Optimising the shape of photometric redshift distributions with clustering cross-correlations

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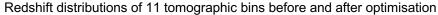


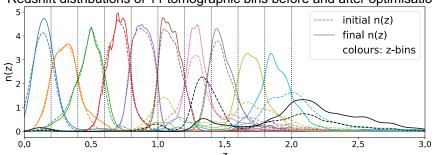
Goals:

Optimising the assignment of galaxies to tomographic redshift bins and reducing the rate of catastrophic outliers in the redshift distribution.

Methods: Self-organising map (SOM), trained 175 on the observed colours of galaxies Clustering cross-correlation 2.0 measurements between a photometric galaxy sample and a reference sample with known redshifts Reassignment of galaxies to tomographic bins via simulated 100 125 150 annealing SOM dimension 1 Optimisation SOM node selection Colour Final tomographic measurements Photometric Initial tomographic edshift estimates Cross-correlation Redshift bin measurement reassignment

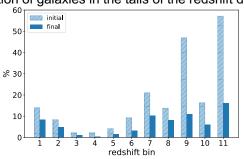
Results:





Comparison of the initial redshift distribution of 11 tomographic bins, obtained using the photometric redshift estimate of individual galaxies (dashed lines), and the redshift distribution after optimisation via simulated annealing (solid lines). Dotted lines indicate the redshift bin edges.

Fraction of galaxies in the tails of the redshift distribution



Comparison of the percentage of the redshift distribution that is located within the tails of the distribution before and after optimisation.