

Tidal Stream Detection in HSC-SSP with Deep Learning

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With the **valuable contribution** of:

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- I. Damjanov (University St. Mery)
- H. Souchereau (University St. Mery)
- O. Karishma (University St. Mery)
- E. Kado-Fong (Princeton University)
- K. Johnston (Columbia University)
- B. Robertson, R. Hausen (UCSC)

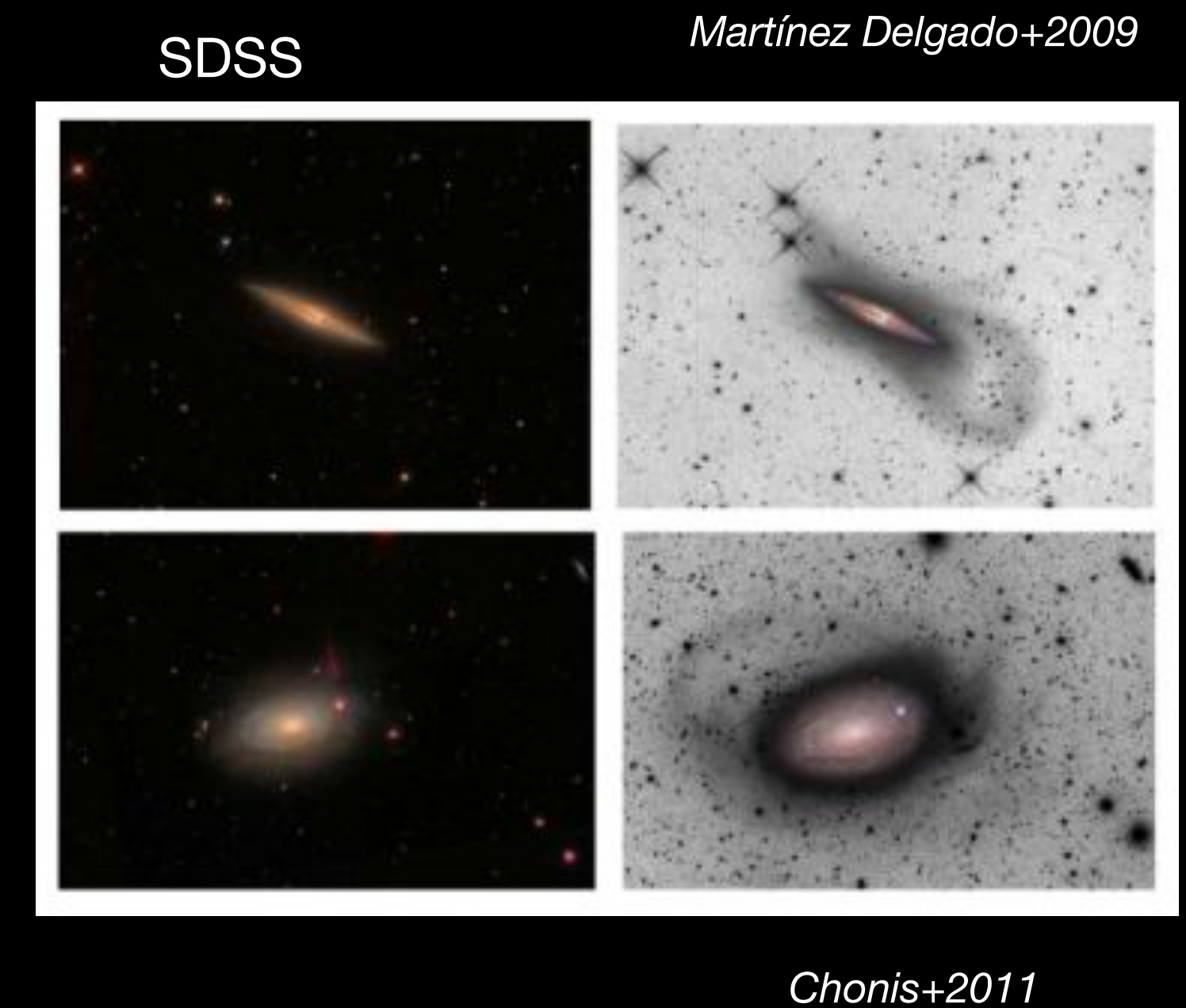


Debating the potential of machine learning in astronomical surveys

21/10/2021

Context: The importance of Tidal Streams

- The frequency and characteristics of low surface brightness features can be used to disentangle the different formation channels (in situ star formation versus accreted stars).
- Classical ‘cat/dog’ problem: optimal for CNN but...
- Tidal streams are **difficult to detect**: rare (only a few dynamical periods) and faint (imprints in the outskirts of galaxies)
- Need deep imaging and large areas! —> Small training samples available



Supervised Learning: training sample

- Pioneer attempt by Walmsley+19

- CFHTLS-Wide Survey

- 1316 galaxies, 305 labelled as tidal streams

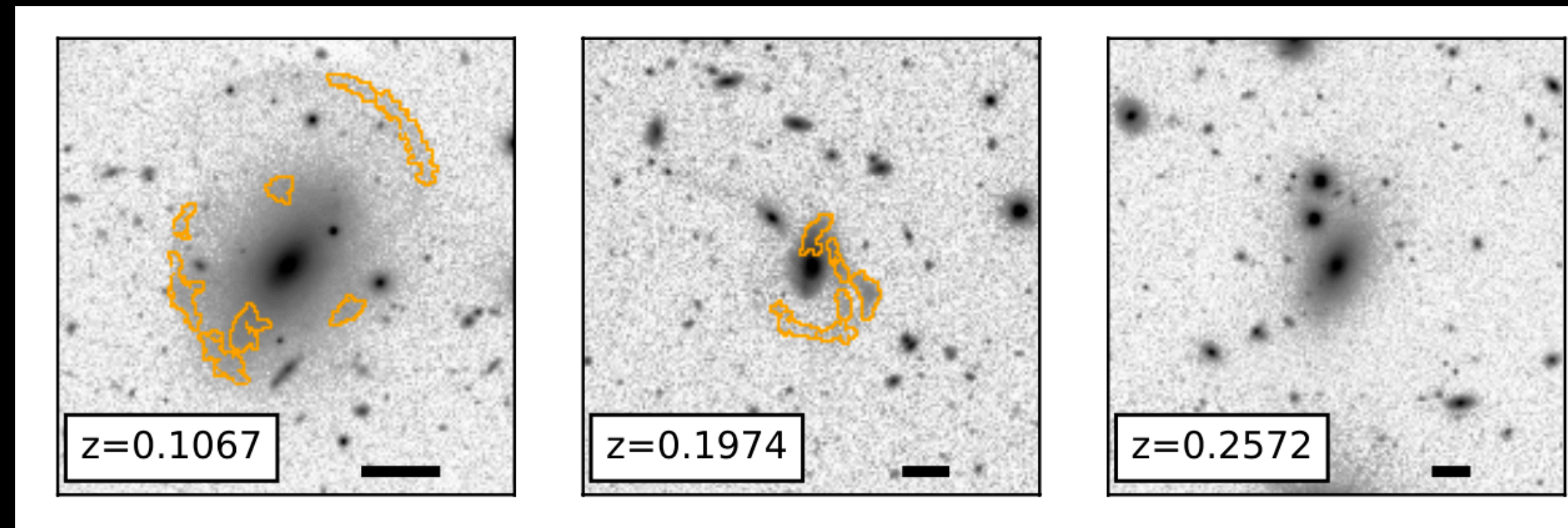
- HSC-SSP data (Hyper Suprime-Cam Subaru Strategic Program)

- $m_i \sim 26$ mag (5σ point source), 1400 deg²

- 21200 galaxies at $0.05 < z < 0.45$; $m_r < 17.7$

- Filtering algorithms for high spatial frequency features + Visual inspection

- 1200 show tidal features

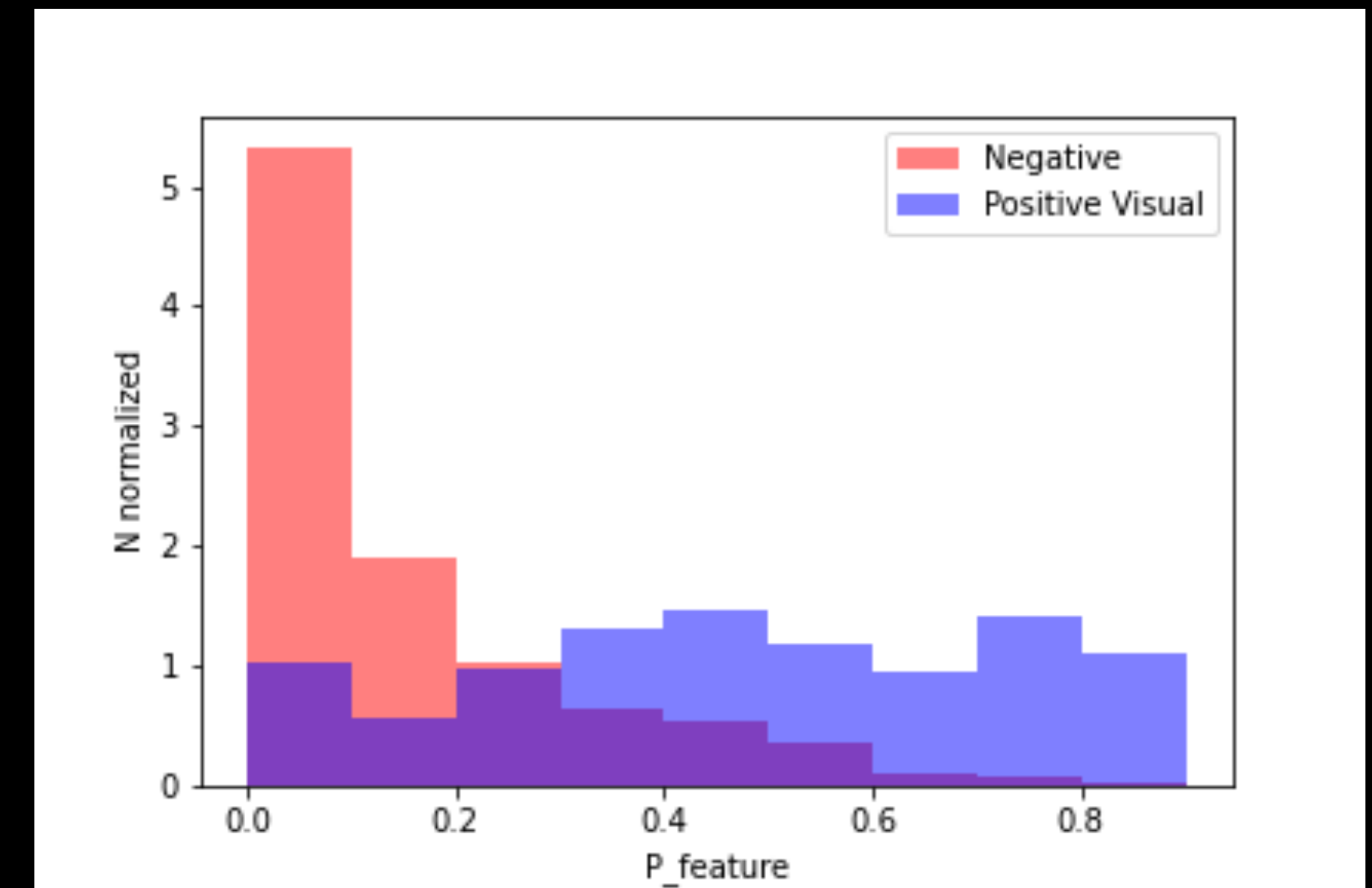
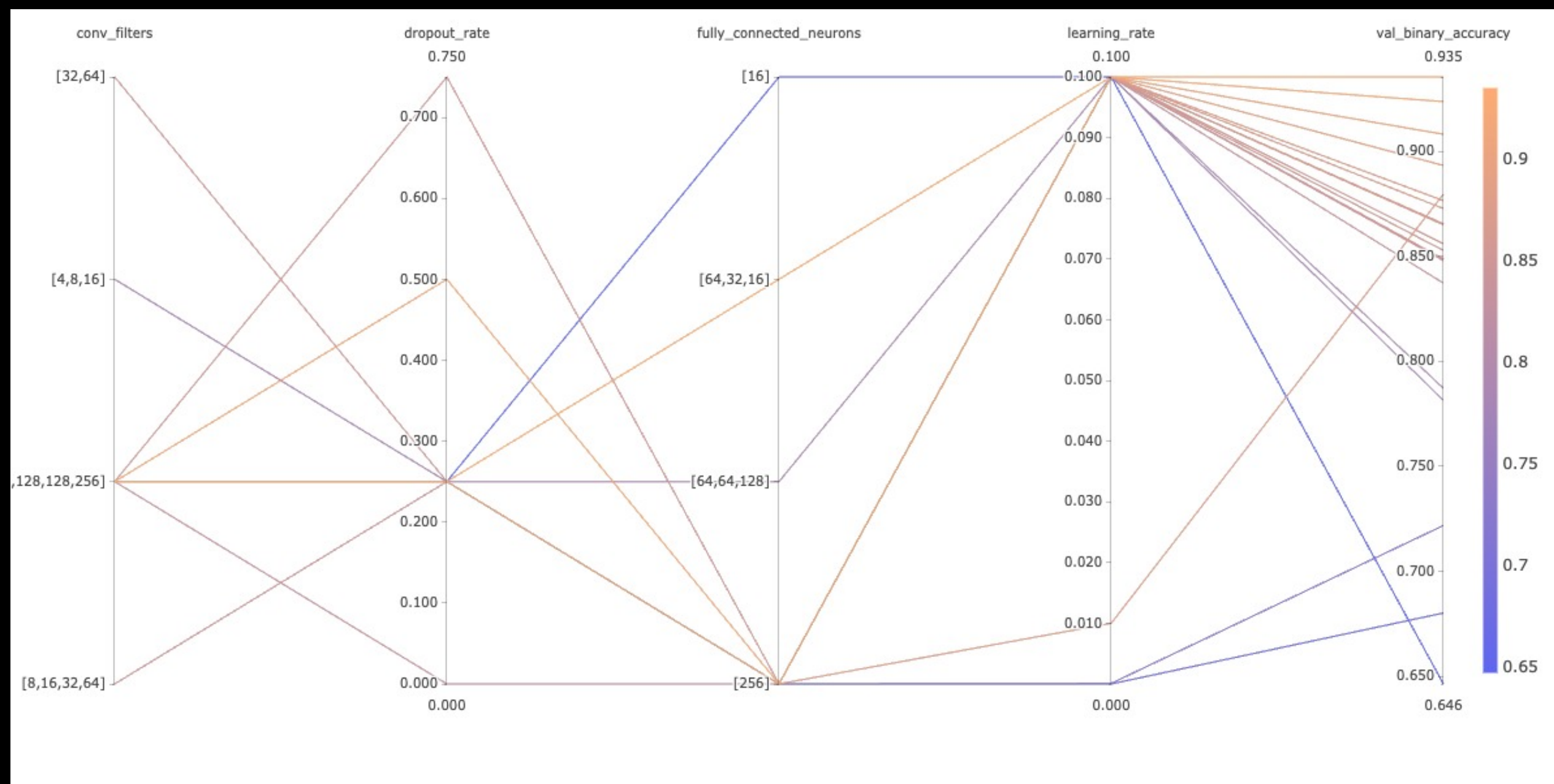


Results (preliminary)

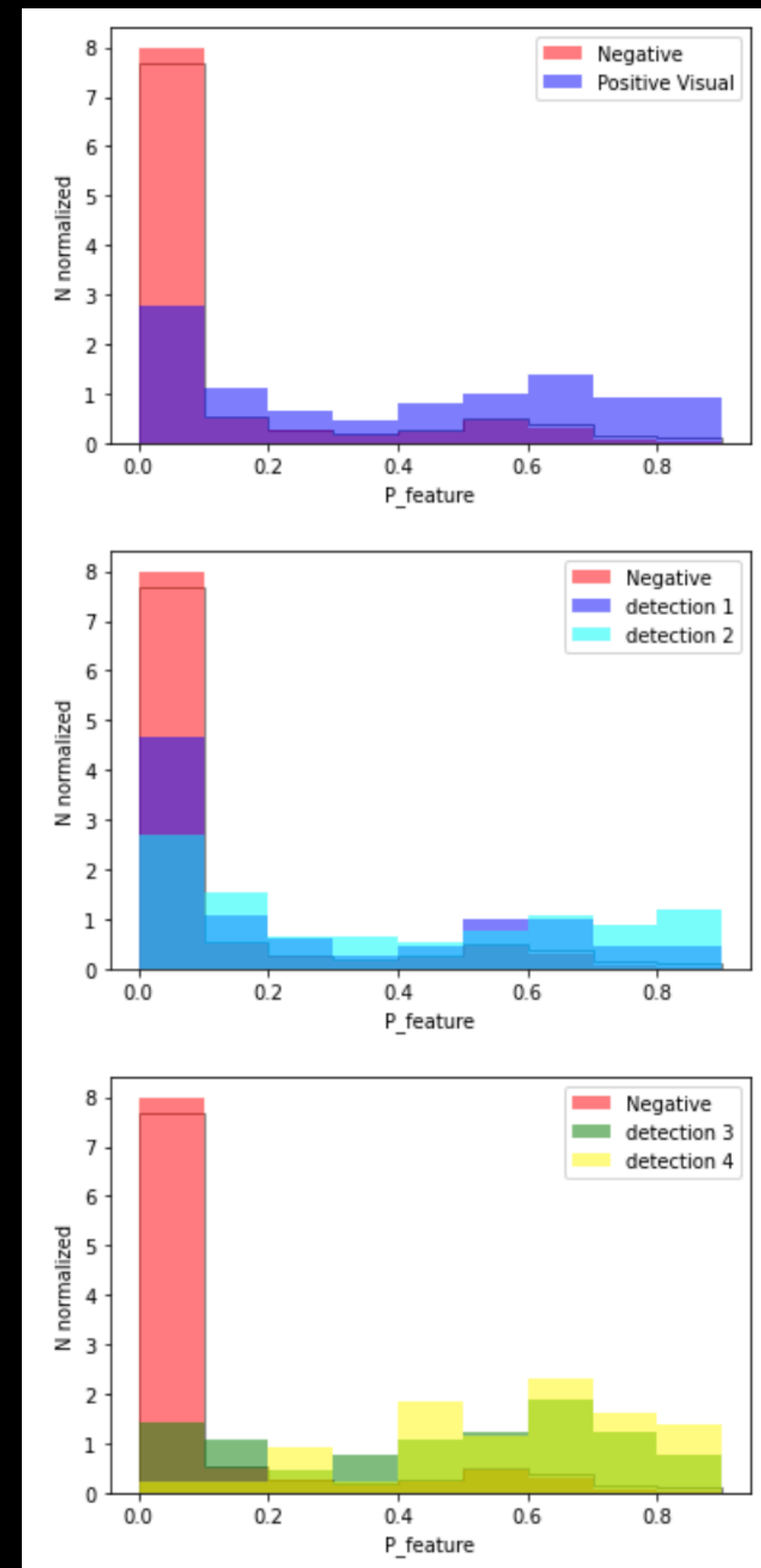
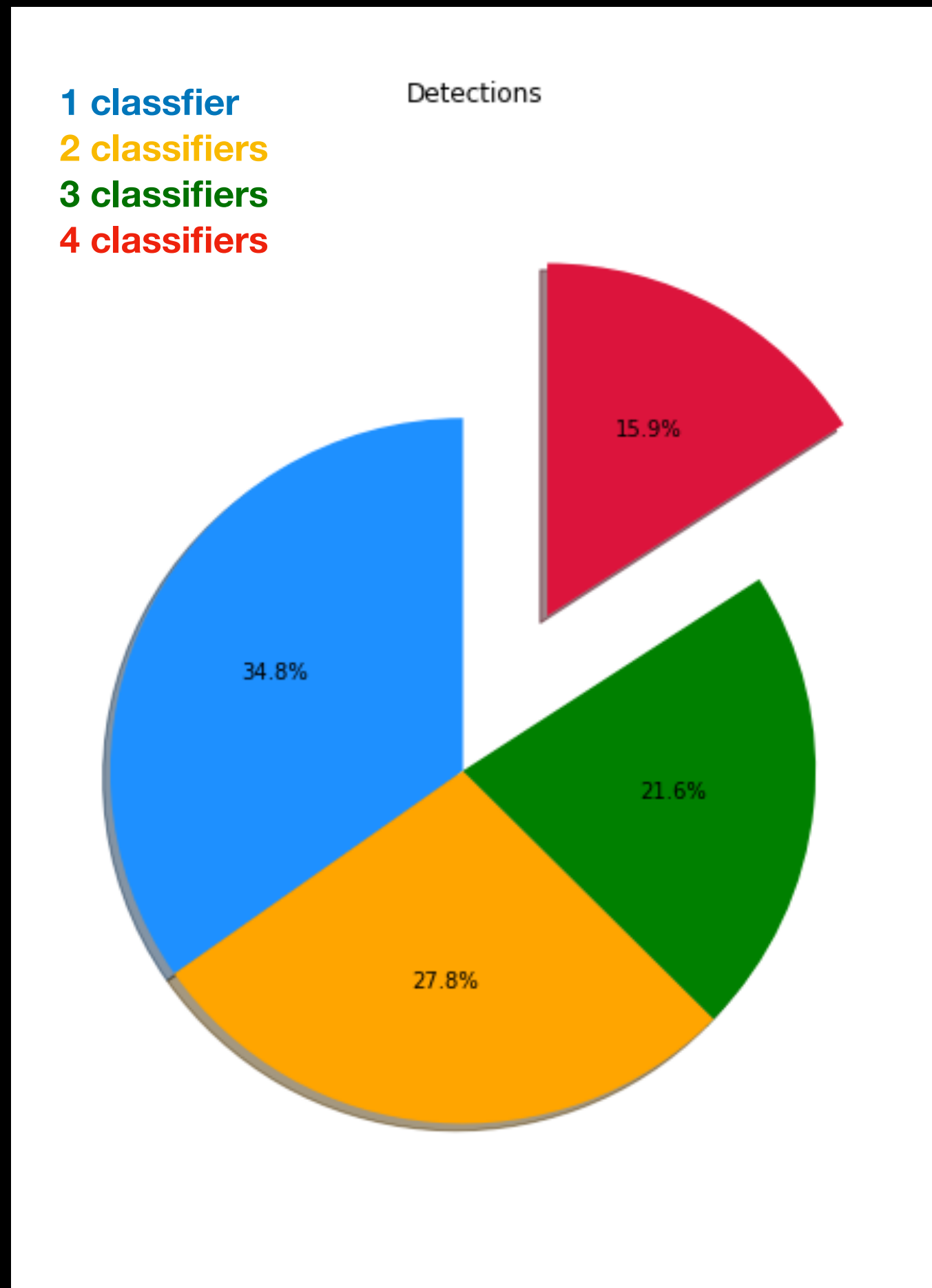
- Hyperparameter optimization
- Image pre-processing

Completeness
Contamination

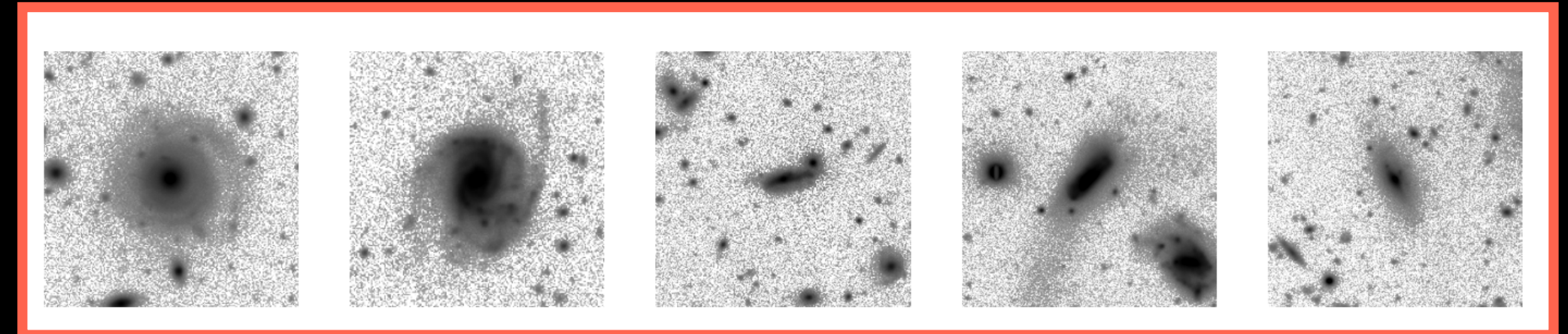
	AUC	Acc.	Recall	Prec.
Unbalanced	0.90	0.94	0.49	0.56
Balanced	0.90	0.74	0.49	0.93



Tidal stream visual identification



- Significant disagreement between classifiers!
- Galaxies with more votes show larger Prob.
- False positives show some features.



Conclusions

Automated detection of tidal features is a **challenging** task due to:

- Low number of real tidal stream detections (small training samples)
- Subjectivity of the tidal stream identification (wrong labels)
- Subtleness of tidal features (low contrast and noisy images)

Future Work

- Neural Network architecture
- Transfer Learning (e.g., SDSS models)
- Combine observational datasets
- Simulations

*Open to collaboration!
Contact me*

