

How to detect GWs

Measure **proper** (i.e. physical) distance between two free-falling test masses

- Each test mass follows geodesics

Consider flat spacetime + GW (TT perturbation)

$$\frac{d^2 x^\mu}{d\tau^2} + \Gamma_{\alpha\beta}^\mu \frac{dx^\alpha}{d\tau} \frac{dx^\beta}{d\tau} = 0$$

$$\frac{d^2 x^i}{dt^2} = -(\Gamma_{tt}^i + 2\Gamma_{tj}^i v^j + \Gamma_{jk}^i v^j v^k) + v^i (\Gamma_{tt}^t + 2\Gamma_{tj}^t v^j + \Gamma_{jk}^t v^j v^k)$$

$$v^i = dx^i/dt \ll 1 \quad \longrightarrow \quad \frac{d^2 x^i}{dt^2} + \Gamma_{tt}^i = 0 \quad \Gamma_{tt}^i = \frac{1}{2} (2\partial_t h_{jt}^{\text{TT}} - \partial_j h_{tt}^{\text{TT}}) = 0$$

- **Coordinate** positions of test masses unaffected, but how about proper distance?

How to detect GWs

GW in z direction, test masses at $x=0, y=0$ and $x=L_c, y=0$

$$L = \int_0^{L_c} dx \sqrt{g_{xx}} = \int_0^{L_c} dx \sqrt{1 + h_{xx}^{\text{TT}}(t, z = 0)}$$

$$\simeq L_c \left[1 + \frac{1}{2} h_{xx}^{\text{TT}}(t, z = 0) \right]$$

$$\frac{\delta L}{L} \simeq \frac{1}{2} h_{xx}^{\text{TT}}(t, z = 0)$$

Measurable effect!

A better derivation

$$ds^2 = -dt^2 + d\mathbf{x}^2 + O\left(\frac{\mathbf{x}^2}{\mathcal{R}^2}\right)$$

- Locally flat coordinates

- Geodesics at $x^i=0$ and $x^i=L^i(t)$

- Proper distance is $\sqrt{L^i L_i}$ up to errors $\sim h L^2 / \lambda^2 \ll 1$
(NB L is Fabry-Perot cavity length for terrestrial detectors)

- Separation vector L^μ between two geodesics obeys geodesic deviation equation

$$\frac{D^2 L^\mu}{d\tau^2} = R^\mu{}_{\alpha\beta\gamma} u^\alpha u^\beta L^\gamma$$

- With $u^\mu = \delta_t^\mu$ and $L^\mu = (0, L^i)$ $\frac{d^2 L^i(t)}{dt^2} = -R_{itjt}(t, \mathbf{0}) L^j(t)$

A better derivation

$$R_{itjt} = -\frac{1}{2}\ddot{h}_{ij}^{\text{TT}} + \Phi_{,ij} + \dot{\Xi}_{(i,j)} - \frac{1}{2}\ddot{\Theta}\delta_{ij}$$

$$\Theta \sim \Phi \sim \frac{\text{mass}}{r} \qquad \Xi \sim \frac{\text{linear momentum}}{r}$$



Only TT piece (=GW) contributes far from the source!

More formally, let's show this from field equations on the board

$$\nabla^2\Theta = -8\pi\rho,$$

$$\nabla^2 S = \dot{\rho},$$

$$T_{tt} = \rho,$$

$$\nabla^2\Phi = 4\pi\left(\rho + 3P - 3\dot{S}\right)$$

$$\nabla^2\sigma = -\frac{3}{2}P + \frac{3}{2}\dot{S},$$

$$T_{ti} = S_i + \partial_i S,$$

$$\nabla^2\Xi_i = -16\pi S_i,$$

$$\nabla^2\sigma_i = 2\dot{S}_i.$$

$$T_{ij} = P\delta_{ij} + \sigma_{ij} + \partial_{(i}\sigma_{j)} + \left(\partial_i\partial_j - \frac{1}{3}\delta_{ij}\nabla^2\right)\sigma,$$

A back-of-the-envelope derivation of the quadrupole formula

Moments of mass and current distributions:

$$M_0 \equiv \int \rho d^3x = M$$

$$S_1 \equiv \int \rho v_j x_k \epsilon_{ijk} d^3x = S_i$$

$$M_1 \equiv \int \rho x_i d^3x = ML_i$$

$$M_2 \equiv \int \rho x_i x_j d^3x = ML_{ij}$$

~~$$h \sim \frac{G}{c^2} \frac{M_0}{r}$$~~

Conservation of mass

~~$$h \sim \frac{G}{c^3} \frac{d}{dt} \frac{M_1}{r} \sim \frac{G}{c^3} \frac{P}{r}$$~~

Conservation of linear momentum

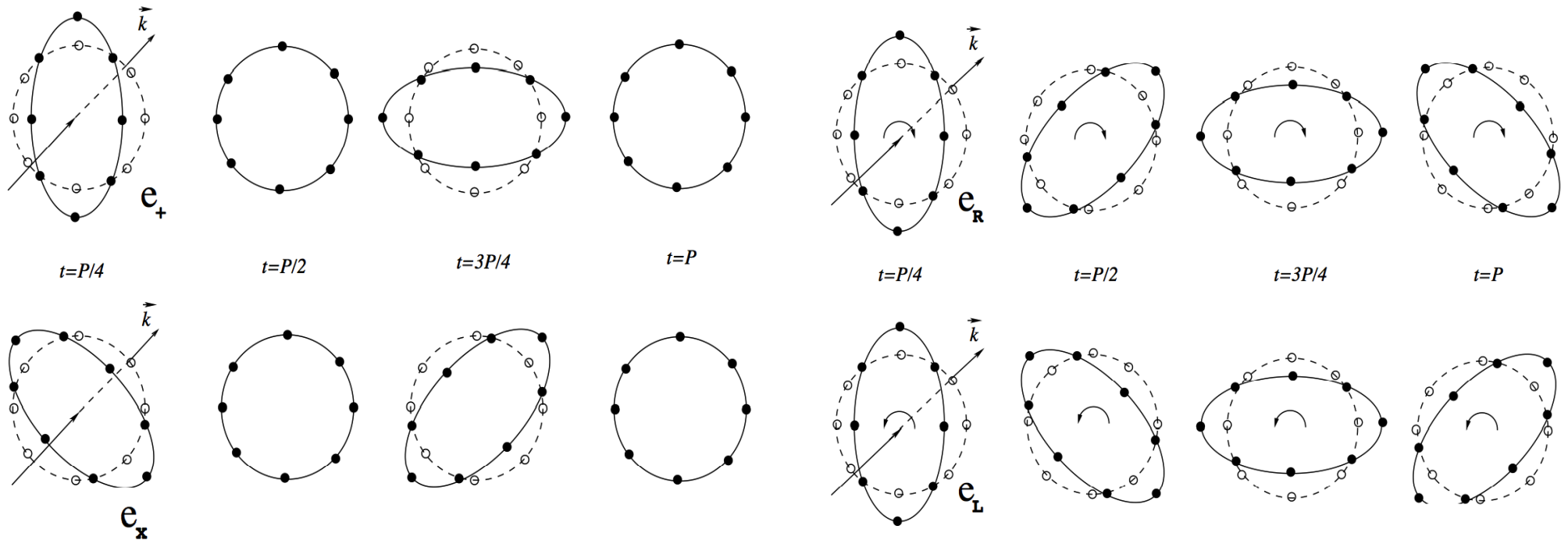
~~$$h \sim \frac{G}{c^4} \frac{d}{dt} \frac{S_1}{r}$$~~

Conservation of angular momentum

$$h \sim \frac{G}{c^4} \frac{d^2}{dt^2} \frac{M_2}{r}$$

Geometrical meaning of \mathbf{h}_+ and \mathbf{h}_x

Figures from Rezzolla's notes



$$\mathbf{e}_+ \equiv \vec{e}_x \otimes \vec{e}_x - \vec{e}_y \otimes \vec{e}_y$$

$$\mathbf{e}_x \equiv \vec{e}_x \otimes \vec{e}_x + \vec{e}_y \otimes \vec{e}_y$$

$$\mathbf{e}_R \equiv \frac{\mathbf{e}_+ + i\mathbf{e}_x}{\sqrt{2}} \quad \mathbf{e}_L \equiv \frac{\mathbf{e}_+ - i\mathbf{e}_x}{\sqrt{2}}$$

Beyond GR: more polarizations?

Similar decomposition of Riemann tensor in vacuum via Newman-Penrose scalars

$$\Psi_2(u) = -\frac{1}{6}R_{z0z0}(u),$$

$$\Psi_3(u) = -\frac{1}{2}R_{x0z0} + \frac{1}{2}i R_{y0z0},$$

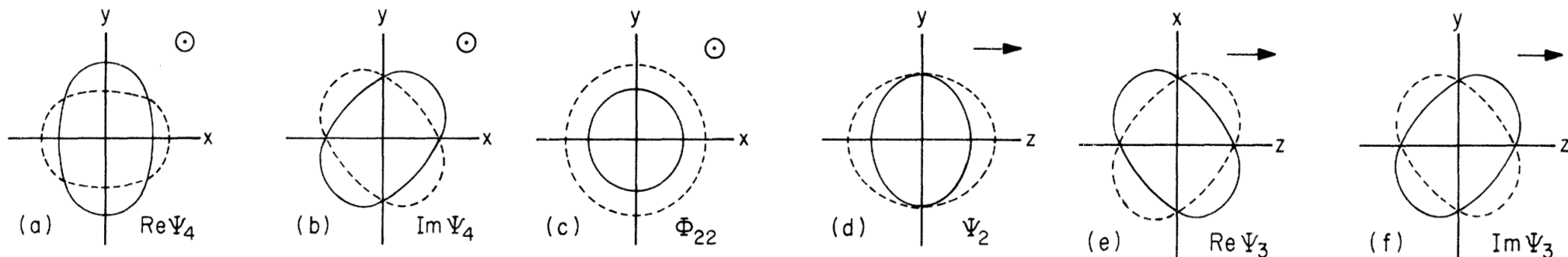
$$\Psi_4(u) = -R_{x0x0} + R_{y0y0} + 2i R_{x0y0},$$

$$\Phi_{22}(u) = -R_{x0x0} - R_{y0y0}.$$

$$\Psi_2(u) \quad (s=0), \quad \Phi_{22}(u) \quad (s=0)$$

$$\Psi_3(u) \quad (s=\pm 1), \quad \Psi_4(u) \quad (s=\pm 2)$$

Figures from Eardley, Lee and Lightman 1973



e.g. Dipolar emission if equivalence principle is violated (Brans-Dicke, scalar tensor theories, etc)

$$h \sim \frac{1}{R} \frac{d}{dt} (m_{\text{GW},1} \mathbf{x}_1 + m_{\text{GW},2} \mathbf{x}_2) \sim \frac{\eta m}{R} \mathbf{v} \left(\frac{m_{\text{GW},1}}{m_{\text{I},1}} - \frac{m_{\text{GW},2}}{m_{\text{I},2}} \right)$$

A real detector: frequencies and not distances

- Detectors are laser interferometers
- Photons accumulate phase change $\delta\phi = 2\pi\delta L/\lambda$ when proper distance between "mirrors" change

Analysis valid for $L \ll \lambda$

- More in general, we can integrate photon geodesics between mirrors; photon frequency will change due to GW and produce phase change

$$\frac{\Delta\nu}{\nu} = \frac{1}{2}(1 + \cos\theta)\Psi(t) - \cos\theta\Psi(t + \tau(1 - \cos\theta)/2) - \frac{1}{2}(1 - \cos\theta)\Psi(t + \tau),$$

$$\cos\theta \equiv \boldsymbol{\sigma} \cdot \mathbf{n}, \quad \Delta\Phi = 2\pi \int_0^t \Delta\nu(t') dt' \quad \text{Estabrook \& Wahlquist 1975}$$

$$\Psi(t) \equiv \frac{h_{ij}^{\text{TT}} \sigma^i \sigma^j}{\sin^2\theta}$$

τ = round-trip laser travel time between mirrors
 $\boldsymbol{\sigma}$ and \mathbf{n} = propagation directions of laser and GW

The detector transfer function

- Exercise: from the shift in the laser frequency, show that a monochromatic GW with frequency f propagating orthogonally to the detector causes an effective change in each detector arm, given by

$$\delta L = h_+ \frac{L \sin(\pi f \tau)}{2 \pi f \tau} = h_+ \frac{L \sin(\pi f / \Lambda)}{2 \pi f / \Lambda}$$

- This **transfer function** $T(f)$ explains why the sensitivity of GW interferometers worsens linearly with f at high frequencies

The detector response

At least when $L \ll \lambda$ (i.e. $T(f) \sim 1$) an interferometer measures

$$\begin{aligned} h(t) &= \frac{1}{2} (h_{ij} u^i u^j - h_{ij} v^i v^j) \\ &= D^{ij} h_{ij}(t) = F_+ h_+(t) + F_\times h_\times(t) \end{aligned}$$

$$D^{ij} = \frac{1}{2} (u^i u^j - v^i v^j) \quad \text{Detector tensor}$$

$$F_+, F_\times \quad \text{Beam pattern functions}$$

Pattern functions

Exercise: derive pattern functions for detectors at 90 and 60 degrees, and plot them

$$F_{+}^{(90^{\circ})} = \frac{1}{2} (1 + \cos^2 \theta) \cos 2\phi \cos 2\psi - \cos \theta \sin 2\phi \sin 2\psi,$$

$$F_{\times}^{(90^{\circ})} = \frac{1}{2} (1 + \cos^2 \theta) \cos 2\phi \sin 2\psi + \cos \theta \sin 2\phi \cos 2\psi.$$

$$F_{+}^{(60^{\circ})} = \frac{\sqrt{3}}{2} \left[\frac{1}{2} (1 + \cos^2 \theta) \cos 2\phi \cos 2\psi - \cos \theta \sin 2\phi \sin 2\psi \right],$$

$$F_{\times}^{(60^{\circ})} = \frac{\sqrt{3}}{2} \left[\frac{1}{2} (1 + \cos^2 \theta) \cos 2\phi \sin 2\psi + \cos \theta \sin 2\phi \cos 2\psi \right].$$