

Photometric Redshifts for Cosmology: Improving Accuracy Evan Jones¹, Tuan Do¹, Bernadette Boscoe^{1,3}, Yujie Wan¹, Zoey Nguyen¹, Jack Singal² and Uncertainty Estimates Using Bayesian Neural Networks

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Abstract

We present results exploring the role that probabilistic deep learning models can play in cosmology from large scale astronomical surveys through estimating the distances to galaxies (redshifts) from photometry. Due to the massive scale of data coming from these new and upcoming sky surveys, machine learning techniques using galaxy photometry are increasingly adopted to predict galactic redshifts which are important for inferring cosmological parameters such as the nature of Dark Energy. Associated uncertainty estimates are also critical measurements, however, common machine learning methods typically provide only point estimates and lack uncertainty information as outputs. We turn to Bayesian Neural Networks (BNNs) as a promising way to provide accurate predictions of redshift values. We have compiled a new galaxy training dataset from the Hyper Suprime-Cam Survey, designed to mimic large surveys, but over a smaller portion of the sky. We evaluate the performance and accuracy of photometric redshift (photo-z) predictions from photometry using machine learning, astronomical and probabilistic metrics. We find that while the Bayesian Neural Network did not perform as well as non-Bayesian Neural Networks if evaluated solely by point estimate photo-z values, BNNs can provide uncertainty estimates that are necessary for cosmology.

Data: Galaxy Observations

Comparison of a Neural Network to Bayesian Neural Network

0 Fig 1. (left): Typical galaxy image in the i-band. We We find that the non-Bayesian NN produces better point estimate photo-z predictions compared to the BNN, however the BNN use g,r,i,z,y photometry from the Hyper-Suprime has the advantage of producing uncertainties for photo-z predictions. We note that the photo-z uncertainties are overestimated Cam (HSC) Public Data Release 2 (PDR2)¹, which is at low redshifts (0 < z < 2.5) and underestimated at higher redshifts. We believe the BNN may require more information than is y pixel designed to reach similar depths as LSST² but over present in the g,r,i,z,y inputs in order to improve performance; in a future work we will investigate galaxy images as input in a 60 a smaller portion of the sky. bayesian convolutional neural network. Bavesian non-Bayesian bias bias $z_{\rm phot}-z_{\rm spec}$ 20000 0.8 bias MAD MAD 120 0.6 $1 + z_{\rm spec}$ 0.6 O O 60 120 0.4 0.4 x pixel metrics metrics 0.2 0.2 Fig. 2 (right): Number of galaxies in 0.0 V(Z) 0.0 $-\mathbf{z}_{\mathbf{spec}}$ $\mathbf{z}_{\mathbf{phot}}$ our sample as a function of redshift. In -0.2-0.2 $\bar{1} + z_{\rm spec}$ total. our data consists of 286,401 -0.4 -0.4 galaxies with broad-band g,r,i,z,y -0.6 -0.6 -0.8 -0.8 photometry and known spectroscopic Ω 2 redshifts³. We use 80% for training $-\mathbf{z}_{\text{spec}}$ ò $\mathbf{z}_{\mathbf{phot}}$ 1 2 Zspec Zspec $|O_{\rm h}|$ Z_{spec} and 20% for testing. $1 + z_{spec}$ 03 0 Network architectures overestimated o O_{P} 0.8 4 Dense lavers 1 Dense layer Z_{spec}) Inderestimated & $MAD = Median(|\Delta z - Median(\Delta z_i))$ 0 3 coverage outliers 0.4 concatenation NN: g,r,i,z,y photo-z estimate laver ы $\sum_{i=1}^{n} \overline{(\overline{z}_{pdf,i})}$ $- \mathbf{z_{spec,i}}) < \mathbf{z}_{\sigma,i}$ 0.2 coverage =0.2 n_{gals} 0.0 1 IndependentNormal laver 0.0 4 DenseVariational lavers Zspec Zspec Zspec concatenation BNN: References g,r,i,z,y photo-z PDF laver ¹ Aihara, H., AlSayyad, Y., Ando, M., et al. 2019, Pasj ³ https://hsc-release.mtk.nao.ac.jp/doc/index.php/dr1 specz/

²LSST Collaboration 2021, arXiv:1809.01669 [astro-ph].