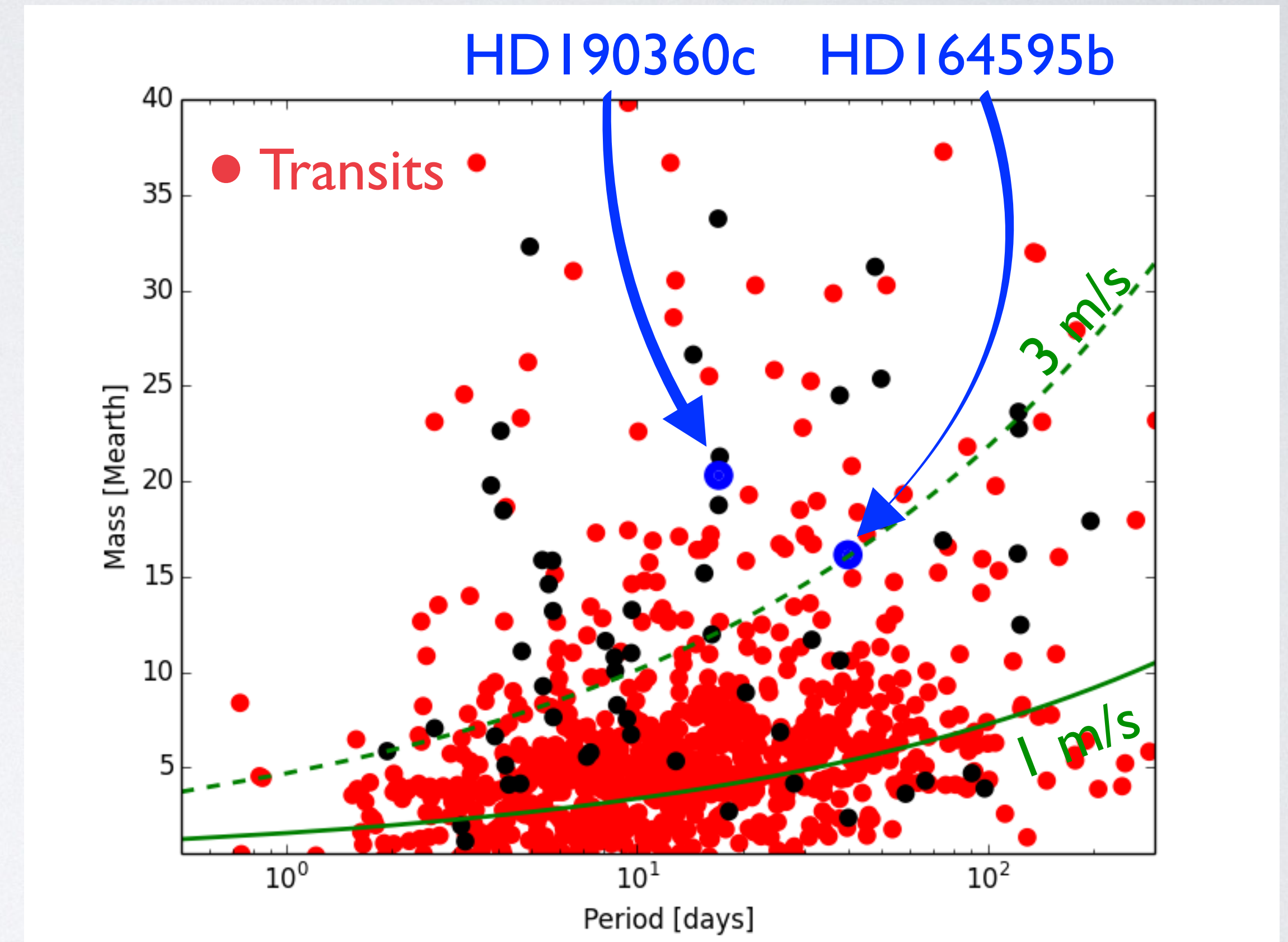
Phase folded RVs of the inner candidate



3.4 days Super-Earth :
 $K=2.6$ m/s !
 $M_{\text{sin}i}=6.3 M_{\text{Earth}}$

THE SMALL PLANETS POPULATION

- Number of small planets known : **>900**

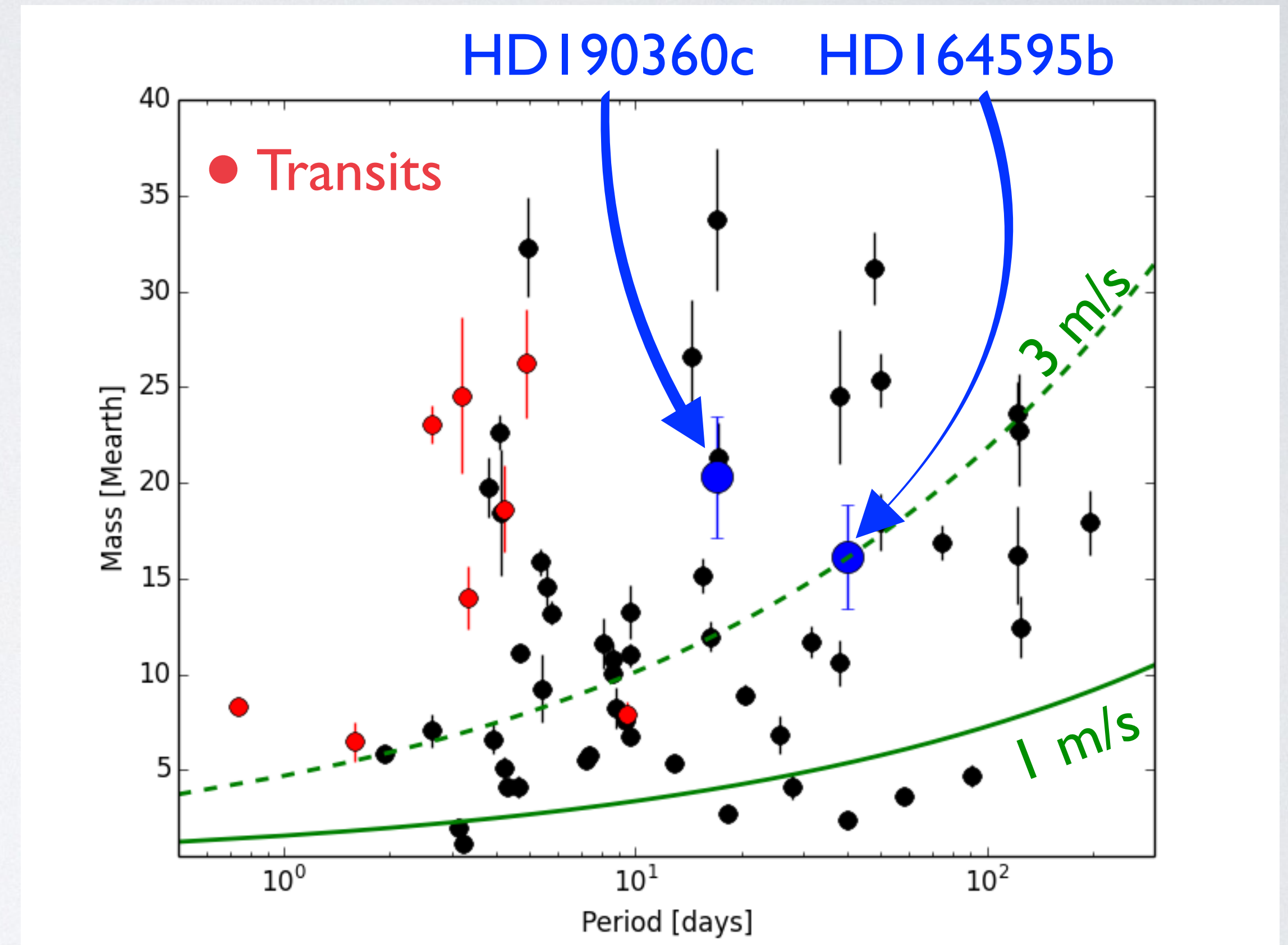


Known small planets

Source : exoplanets.org

THE SMALL PLANETS POPULATION WITH PRECISE RV MASS MEASUREMENTS

- Number of small planets known : **>900**
- Number of small planets with **precise RV mass measurements** : **~60**
 - 12 Neptunes in the 20-100 days range + HDI 64595b
 - **8 transiting planets only !**
No period > 10 days
- Overall transit probability : ~300%
(HDI 64595b transit probability : 2%)



Known small planets with a precision on $M_{\text{sin } i}$ better than **20%**
Source : exoplanets.org

SYNERGIES WITH



CHEOPS : Transit follow-up ESA mission, launch in 2017-2018

- CHEOPS and SOPHIE both focus on **bright stars** to :
 - **Precisely characterize « hot » to « warm » objects** (> 10 days) \Rightarrow golden targets for atmospheric spectroscopy
 - **Cover the parameters space of the Neptune/Super-Earth transition** : period, mass, radius, density, stellar masses, metallicities, irradiation, planet multiplicity...
 \Rightarrow we need more planets !
- In practice, SOPHIE is a target provider for CHEOPS in the northern hemisphere (**~ 15 expected new planets in 2018**) and perform preliminary observations

CONCLUSION

- **Lack of planets** with precise mass (and radius) to cover the parameters space of the Neptune/Super-Earth transition
- Low mass planets require **dedicated telescopes** (100+ measurements) with **high precision spectrographs** (1-2 m/s)
- **SOPHIE is now able to detect and characterize such low mass planets**
- Upcoming software and hardware improvements on SOPHIE, **objective : 1 m/s**