



# LEGA-C + CANDELS

## Constraining the star formation histories of galaxies

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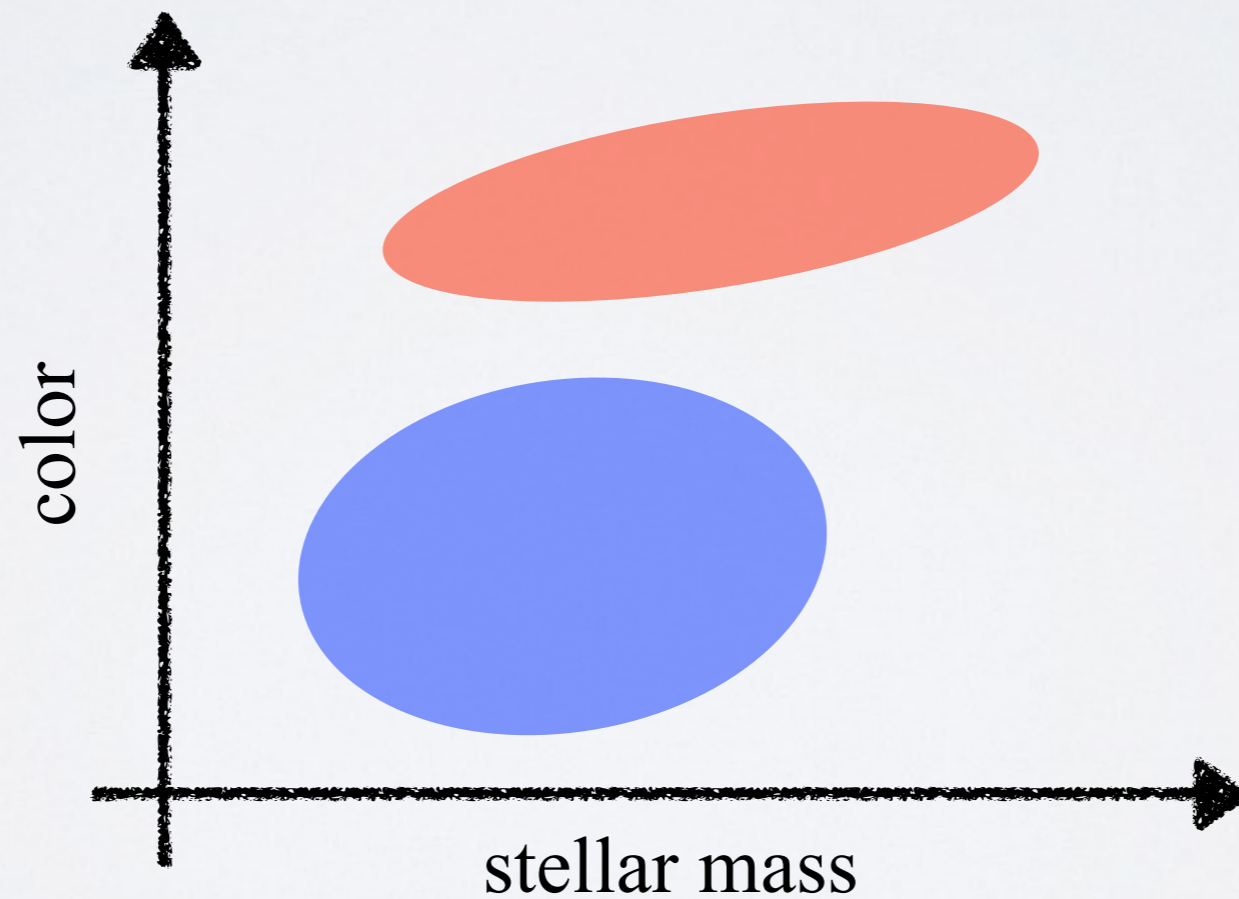
# The evolution of galaxies



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Bimodality in galaxy properties



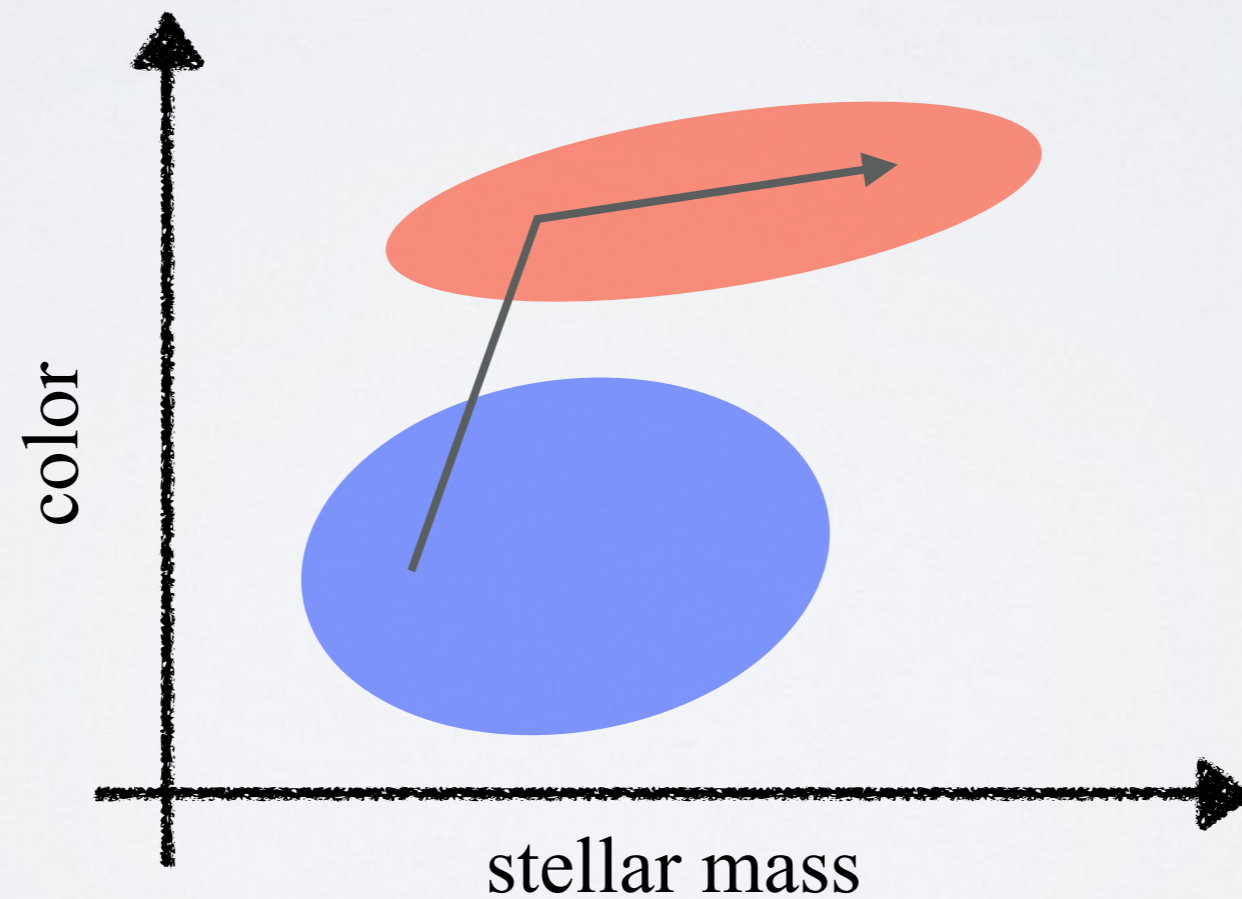
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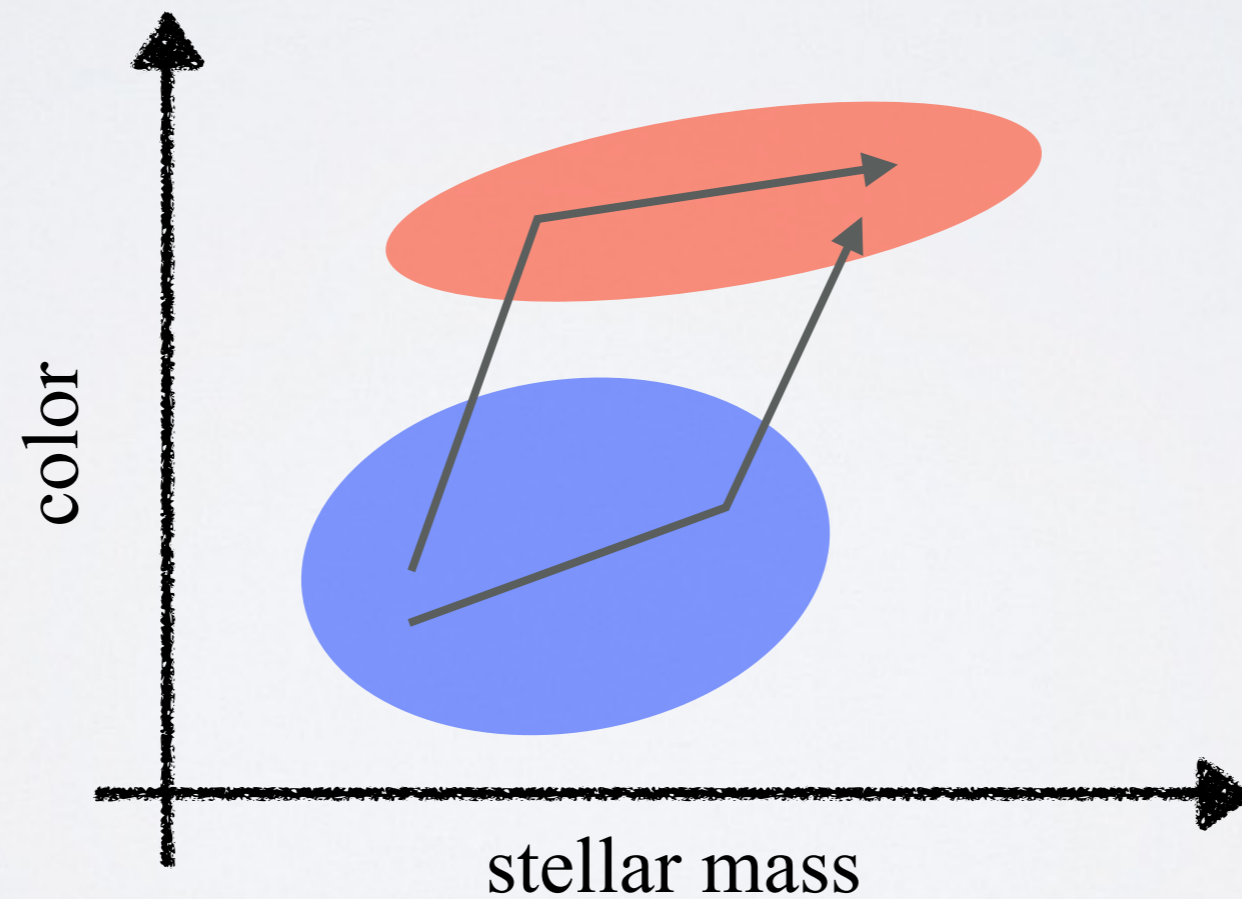
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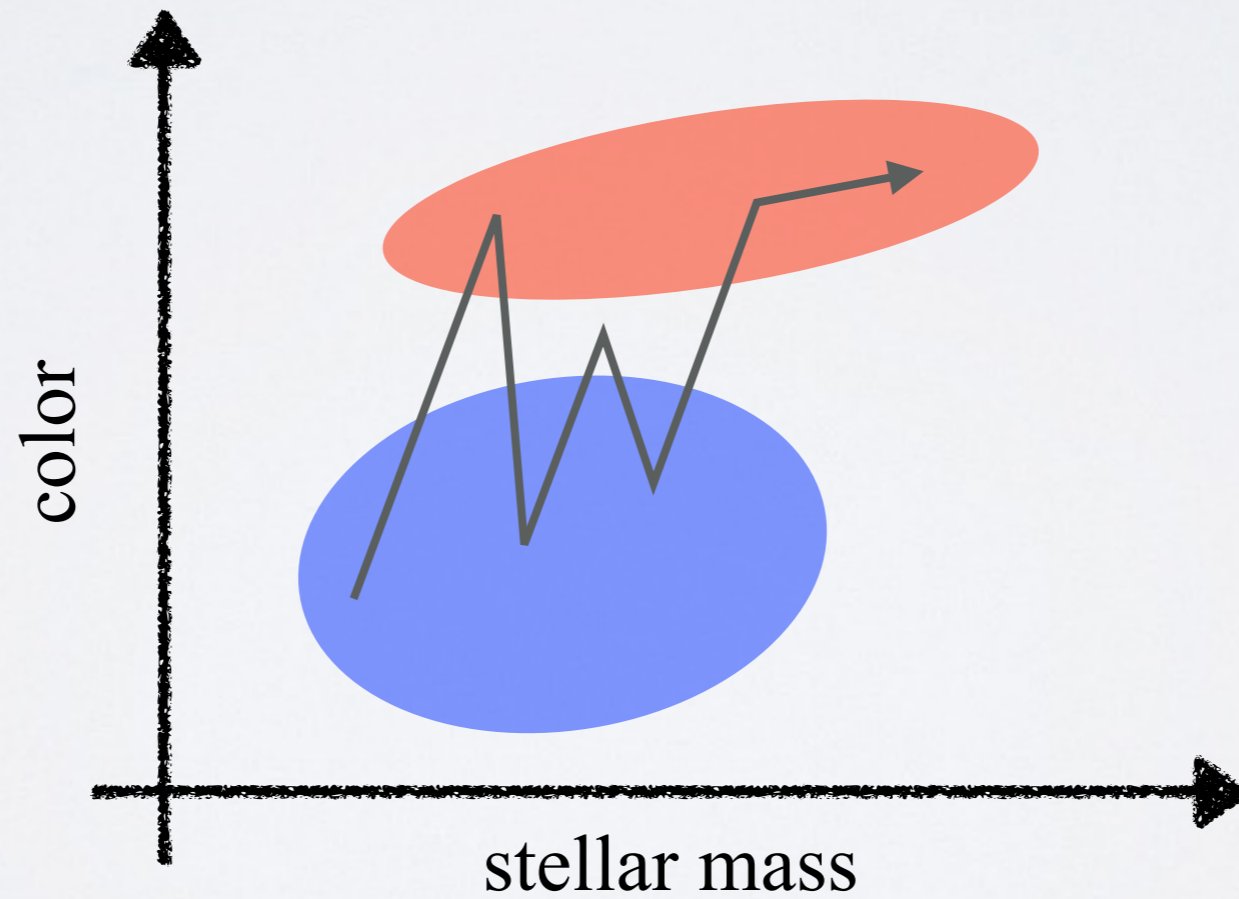
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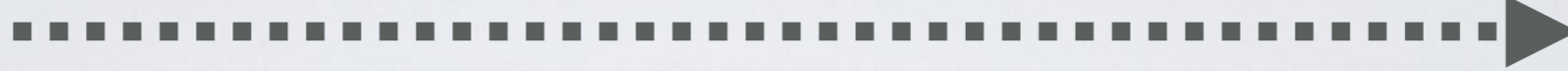
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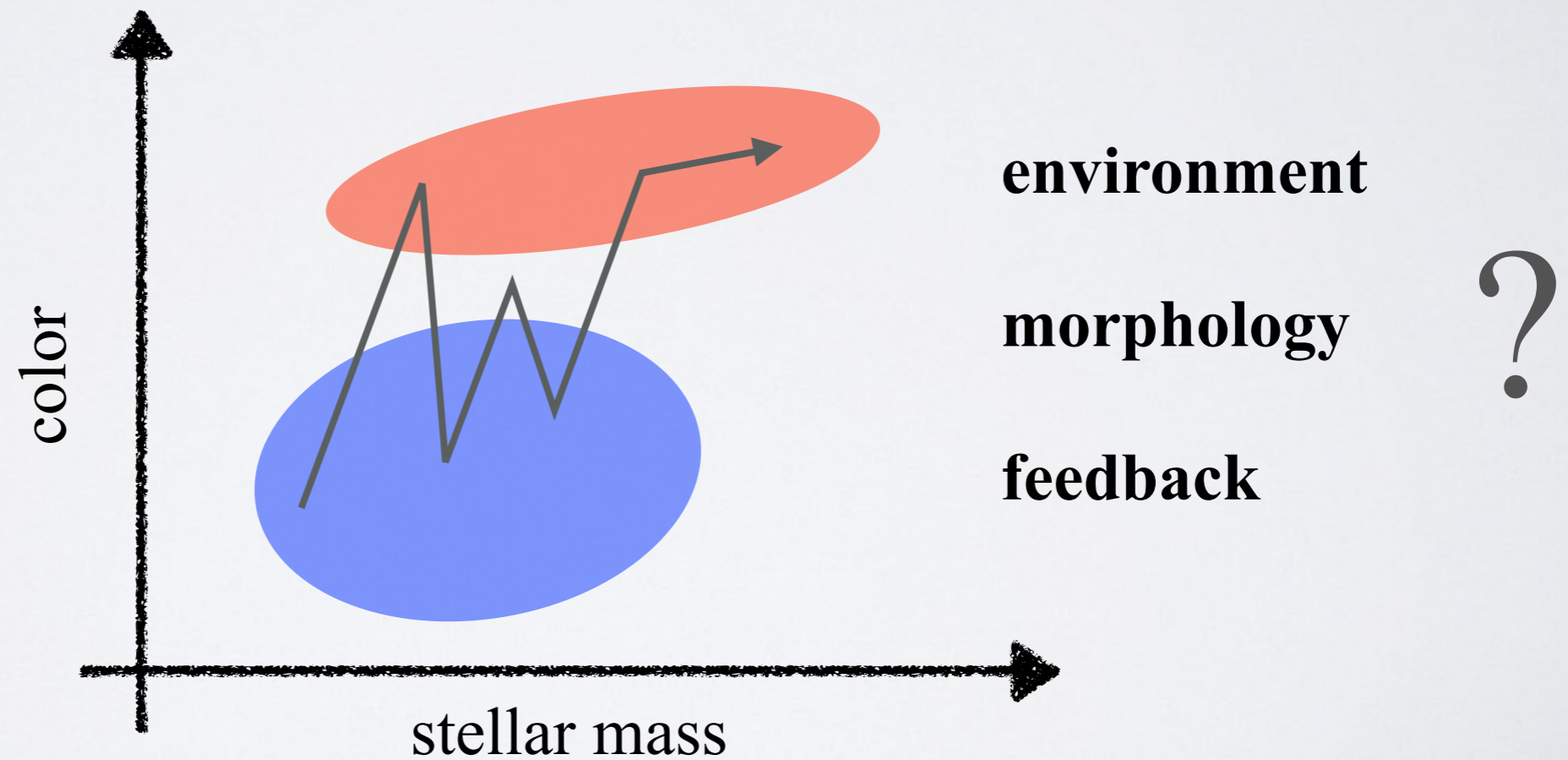
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## Measuring ages and metallicities from spectra

Worthy et al. 1992, Faber et al. 1995, Trager et al. 2000, Schiavon et al. 2007,  
Choi et al, 2014, Gallazzi et al. 2014

## Connecting populations of galaxies via abundance matching

van Dokkum et al. 2013, Patel et al. 2013, Papovich et al. 2015

## Measuring SFHs form the fossil record in the data

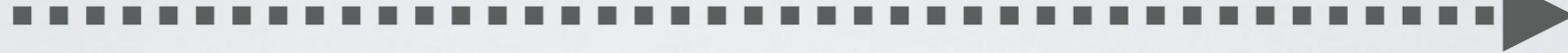
Panter et al. 2003, 2007, McDermid et al. 2015, Pacifici et al. 2016

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# Models of galaxy spectral energy distributions

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Build a library of model spectra to best reproduce the observations

Ingredients:

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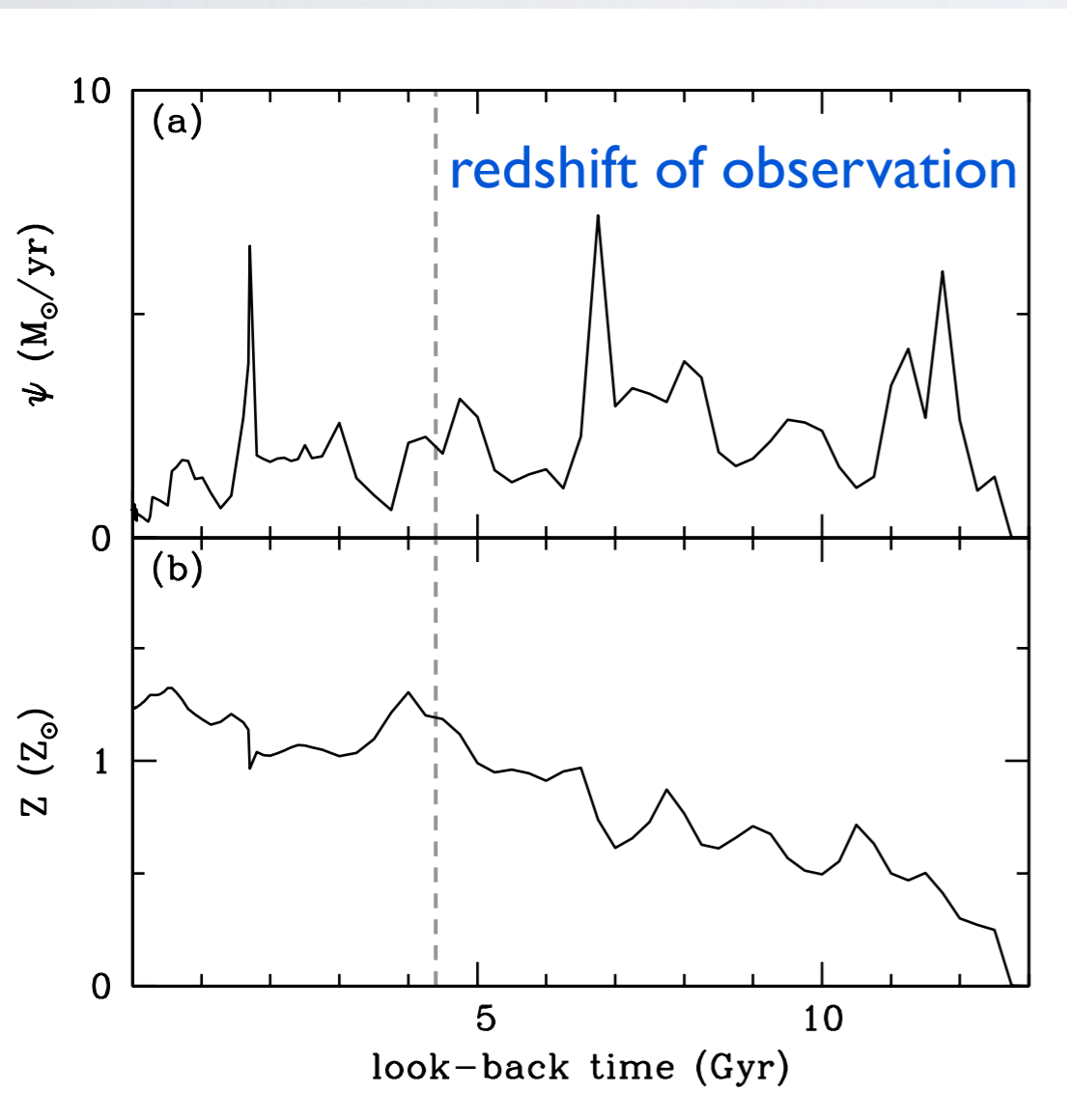
star formation and chemical enrichment histories

# Models of galaxy spectral energy distributions

Build a library of model spectra to best reproduce the observations

Ingredients:

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semi-analytic post-processing  
of the **Millennium Simulation**

Springel et al. (2005), Croton et al. (2006)  
De Lucia & Blaizot (2007), Hirschmann et al. (2015)

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Build a library of model spectra to best reproduce the observations

Ingredients:

star formation and chemical enrichment histories

emission by the stars

emission by the gas

effect of the dust

use models as priors to interpret the observations

(e.g. BEAGLE)



# Application to different datasets

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- DEEP 2 photometry and spectroscopy at  $z \sim 1$
- 3D-HST photometry and grism spectroscopy
- GALEX + SDSS + WISE photometry at  $z \sim 0$
- CANDELS photometry at  $0.2 < z < 2.1$
- LEGA-C spectroscopy +  
UltraVISTA/CANDELS photometry

# Application to different datasets

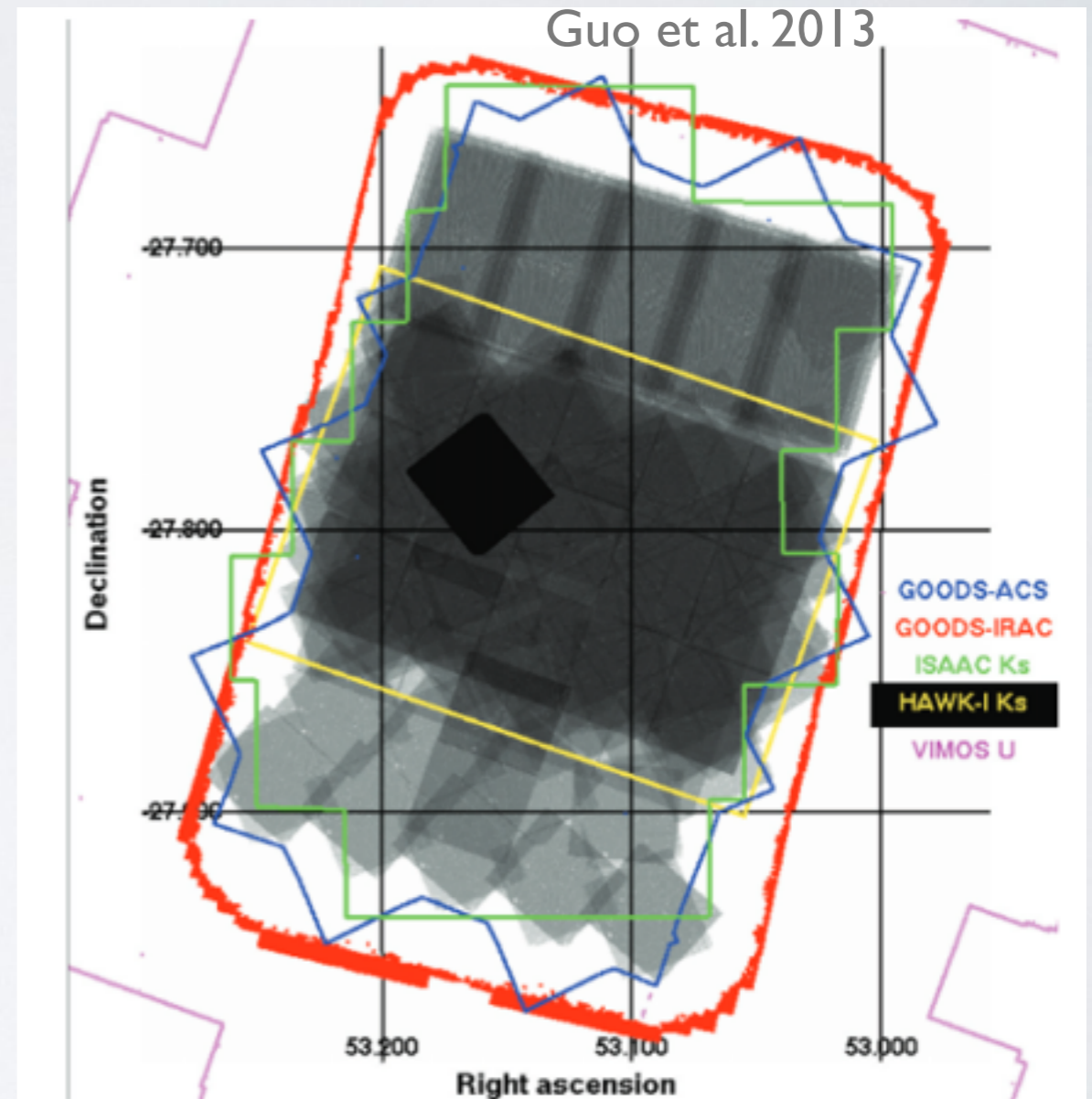
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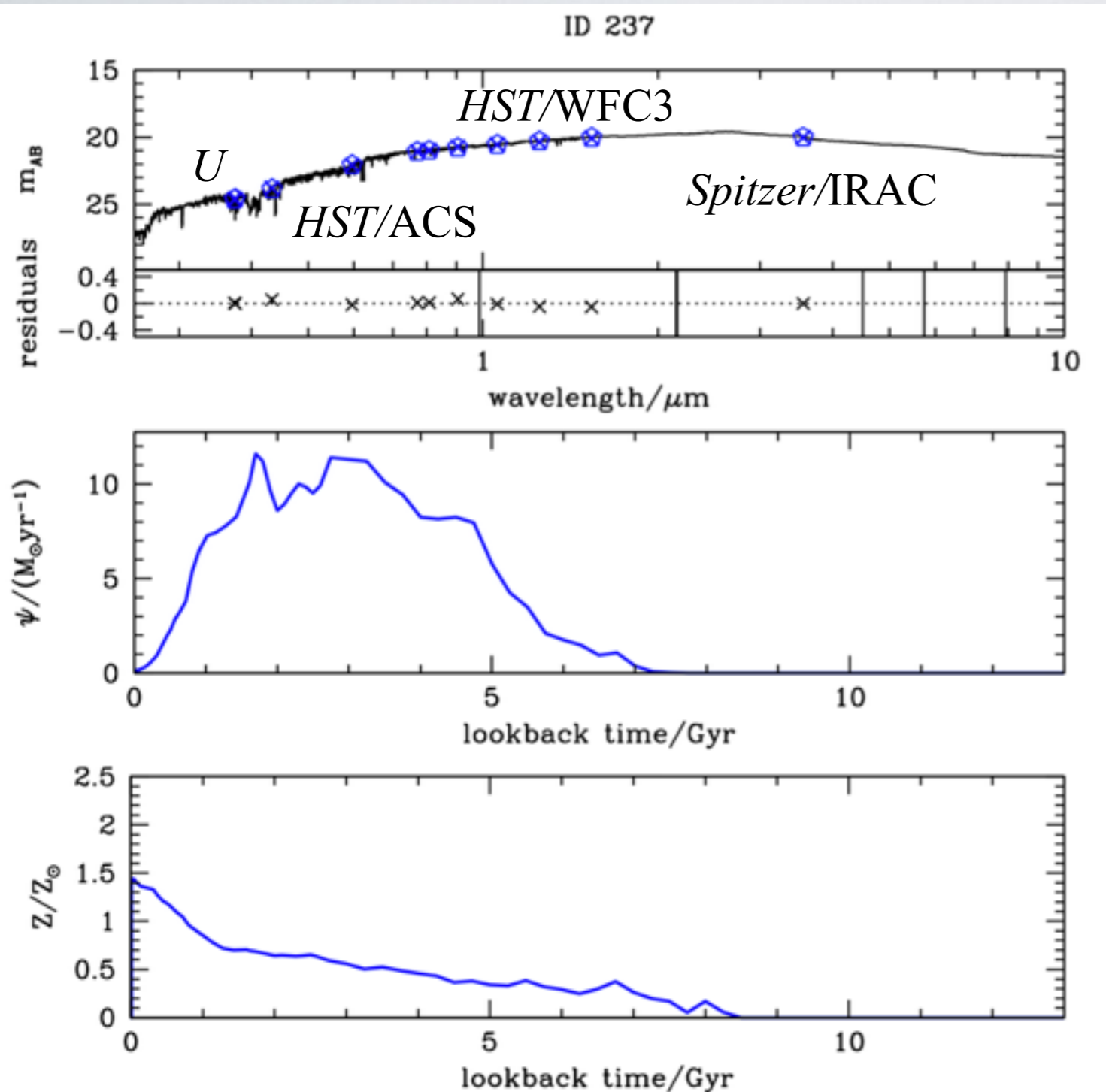
# CANDELS sample

- CANDELS observations
  - 17 photometric bands sampling the rest-frame UV to NIR
- GOODS-South and GOODS-North
- $0.2 < z < 2.1$
- $H < 26$  and stellar mass  $\log(M/M_{\odot}) > 9$
- We select 6183 galaxies
- 861 are quiescent galaxies

GOODS-South  
Guo et al. 2013



# Fitting procedure - photometry



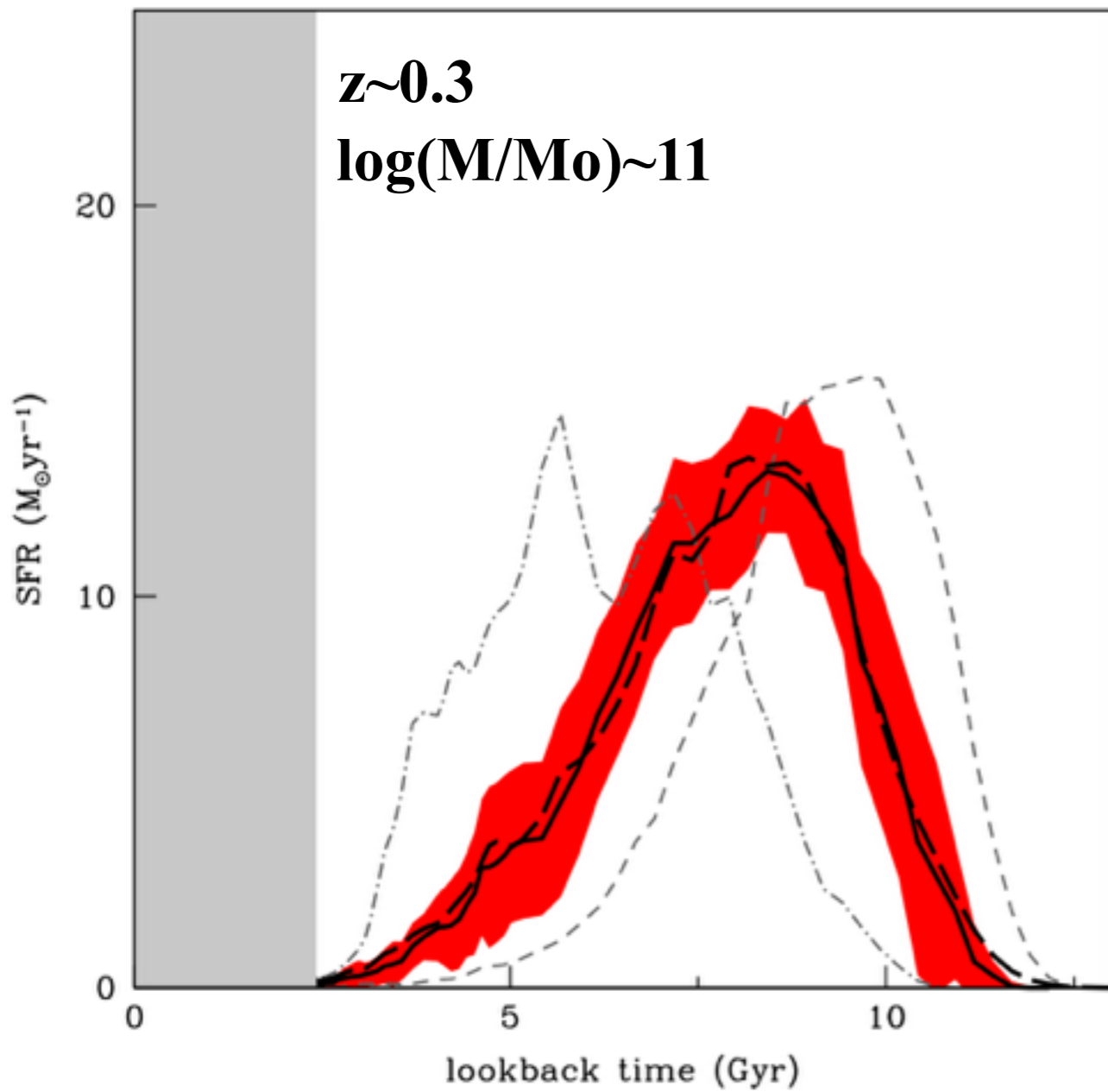
fit the SED, using our library of physically motivated models

extract the best-estimate **star formation history** by averaging all the model SFHs weighted by their likelihoods

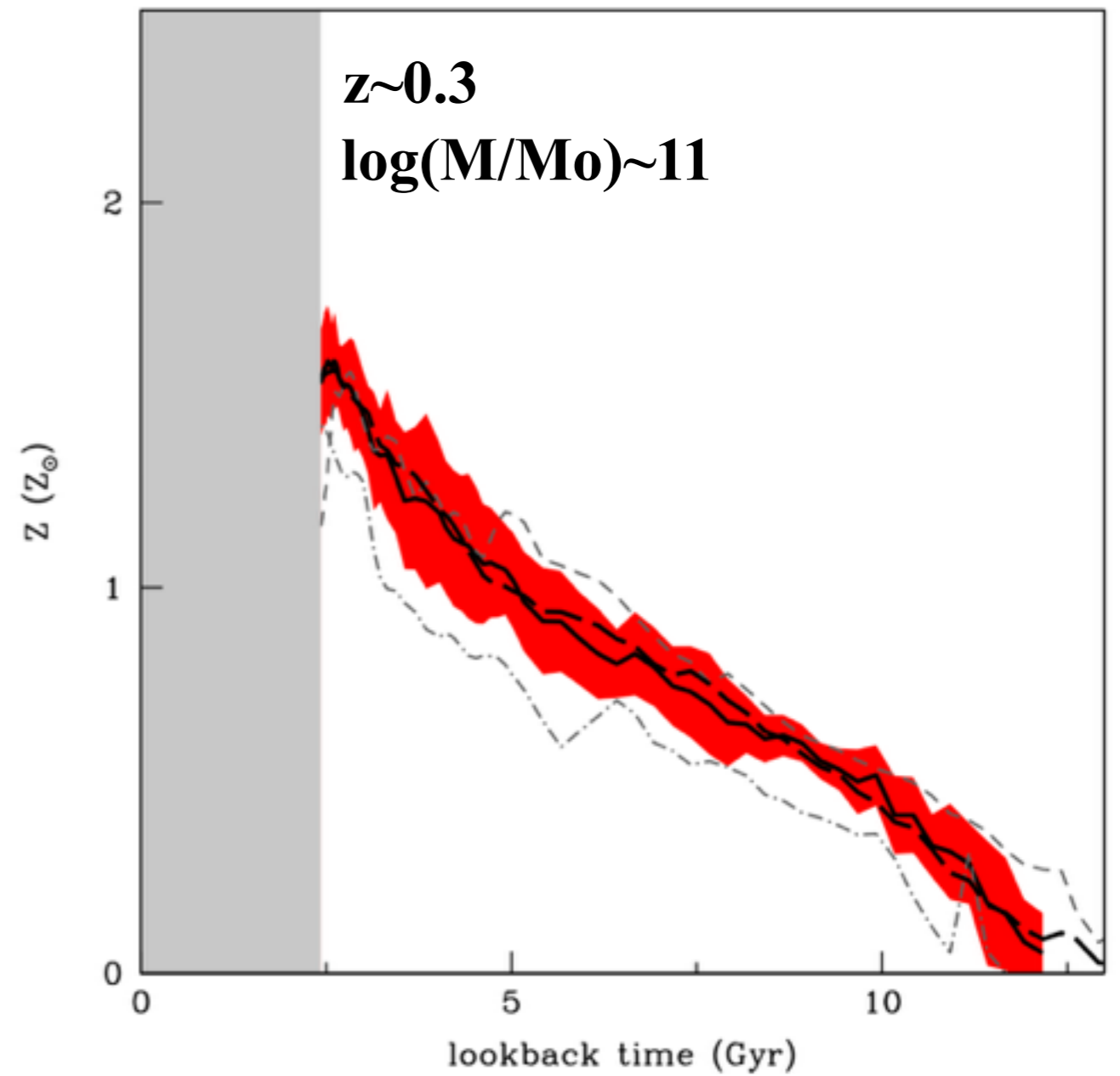
extract the best-estimate **metal enrichment history** by averaging all the model histories weighted by their likelihoods

# The star formation histories of galaxies

median star formation history

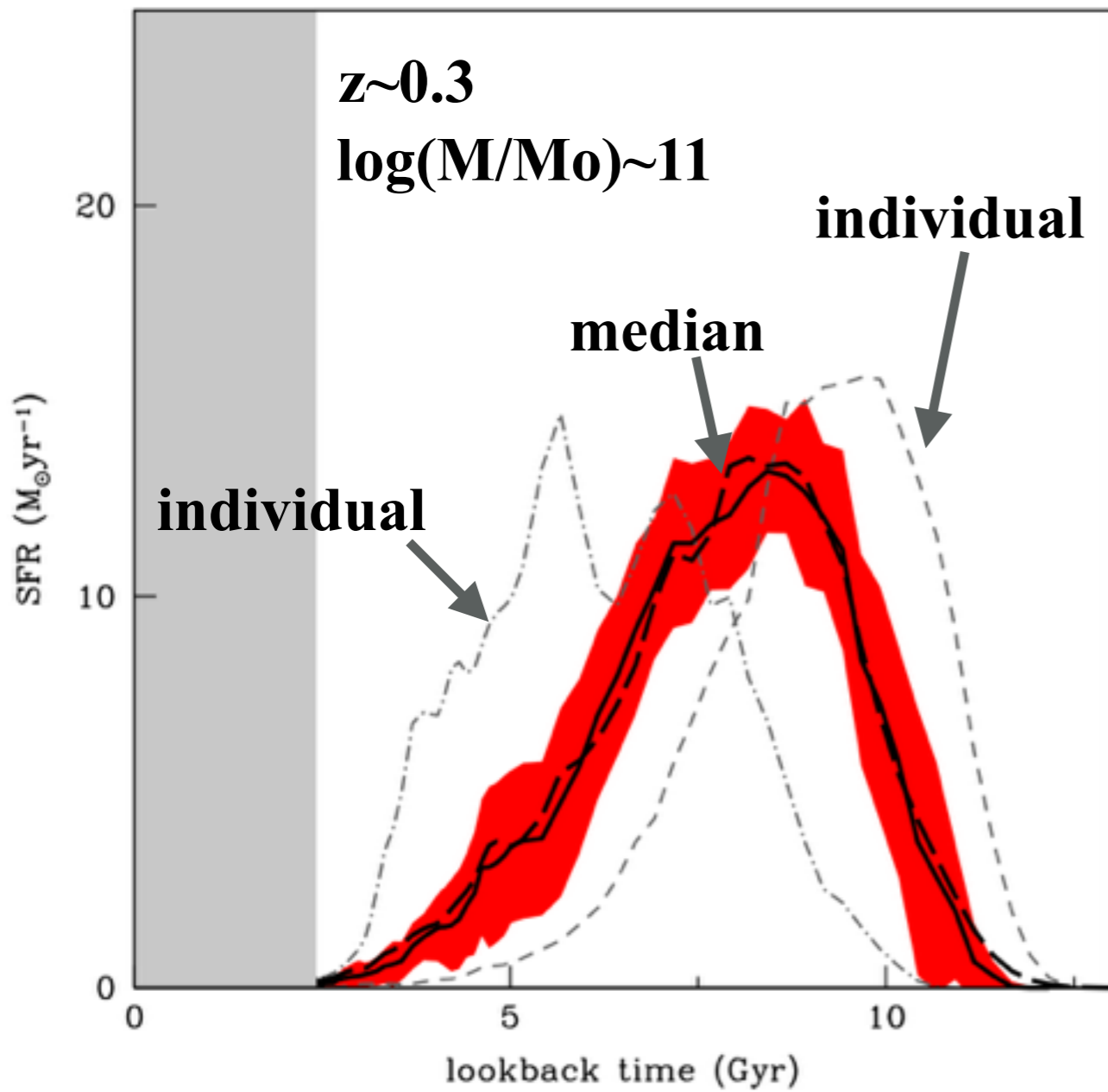


median metal enrichment history

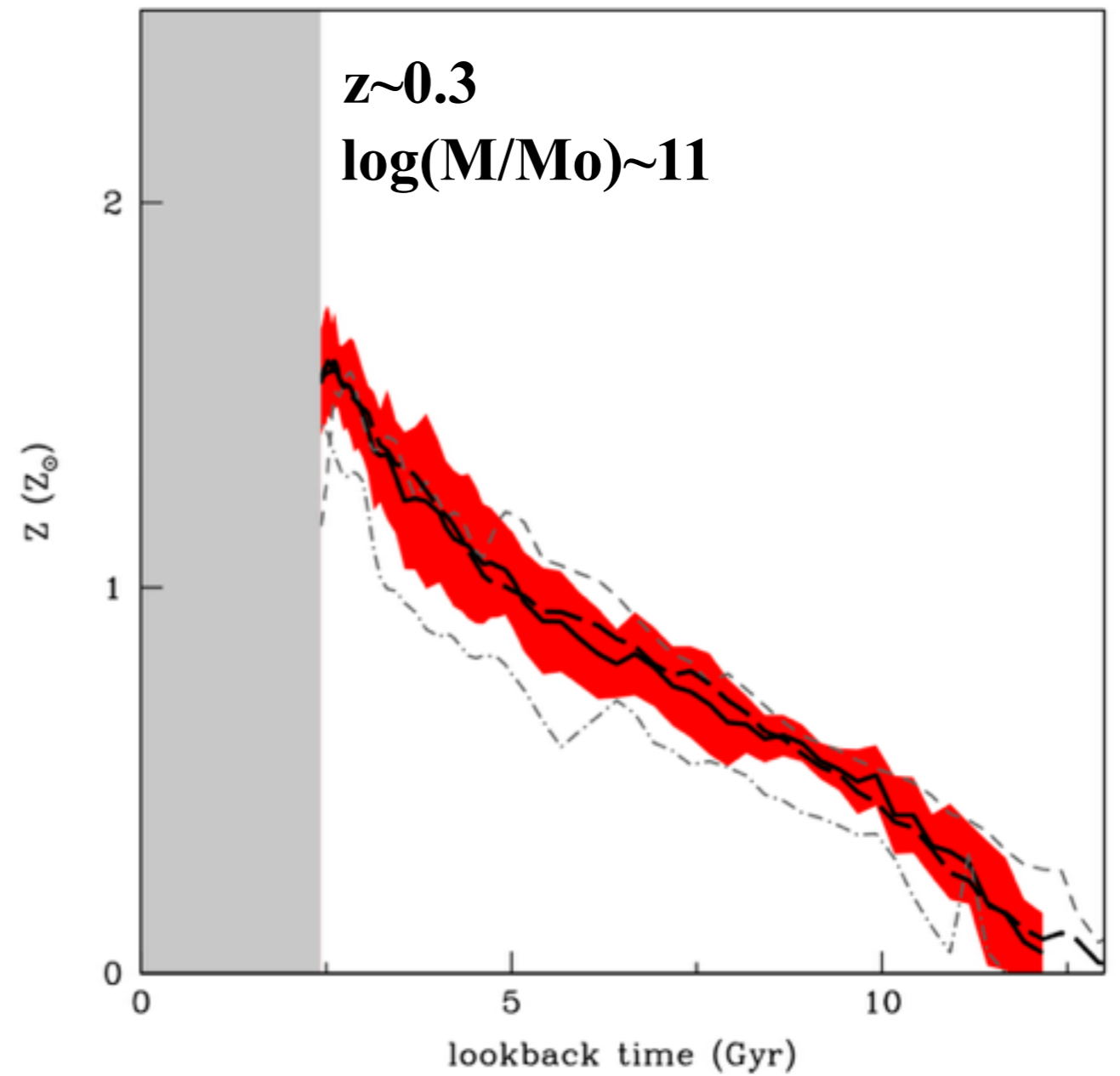


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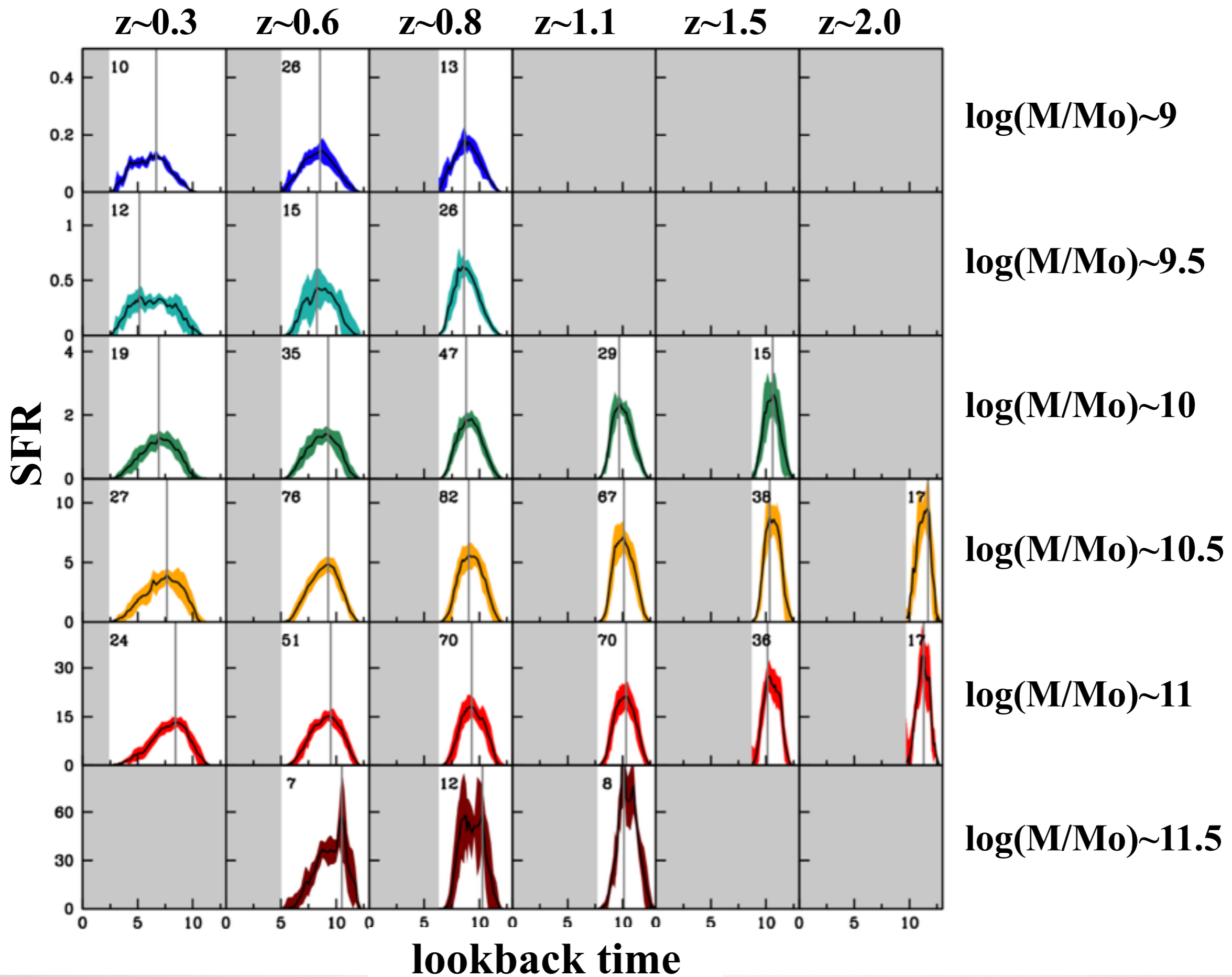
median star formation history



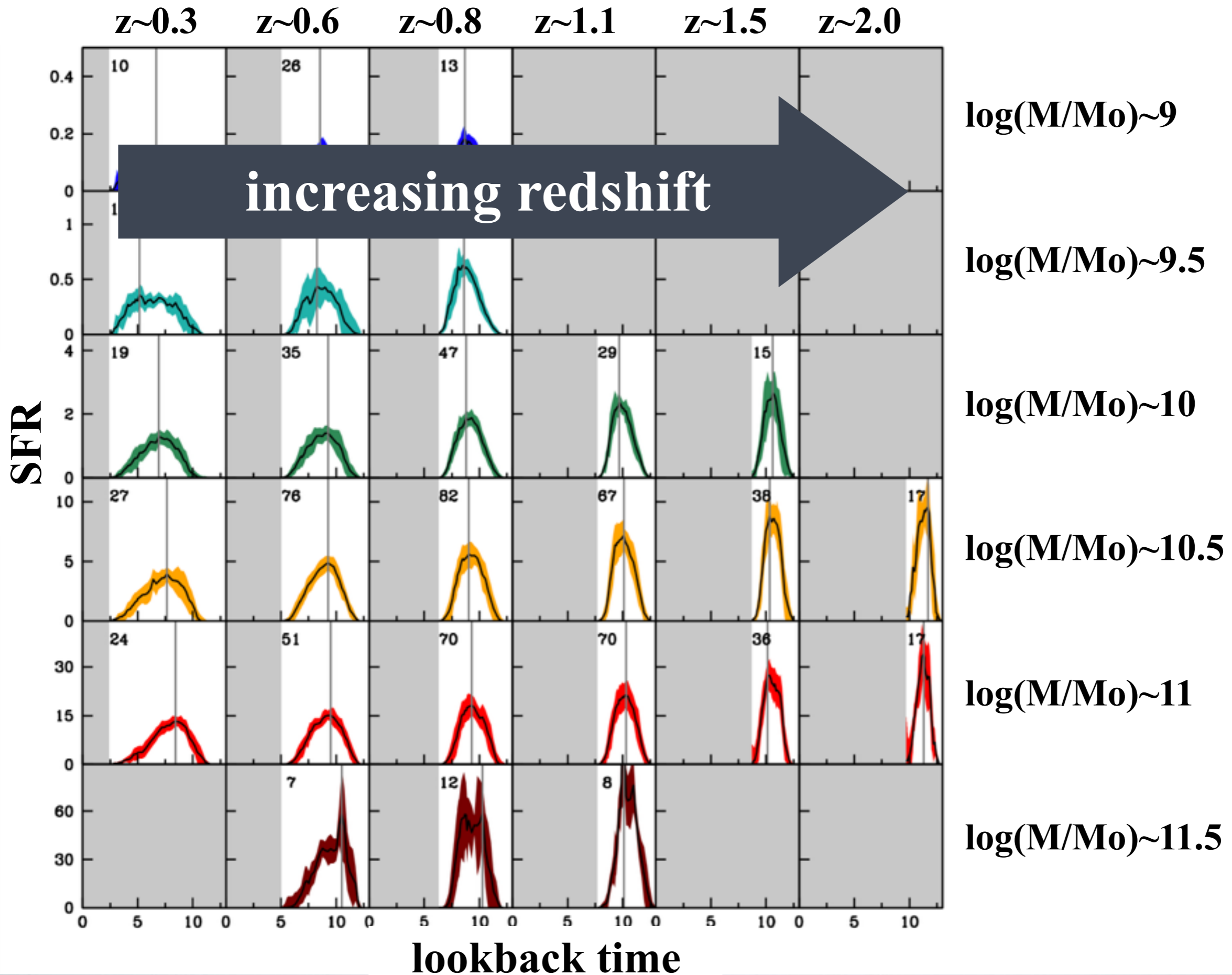
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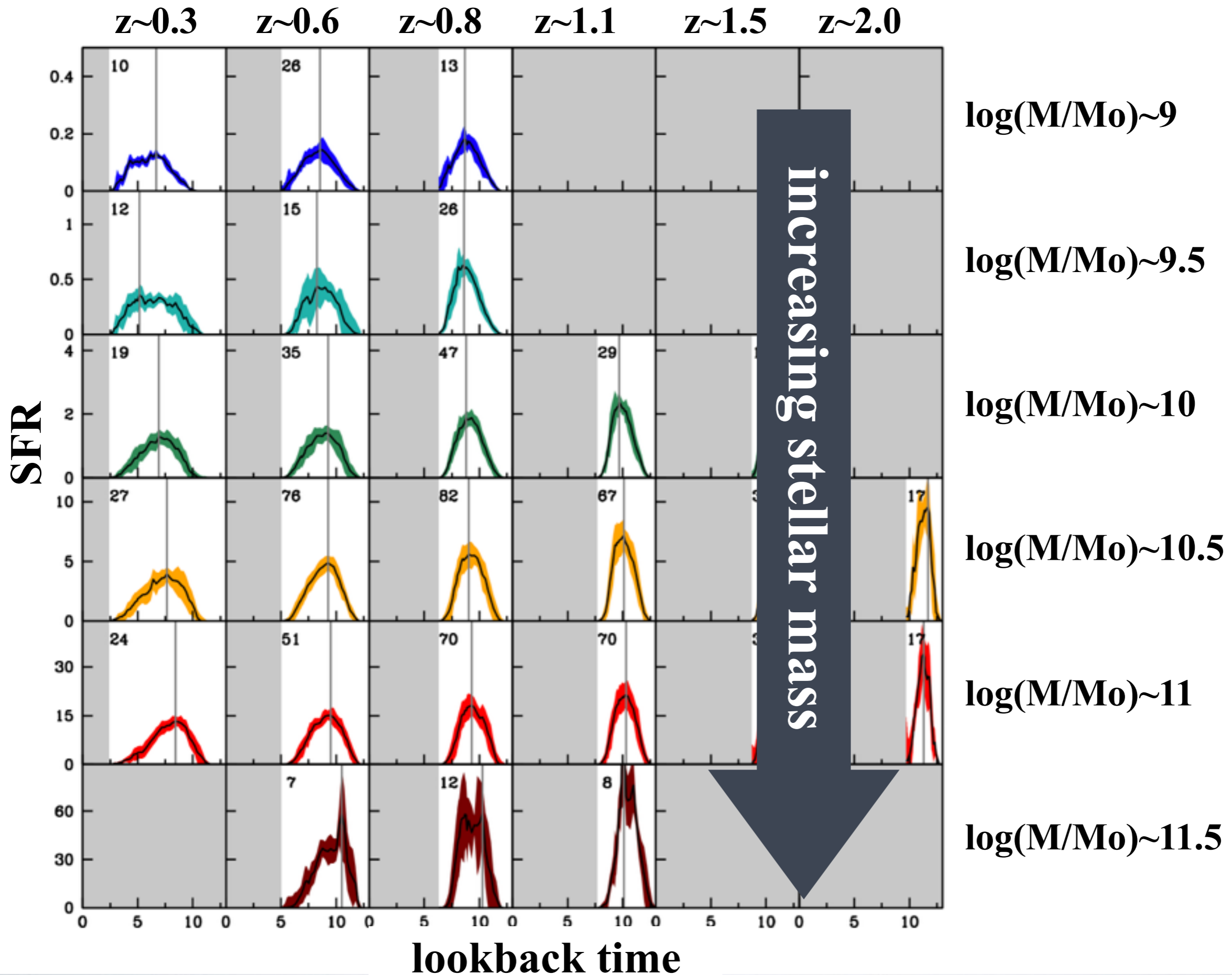


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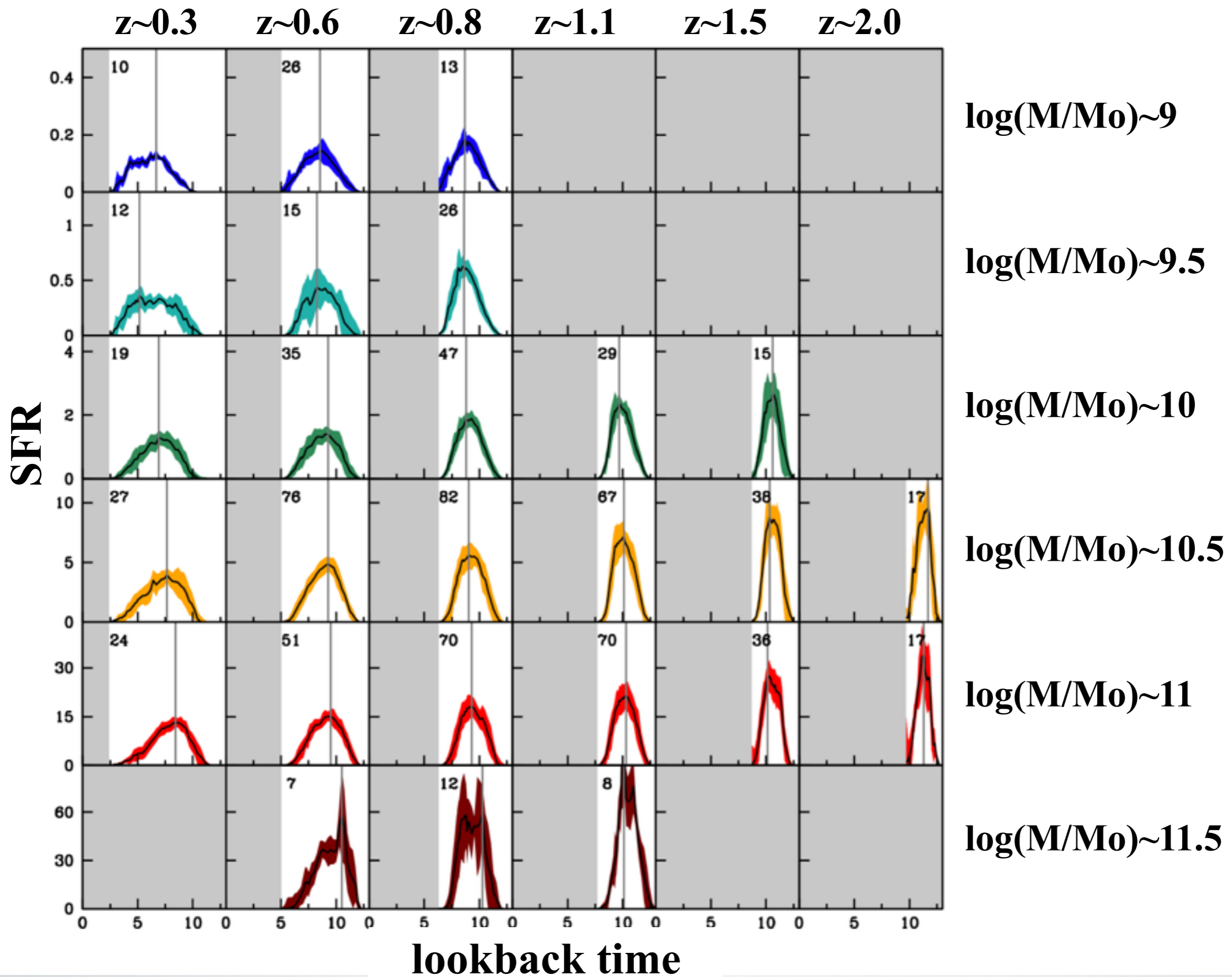




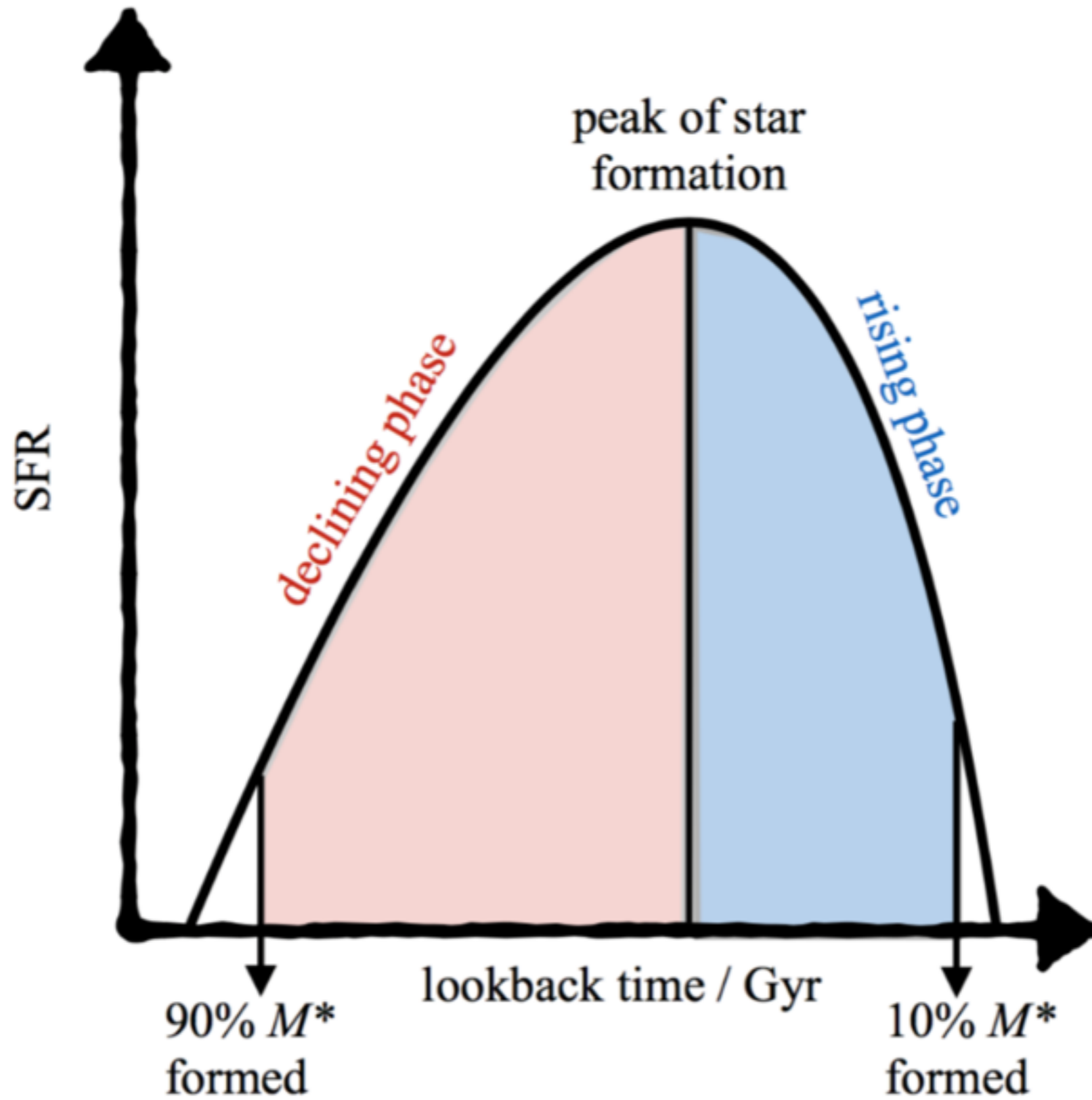
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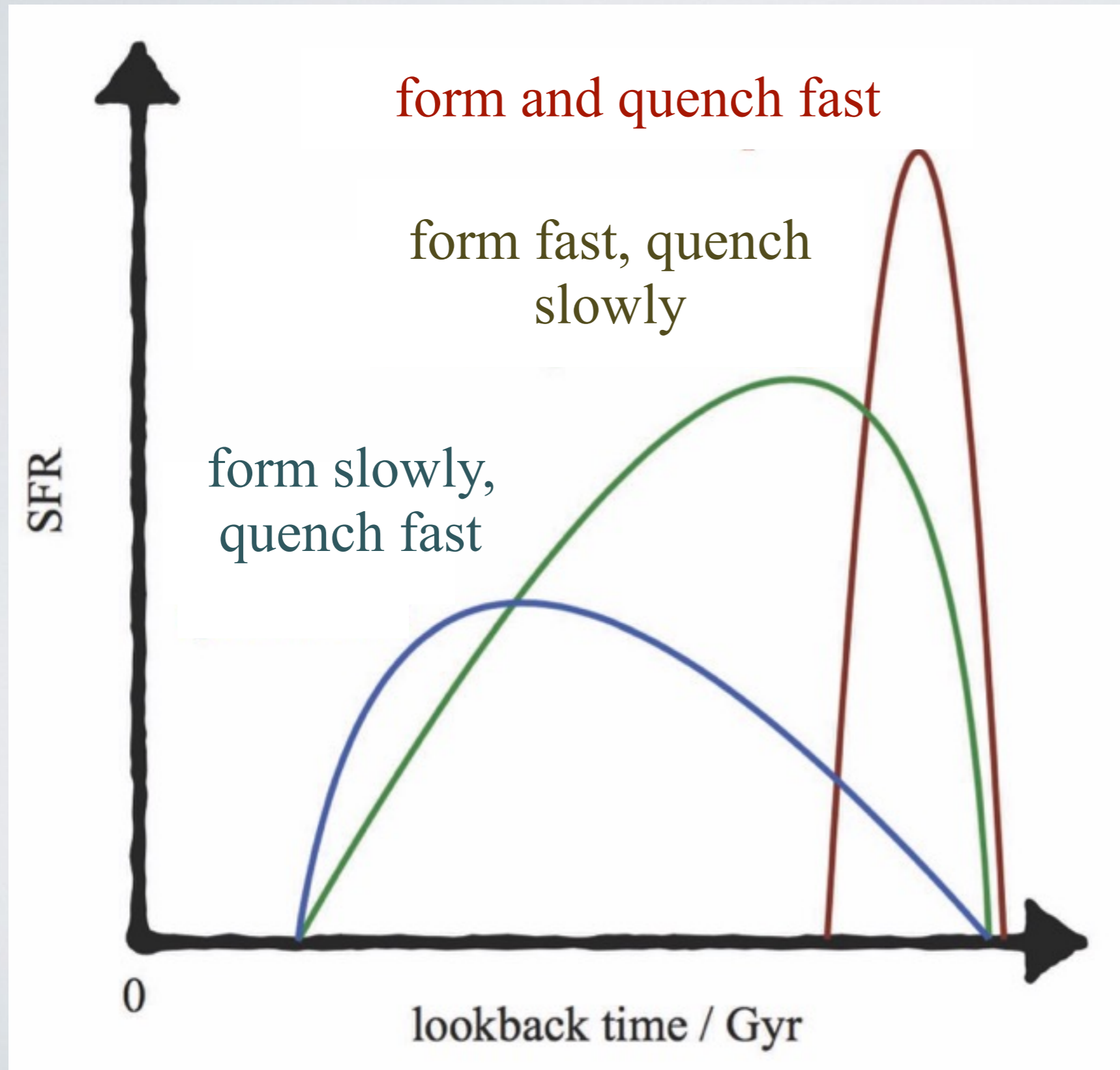
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**high mass, high z**

$10^{11}M_{\odot}$ ,  $z \sim 2$

**high mass, low z**

$10^{11}M_{\odot}$ ,  $z \sim 0.4$

**low mass, low z**

$10^{9.5}M_{\odot}$ ,  $z \sim 0.4$

# Application to different datasets

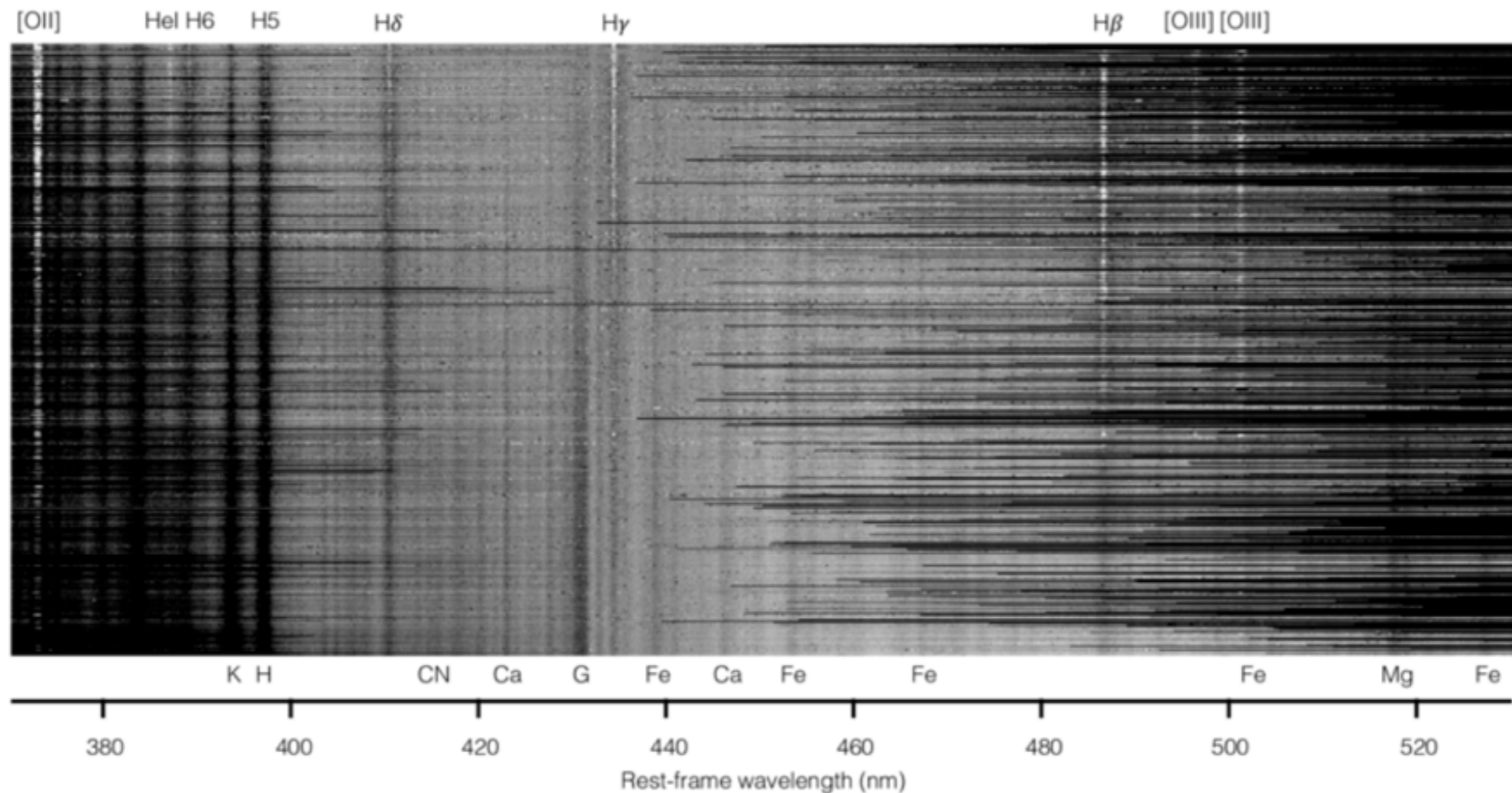
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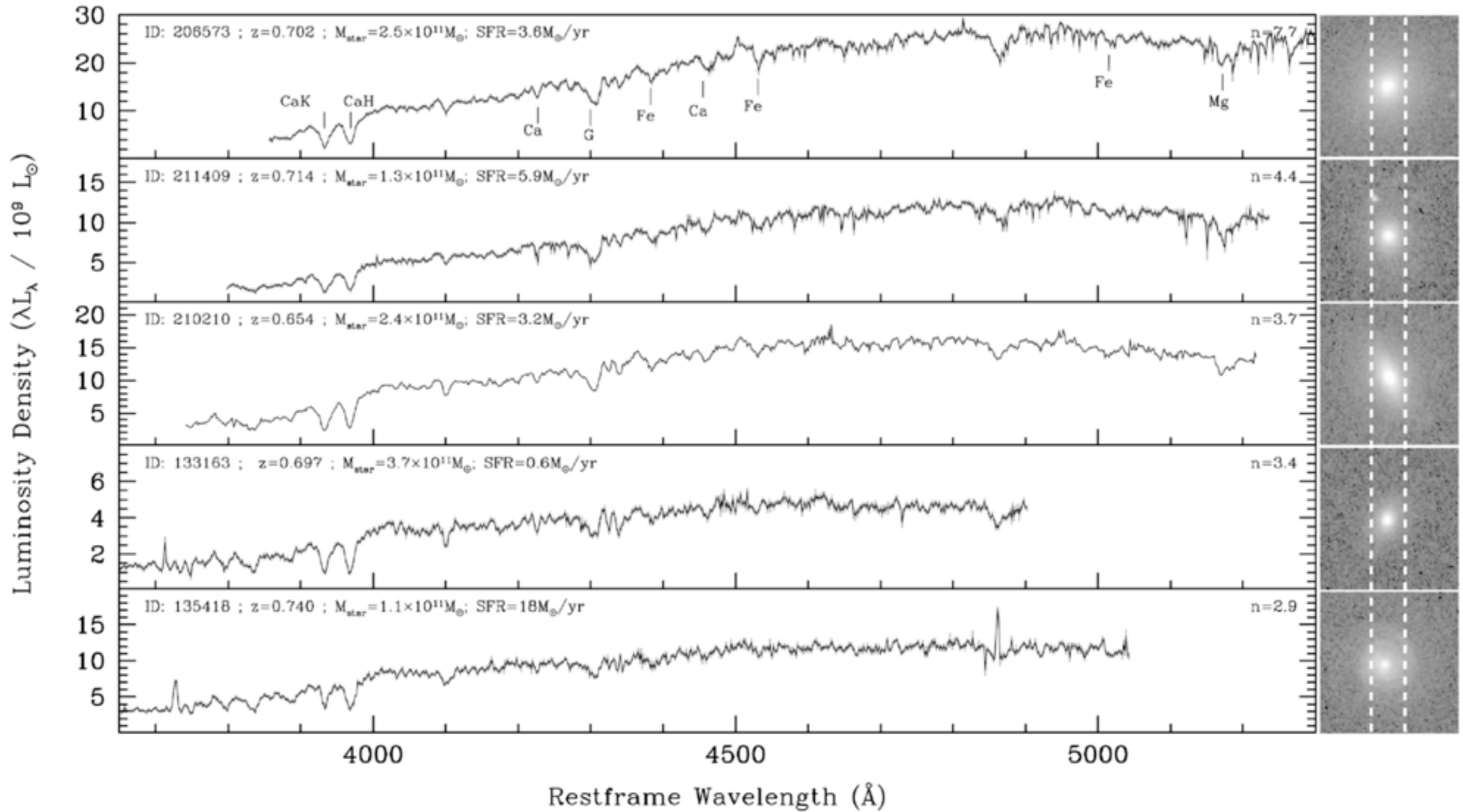
# LEGA-C - spectroscopy

- VLT/VIMOS, in COSMOS field
- $0.6 < z < 1.0$
- targets selected from the UltraVISTA catalog in K band
- In total  $\sim 3000$  galaxies,  $\sim 900$  so far

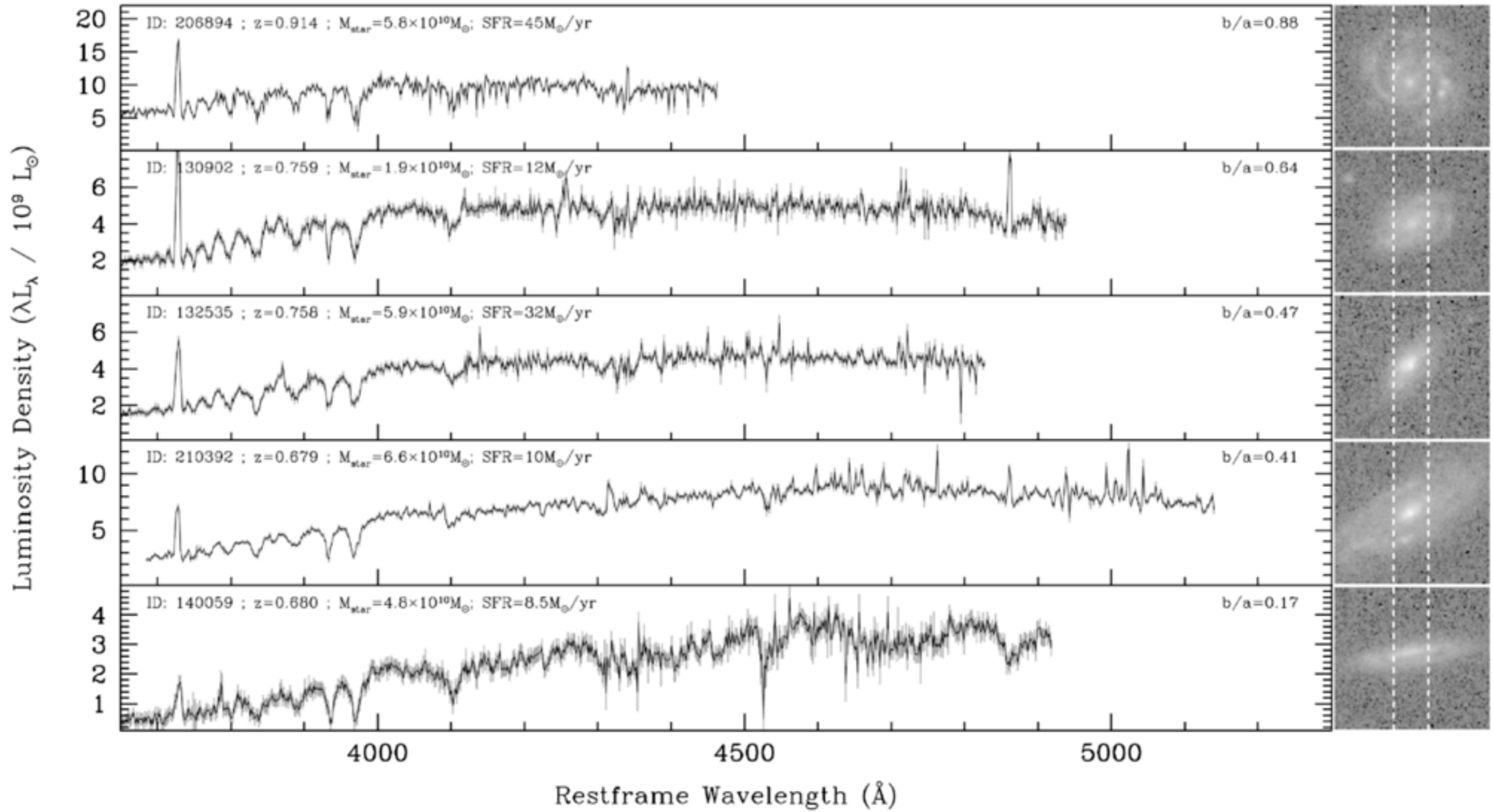
van der Wel et al. 2016



# LEGA-C

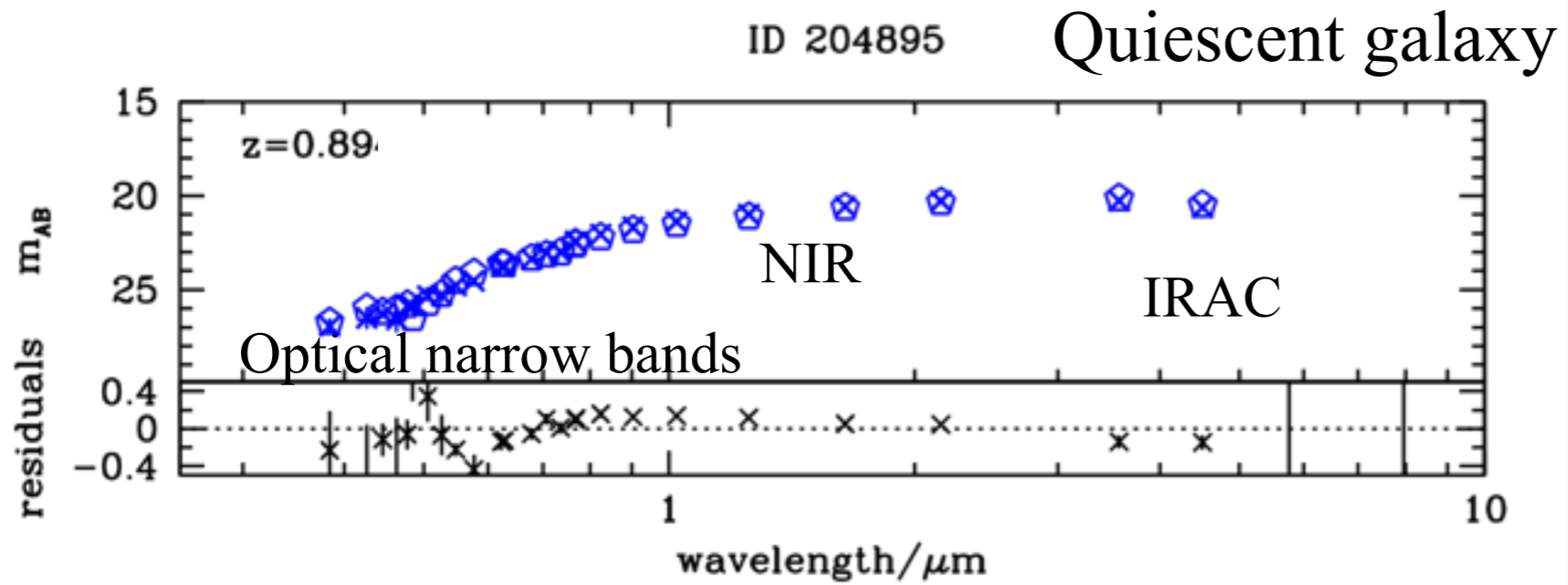


# LEGA-C





# Fitting procedure - photometry+spectroscopy

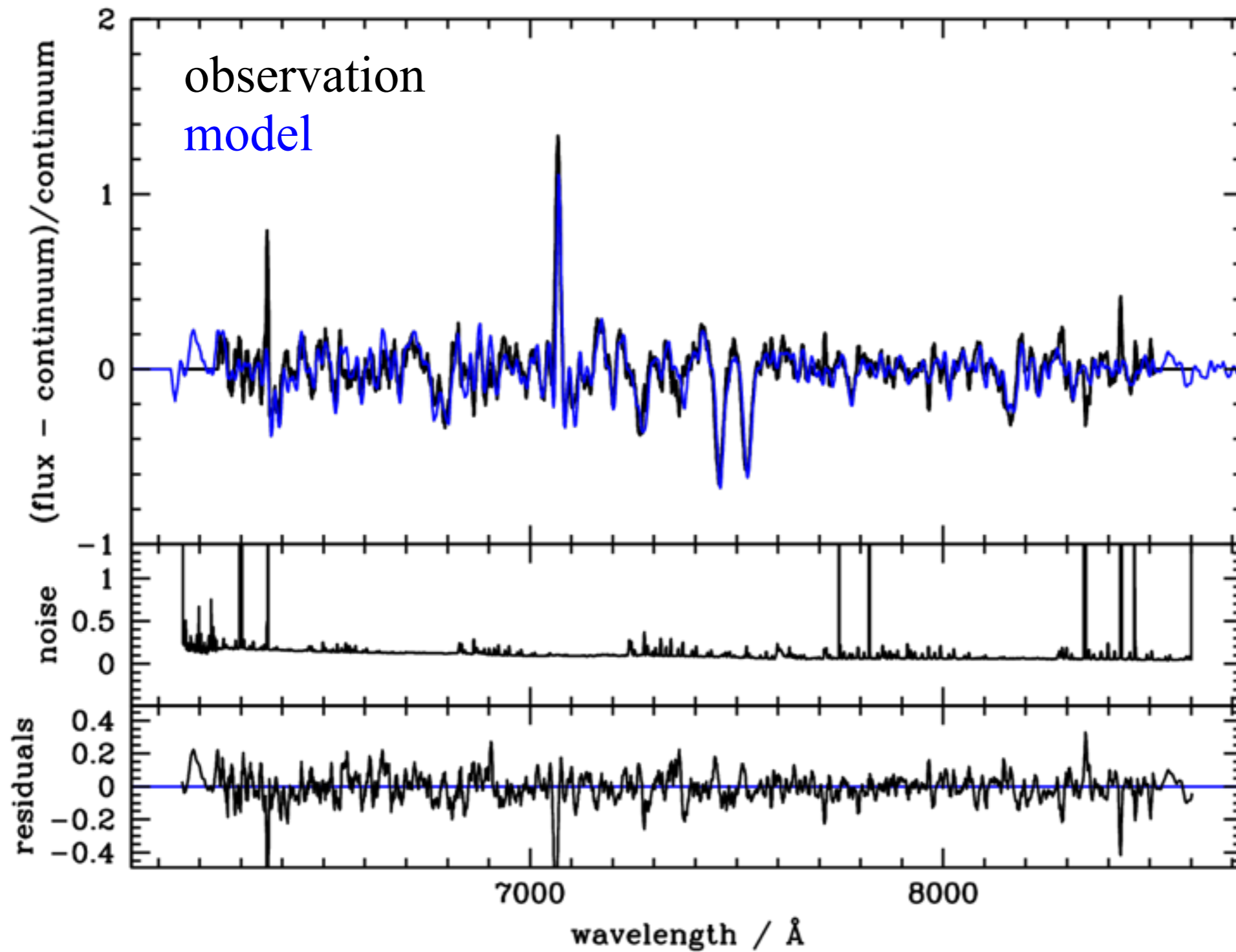


UltraVISTA  
photometry

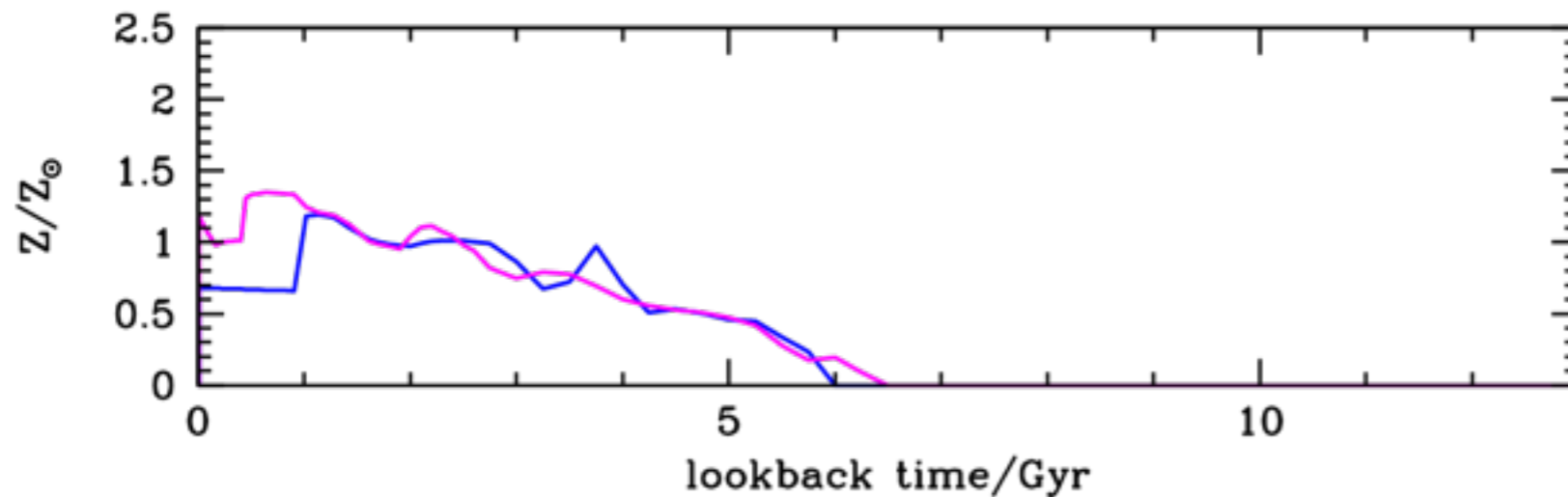
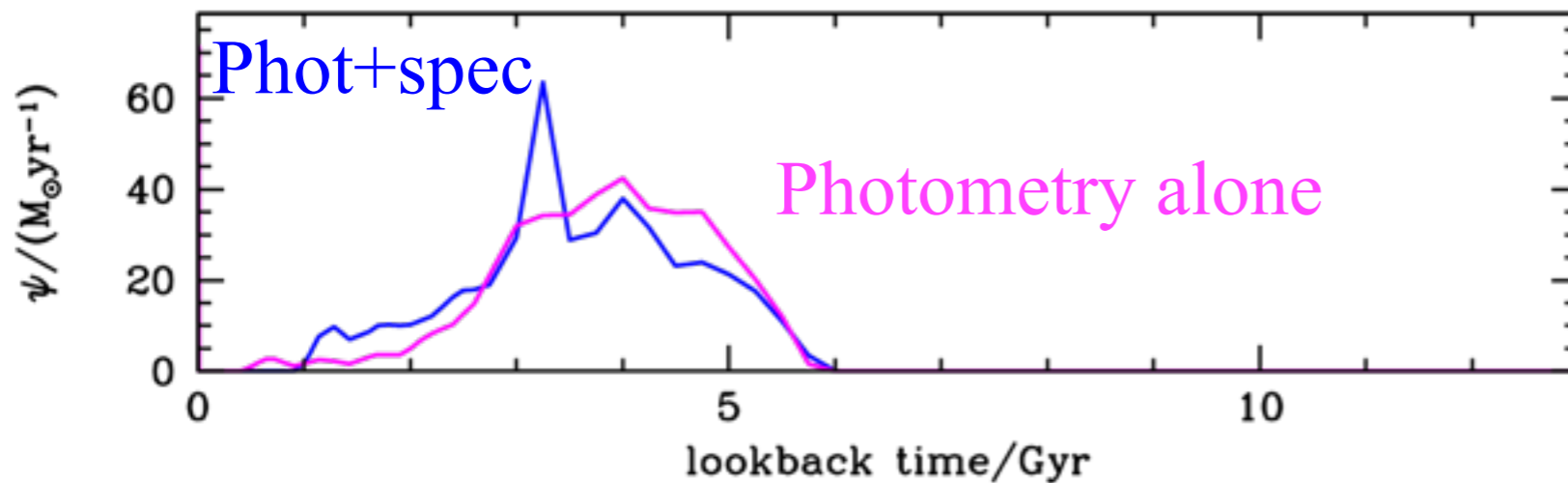
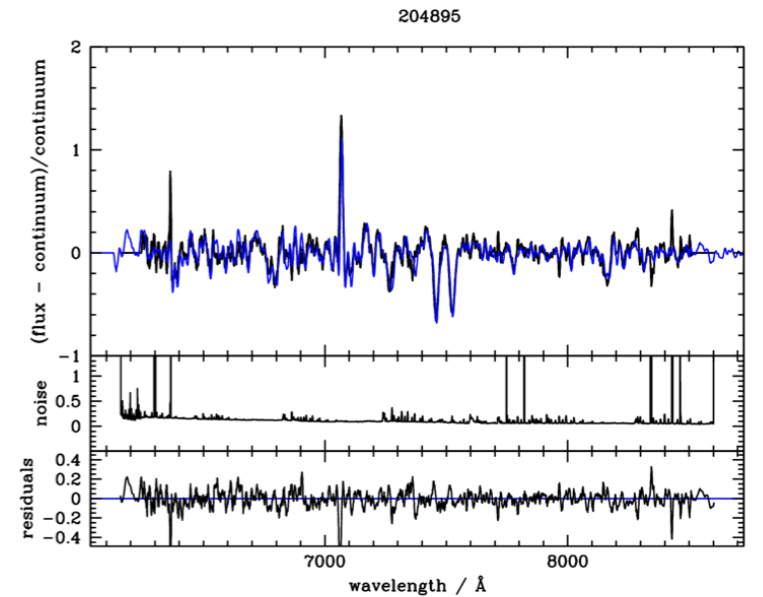
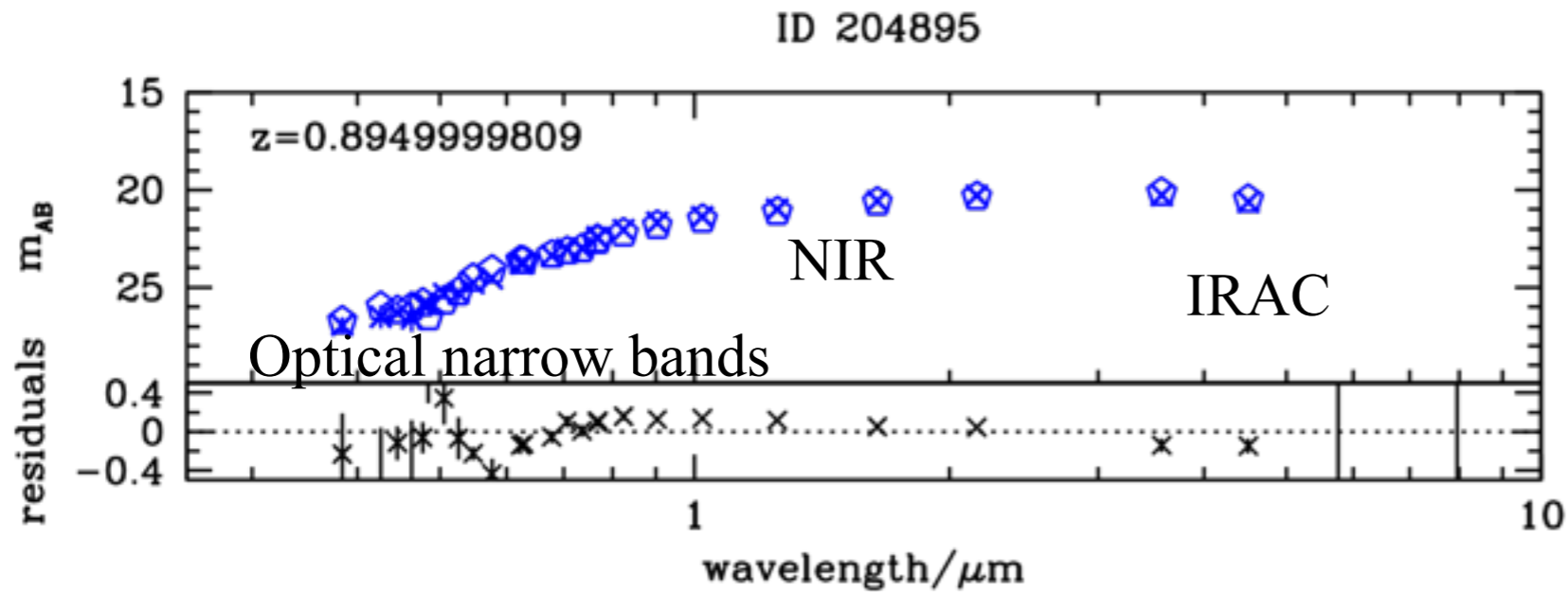
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204895

Quiescent galaxy



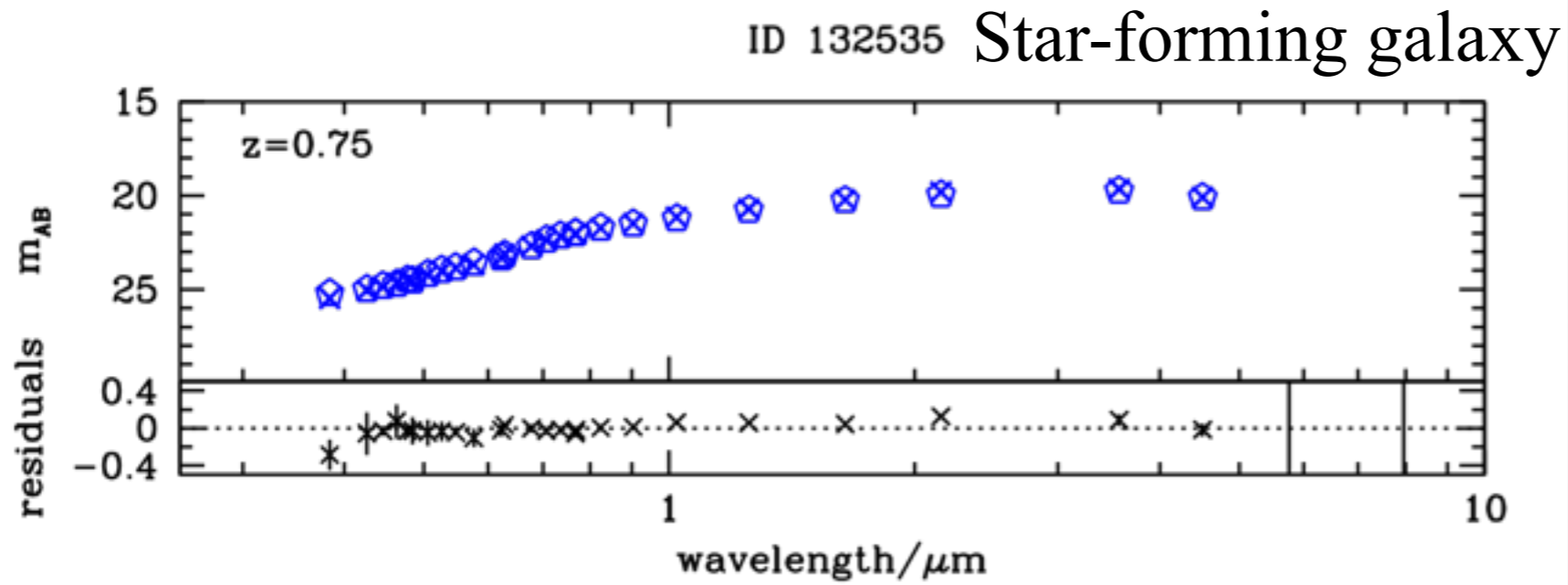
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Star formation history

Metal enrichment history

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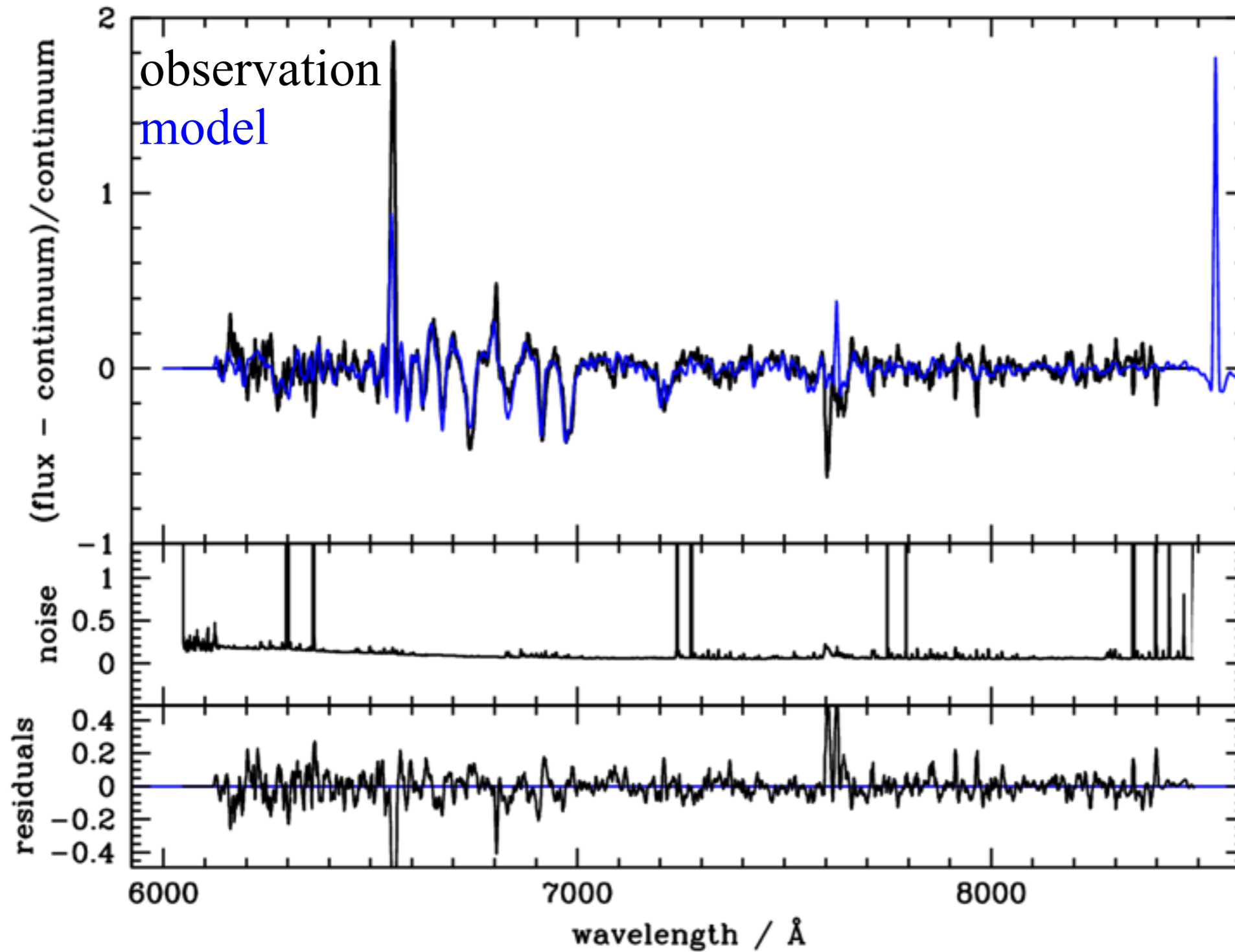


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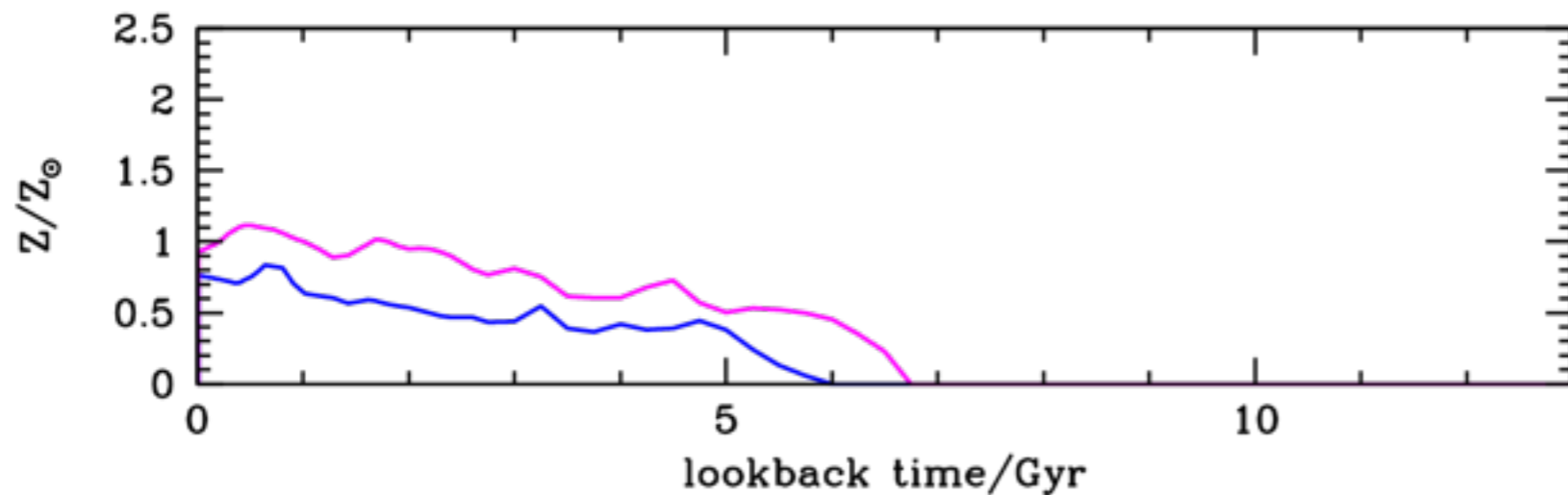
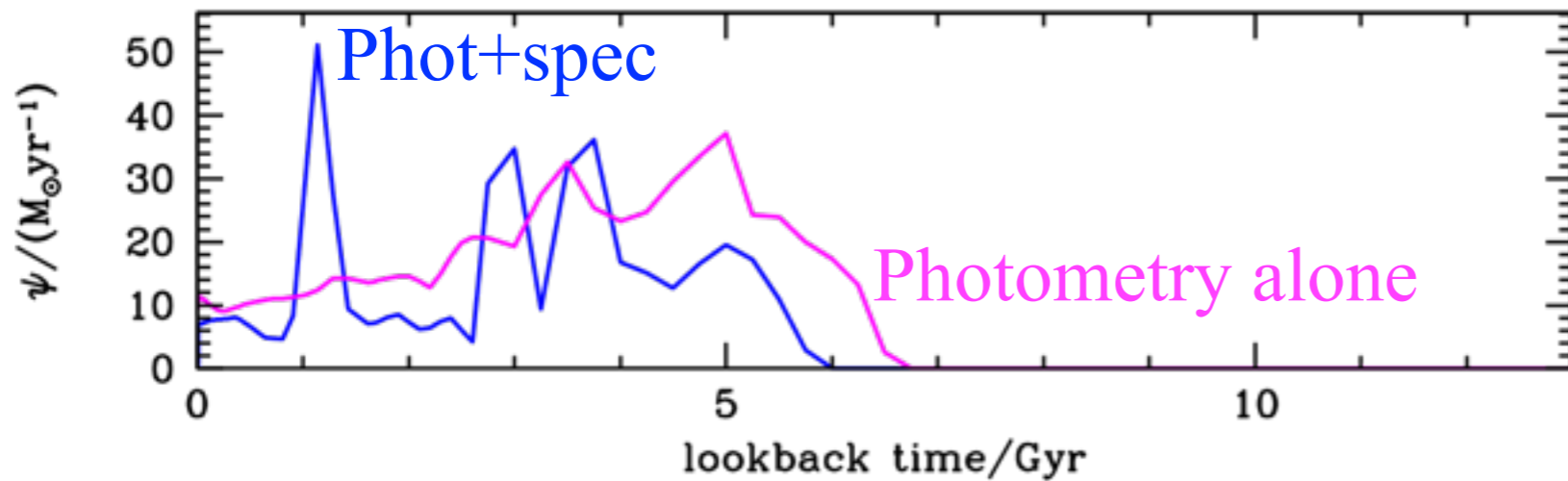
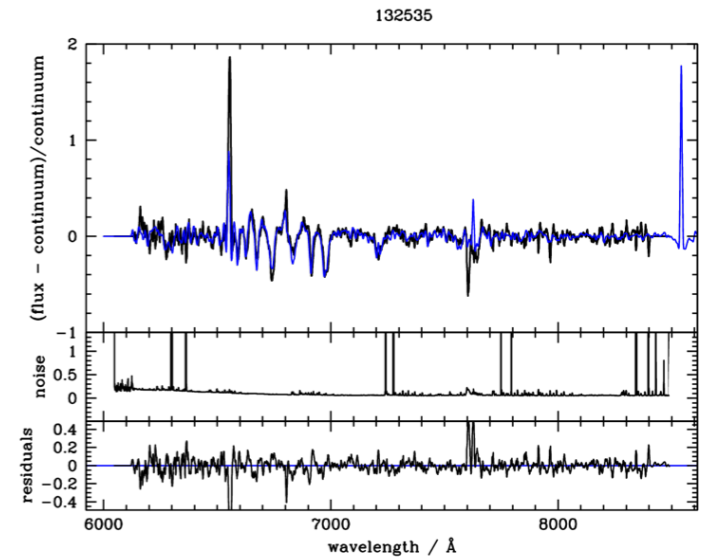
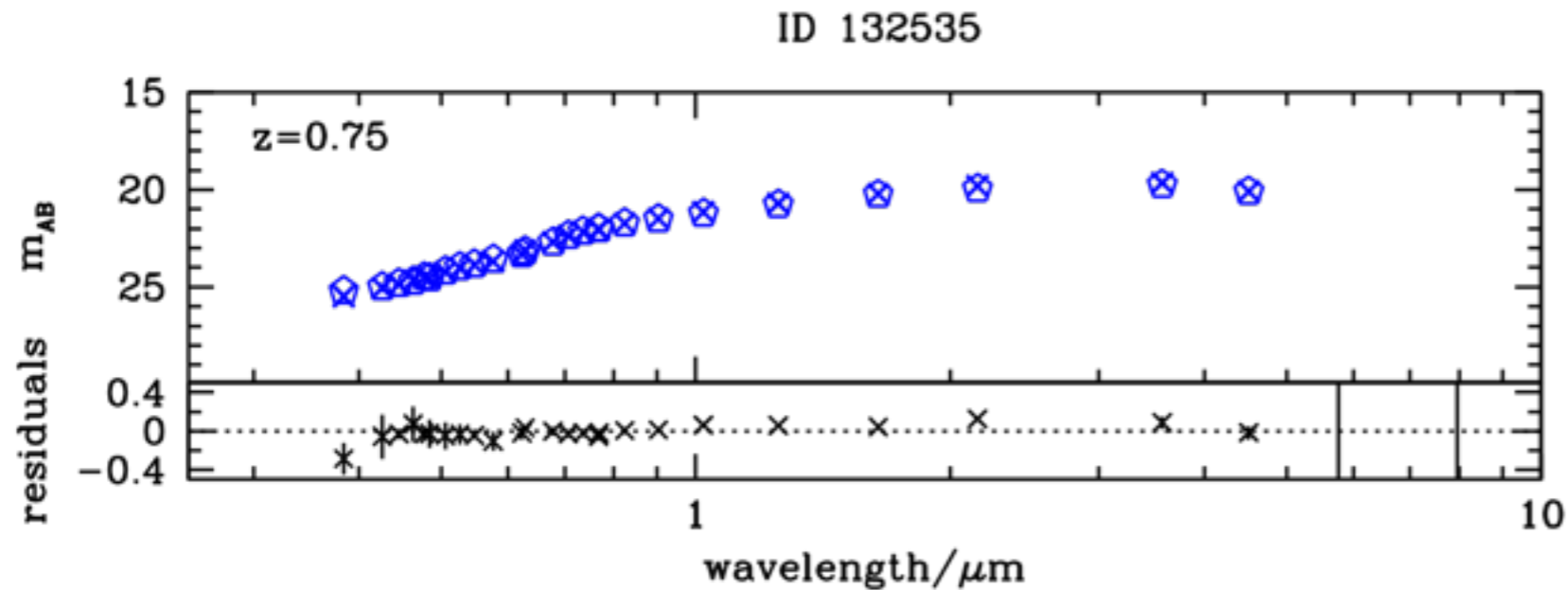
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132535

Star-forming galaxy



# Fitting procedure - photometry+spectroscopy



Star formation history

Metal enrichment history

# Conclusions

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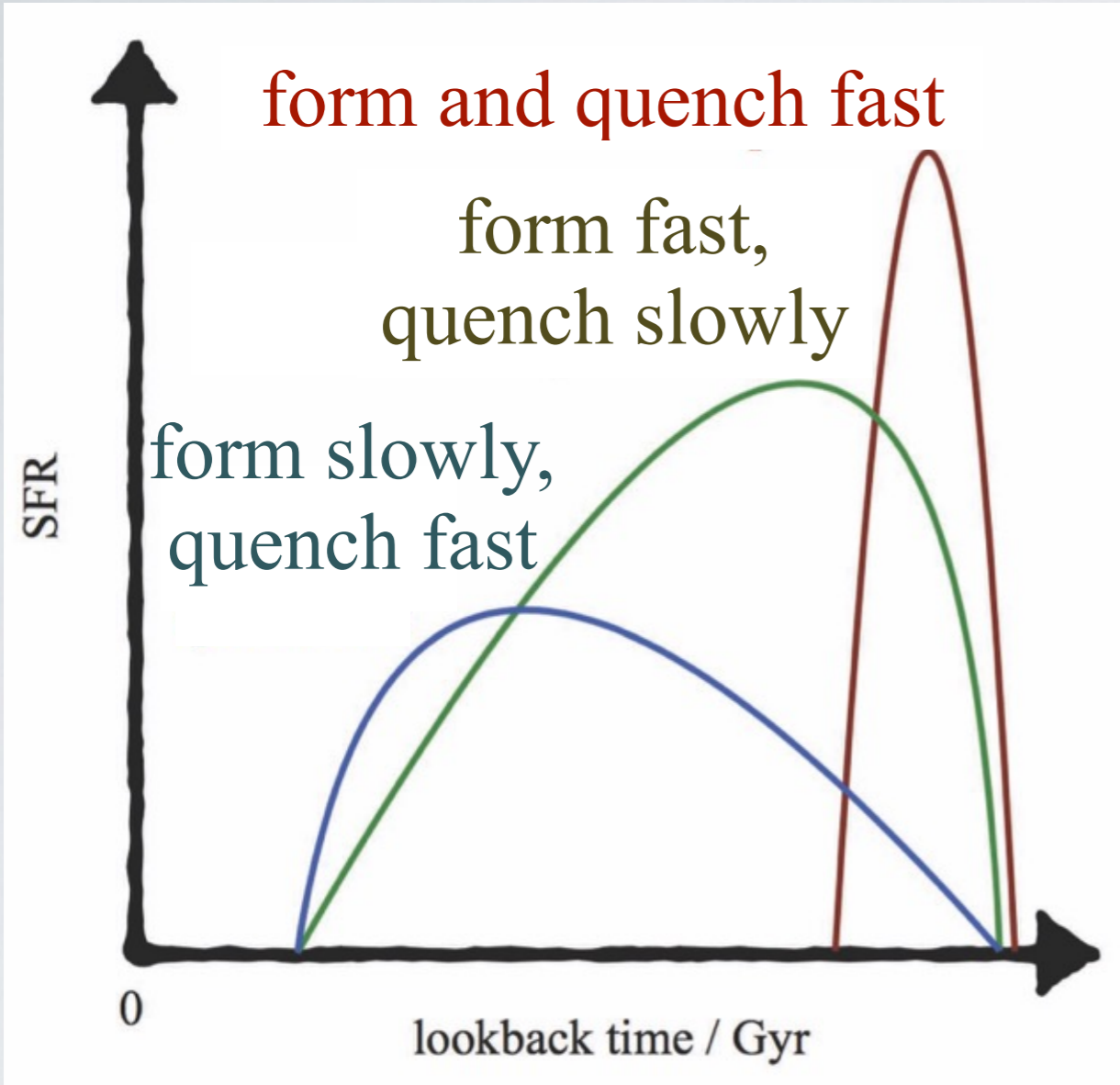
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*Julianne Dalcanton: “In Romeo and Juliet, it is not just ‘boy and girl meet, parents don’t approve, boy and girl die’. There is much more than that.”*

# Conclusions



**high mass, high z**

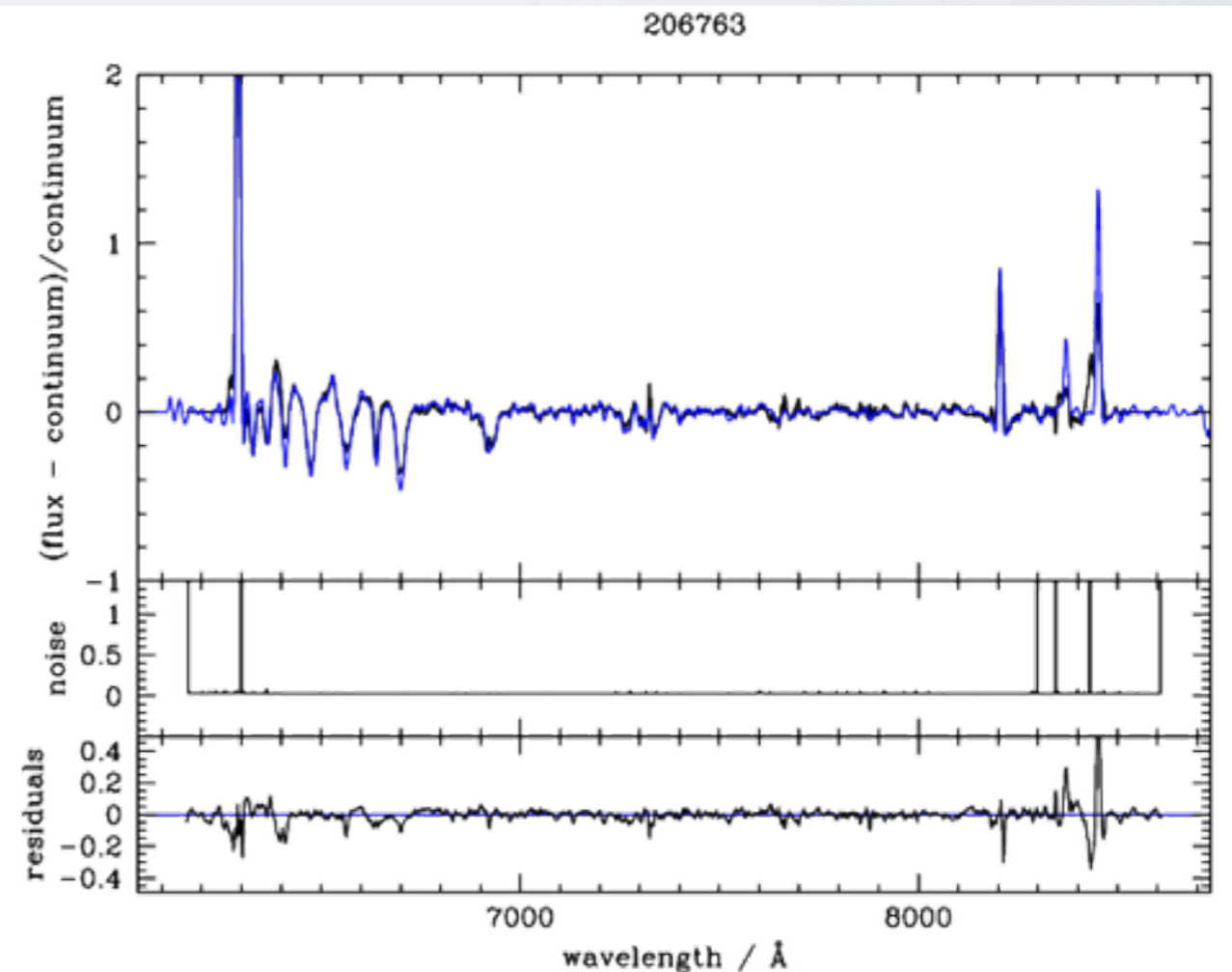
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*Thanks...Merci...Grazie*

